

# Chiral Composite Asymmetric Dark Matter

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Masahiro Ibe, Shin Kobayashi and Keiichi Watanabe  
(Tokyo U, ICRR)

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# Contents

0. The Introduction of Asymmetric Dark Matter (ADM)
1. A model of (Non-chiral) Composite ADM
2. A Model of Chiral Composite ADM
3. Conclusion

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# 0. The Introduction of Asymmetric Dark Matter (ADM)

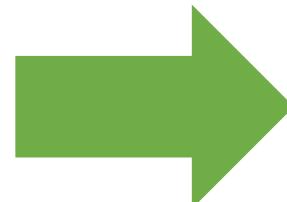
- Baryon asymmetry exists in the Universe.

$$\eta \equiv (n_b - n_{\bar{b}})/n_\gamma \sim 10^{-10}$$

- The ratio of  $\Omega_{\text{DM}} : \Omega_B$  is  $\mathcal{O}(1)$ .

$$\Omega_{\text{DM}}/\Omega_B \simeq 5 \sim \mathcal{O}(1)$$

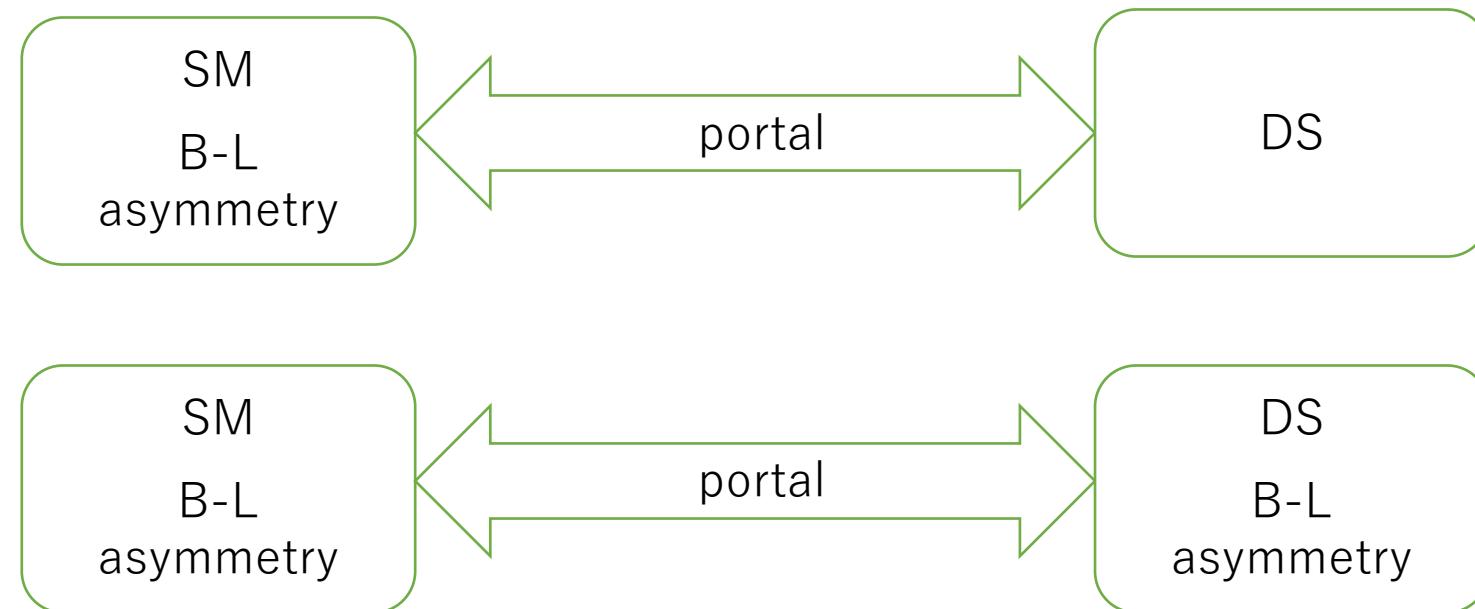
Do these matters have  
a common origin?



