

# Electroweak phase transition in a complex singlet extension of the Standard Model with degenerate scalars

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**arXiv: 2105.11830**

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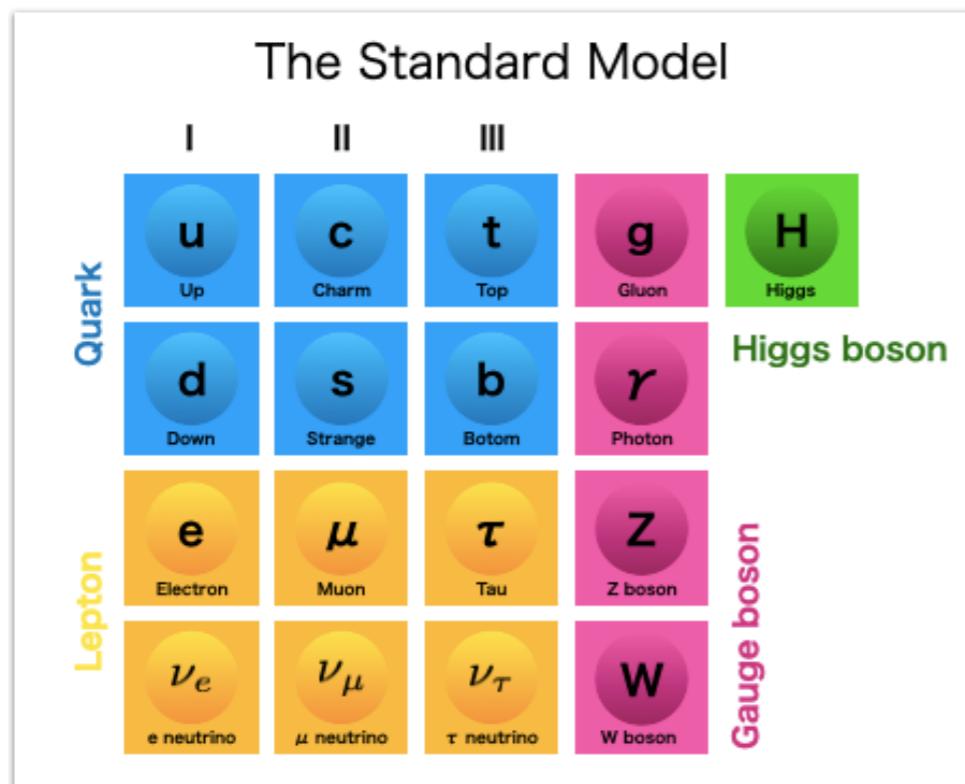
# Back ground: Baryon Asymmetry

**Baryon asymmetry:** imbalance in particles and antiparticles in the observable universe

## Sakharov conditions

**Electroweak  
baryogenesis**

1. Baryon number violation  
→ Sphaleron
2. C symmetry and CP symmetry violation  
→ Chiral gauge interaction, CKM matrix
3. Interaction out of thermal equilibrium  
→ Strong 1st order phase transition



The parameters of the SM do not satisfy Sakharov conditions.

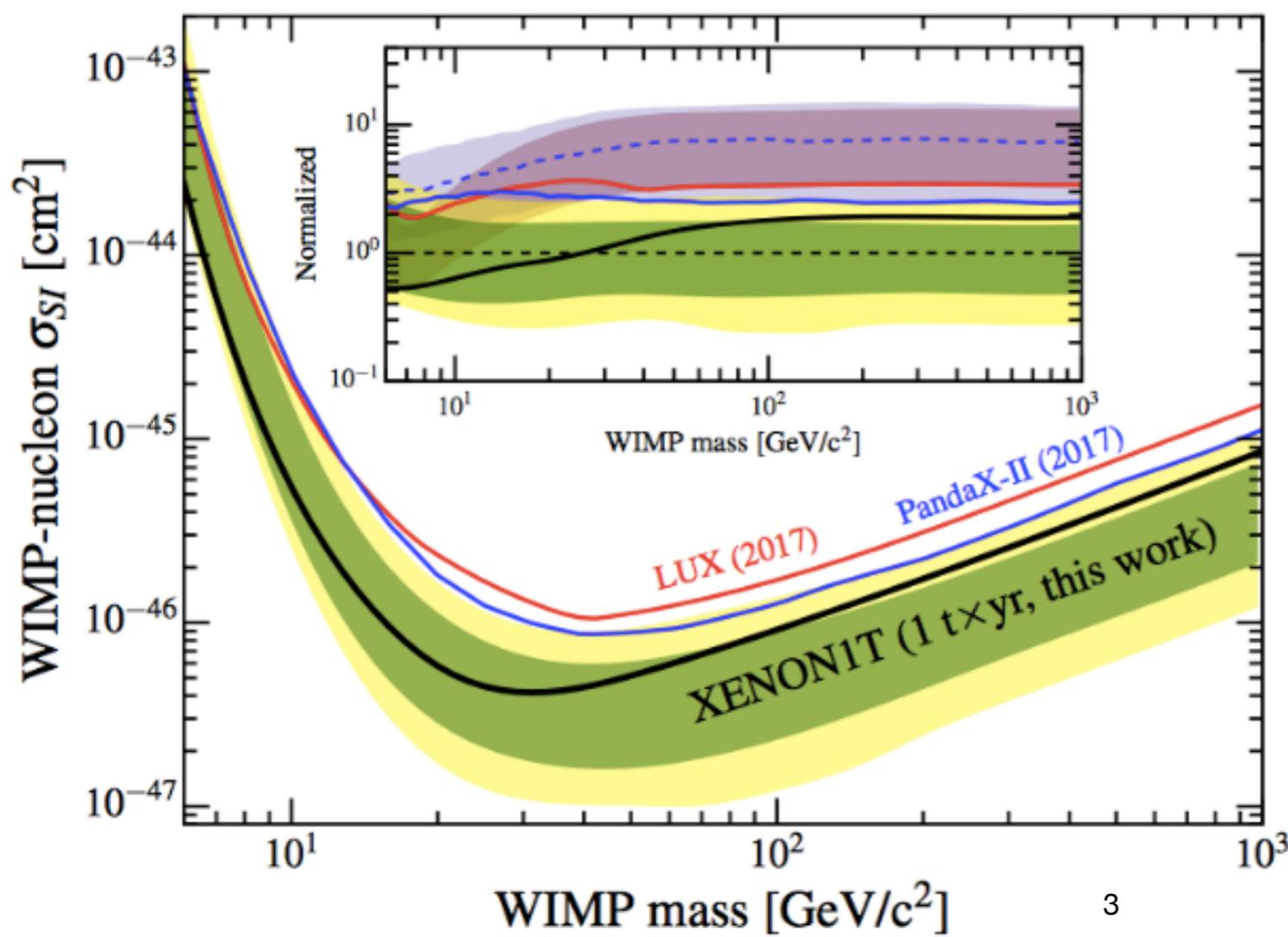