ADM

# 0. The Introduction of Asymmetric Dark Matter (ADM)

$$\mathcal{L}_{B-L \text{ portal}} = \mathcal{O}_{DS} \mathcal{O}_{SM} / M^n$$



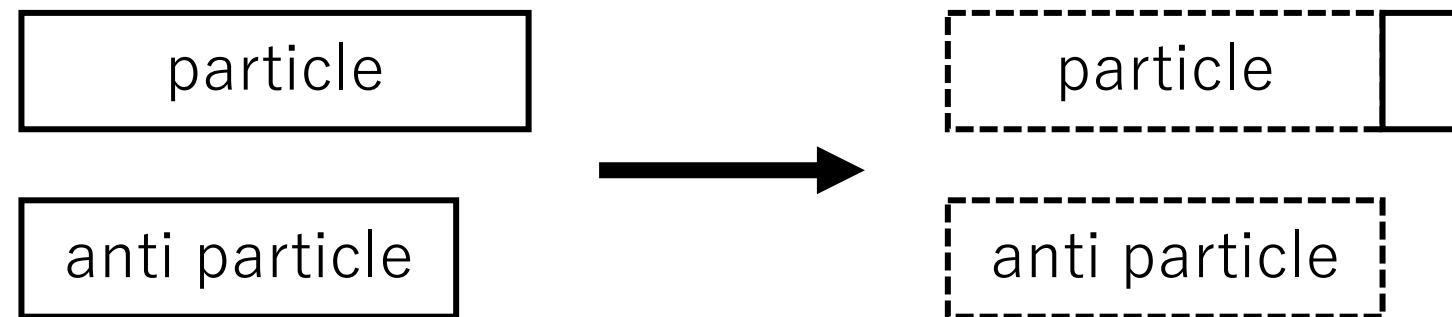
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# 1. A model of (Non-chiral) Composite ADM [1]

Why was composite model considered ?

- Symmetric component almost vanish like baryons



- Confinement scale determines DM mass ( $\mathcal{O}(1)$  GeV)

$$\Lambda_D \sim 10 \times \Lambda_{QCD}$$

# 1. A model of (Non-chiral) Composite ADM [1]

	gauge	global	
	$SU(3)_D$	$U(1)_D$	$U(1)_{B-L}$
$U$	<b>3</b>	$2/3$	$1/3$
$D$	<b>3</b>	$-1/3$	$1/3$
$\bar{U}$	<b><math>\bar{3}</math></b>	$-2/3$	$-1/3$
$\bar{D}$	<b><math>\bar{3}</math></b>	$1/3$	$-1/3$

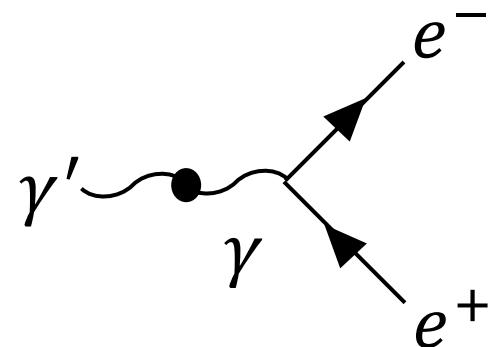
- Dark baryons and dark mesons arise due to the confinement
  - dark baryons  $\rightarrow$  dark matter  
e.g.  $p' \propto UUD, n' \propto UDD$
  - dark mesons  $\rightarrow$  over density  
e.g.  $\pi'^0 \propto U\bar{U} - D\bar{D}, \pi'^+ \propto U\bar{D}$



Necessity of annihilating or decaying dark mesons into SM particles !

# 1. A model of (Non-chiral) Composite ADM [1]

dark meson  $\rightarrow$  SM, How ?  
 $\rightarrow$  dark photon (kinetic mixing)



massless  $\rightarrow N_{\text{eff}}$  X →

dark photon must be  
massive ! ( $\mathcal{O}(10 - 100)$  MeV)

$$2 \times m_e < m_{\gamma'} < m_{\pi'} < m_{\text{DM}}$$

# 1. A model of (Non-chiral) Composite ADM [1]

dark photon mass ( $\mathcal{O}(10 - 100)$  MeV)  $\leftarrow$  dark Higgs

Problems arise here !

( Dark Higgs VEV tuning is needed in addition to  $\Lambda_D$  !

Nothing is said about dark Higgs in [1] !  
(e.g. Thermal history, 4-point coupling with SM Higgs)

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## 2. A Model of Chiral Composite ADM

	gauge	global	
	$SU(3)_D$	$U(1)_D$	$U(1)_{B-L}$
$U$	3	1	1/3
$D$	3	-1	1/3
$S$	3	0	1/3
$\bar{U}$	$\bar{3}$	$-a$	-1/3
$\bar{D}$	$\bar{3}$	$a$	-1/3
$\bar{S}$	$\bar{3}$	0	-1/3

- We assume  $0 < a < 1$
- $SU(3)_L \times SU(3)_R$  approximate flavor symmetry exists
- $U$  and  $D$  can not have the mass



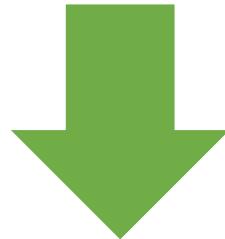
3 flavor  $\leftarrow \mathcal{L}_{B-L \text{ portal}} = \mathcal{O}_{DS} LH/M^n$

## 2. A Model of Chiral Composite ADM

	SU(3) <sub>D</sub>	U(1) <sub>D</sub>	U(1) <sub>B-L</sub>
$U$	<b>3</b>	1	1/3
$D$	<b>3</b>	-1	1/3
$S$	<b>3</b>	0	1/3
$\bar{U}$	<b><math>\bar{3}</math></b>	$-a$	-1/3
$\bar{D}$	<b><math>\bar{3}</math></b>	$a$	-1/3
$\bar{S}$	<b><math>\bar{3}</math></b>	0	-1/3

Below  $\Lambda_D$ ,

$$\begin{aligned} \langle U\bar{U} + \text{h.c.} \rangle &= \langle D\bar{D} + \text{h.c.} \rangle = \langle S\bar{S} + \text{h.c.} \rangle \\ &= \mathcal{O}(\Lambda_D^3) \end{aligned}$$



$$SU(3)_L \times SU(3)_R \rightarrow SU(3)_V$$

U(1)<sub>D</sub> also breaks due to this condensation

## 2. A Model of Chiral Composite ADM

SSB of  $U(1)_D \rightarrow$  massive dark photon ! [2]

Chiral Lagrangian

$$\mathcal{L} = \frac{f_\pi'^2}{4} \text{tr} [(D_\mu U)(D^\mu U)^\dagger] + \text{meson mass term}$$

$$U(x) = \exp \left[ \frac{i}{f_\pi'} \sum_{i=1}^8 \pi'_i(x) \lambda_i \right]$$

$$D_\mu U(x) = \partial_\mu U(x) - ie_D A'_\mu \lambda_3 U(x) +iae_D A'_\mu U(x) \lambda_3$$

## 2. A Model of Chiral Composite ADM

From chiral Lagrangian,

$$m_{\gamma'} = e_D(1 - a)f'_\pi \sim e_D(1 - a)\Lambda_D \quad (\pi'_3 \text{ is would-be NG boson})$$



Dark Higgs VEV tuning is needed in addition to  $\Lambda_D$ !

Nothing is said about dark Higgs in [1] !

e.g. Thermal history, 4-point coupling with SM Higgs

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### 3. Conclusion

We constructed the new composite ADM model that solves the following problems.

- 
- Dark Higgs VEV tuning is needed in addition to  $\Lambda_D$  !
  - Nothing is said about dark Higgs in [1] !
    - e.g. Thermal history, 4-point coupling with SM Higgs

## References

- [1]: M. Ibe, A. Kamada, S. Kobayashi and W. Nakano, JHEP 11 (2018)  
203
- [2]: K. Harigaya and Y. Nomura, Phys. Rev. D 94 no.3 (2016), 035013

Thank you !

# BACK UP

# ADM mass

$$m_{\text{DM}} \simeq \frac{\Omega_{\text{DM}}}{\Omega_B} \frac{A_B}{A_{\text{SM}}} \frac{A_{\text{SM}}}{A_{\text{DM}}} \times m_N$$

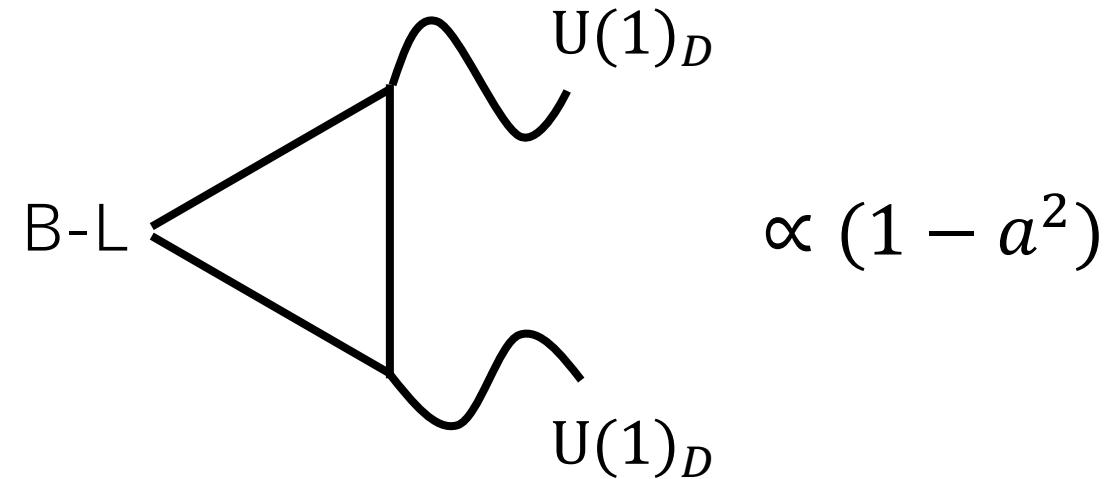
$$A_B \equiv \sum_{i \in \text{SM}} q_{i,B} (n_i - \bar{n}_i), \quad A_{\text{SM}} \equiv \sum_{i \in \text{SM}} q_{i,B-L} (n_i - \bar{n}_i),$$

$$A_{\text{DM}} \equiv \sum_{i \in \text{DM}} q_{i,B-L} (n_i - \bar{n}_i)$$

# Anomaly in Chiral Composite ADM model

	$SU(3)_D$	$U(1)_D$	$U(1)_{B-L}$
$U$	<b>3</b>	1	1/3
$D$	<b>3</b>	-1	1/3
$S$	<b>3</b>	0	1/3
$\bar{U}$	<b><math>\bar{3}</math></b>	$-a$	-1/3
$\bar{D}$	<b><math>\bar{3}</math></b>	$a$	-1/3
$\bar{S}$	<b><math>\bar{3}</math></b>	0	-1/3

$U(1)_{B-L} \times [U(1)_D]^2$  global anomaly exists,



This anomaly does not affect the ADM scenario, unless a dark helical magnetic field exists.

# Phenomenology of Dark Matter

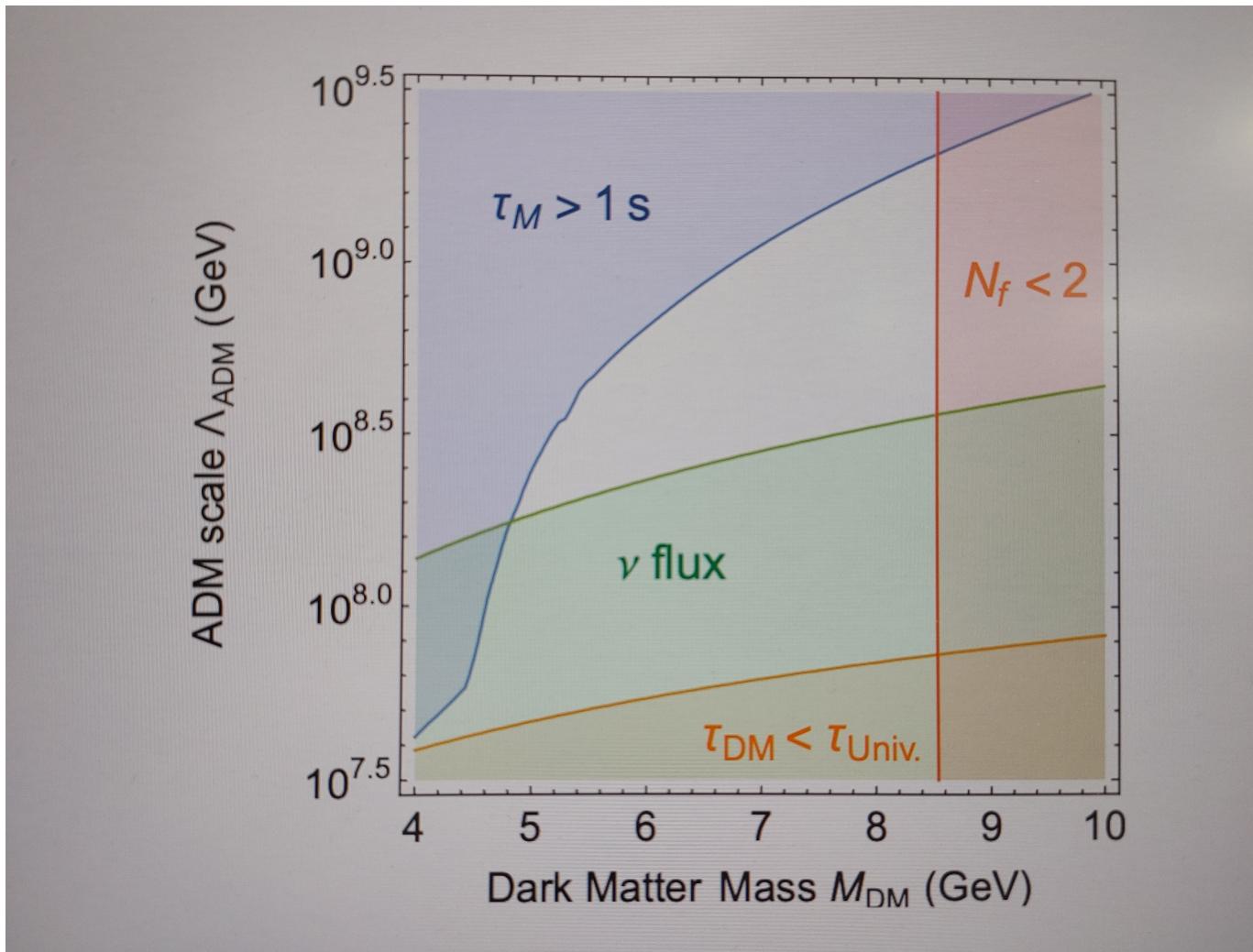
These dark baryons decay due to B-L portal,

$$\mathcal{L}_{B-L \text{ portal}} = \mathcal{O}_{DS} LH/M^3 \quad \mathcal{O}_{DS} = UDS + \bar{U}^\dagger \bar{D}^\dagger S$$

but their lifetime is longer than the age of the Universe  
for  $M \gtrsim 10^{8.2} \text{ GeV}$ ,

$$\tau_{\text{DM}} \gtrsim 10^{21} \text{ sec} = 10^{14} \text{ year}$$

# Lifetime of the DM



# Formula of the Dark Baryon

	$SU(3)_D$	$U(1)_D$	$U(1)_{B-L}$	$U(1)_{3D}$	$U(1)_{8D}$
$U$	<b>3</b>	1	1/3	1	1
$D$	<b>3</b>	-1	1/3	-1	1
$S$	<b>3</b>	0	1/3	0	-2
$\bar{U}$	<b><math>\bar{3}</math></b>	$-a$	-1/3	-1	-1
$\bar{D}$	<b><math>\bar{3}</math></b>	$a$	-1/3	1	-1
$\bar{S}$	<b><math>\bar{3}</math></b>	0	-1/3	0	2

$$B' = \begin{pmatrix} \Sigma'_3 + \Lambda'/\sqrt{3} & \sqrt{2} \Sigma'_1 & \sqrt{2} p' \\ \sqrt{2} \Sigma'_2 & -\Sigma'_3 + \Lambda'/\sqrt{3} & \sqrt{2} n' \\ \sqrt{2} \Xi'_2 & \sqrt{2} \Xi'_1 & -2 \Lambda'/\sqrt{3} \end{pmatrix} \quad \left( B_{\text{SM}} = \begin{pmatrix} \Sigma^0 + \Lambda/\sqrt{3} & \sqrt{2} \Sigma^+ & \sqrt{2} p \\ \sqrt{2} \Sigma^- & -\Sigma^0 + \Lambda/\sqrt{3} & \sqrt{2} n \\ \sqrt{2} \Xi^- & \sqrt{2} \Xi^0 & -2 \Lambda/\sqrt{3} \end{pmatrix} \right)$$

# Formula of the Dark Meson

	$SU(3)_D$	$U(1)_D$	$U(1)_{B-L}$	$U(1)_{3D}$	$U(1)_{8D}$
$U$	<b>3</b>	1	1/3	1	1
$D$	<b>3</b>	-1	1/3	-1	1
$S$	<b>3</b>	0	1/3	0	-2
$\bar{U}$	<b><math>\bar{3}</math></b>	$-a$	-1/3	-1	-1
$\bar{D}$	<b><math>\bar{3}</math></b>	$a$	-1/3	1	-1
$\bar{S}$	<b><math>\bar{3}</math></b>	0	-1/3	0	2

$$M' = \begin{pmatrix} \eta'/\sqrt{3} & \sqrt{2} \pi' & \sqrt{2} K'^1 \\ \sqrt{2} \pi'^{\dagger} & \eta'/\sqrt{3} & \sqrt{2} K'^2 \\ \sqrt{2} K'^1\dagger & \sqrt{2} K'^2\dagger & -2 \eta'/\sqrt{3} \end{pmatrix} \quad M_{\text{SM}} = \begin{pmatrix} \pi^0 + \eta/\sqrt{3} & \sqrt{2} \pi^+ & \sqrt{2} K^+ \\ \sqrt{2} \pi^- & -\pi^0 + \eta/\sqrt{3} & \sqrt{2} K^0 \\ \sqrt{2} K^- & \sqrt{2} \bar{K}^0 & -2 \eta/\sqrt{3} \end{pmatrix}$$

## Dark pion mass

Dark pion also obtains the mass,

$$m_{\pi'}^2 \sim \frac{3a \log 2}{2\pi^2} e_D^2 \Lambda_D^2$$

There is a parameter region  $(a, e_D)$  which satisfies the following relation

$$2 \times m_e < m_{\gamma'} < m_{\text{Dark meson}} < m_{\text{DM}},$$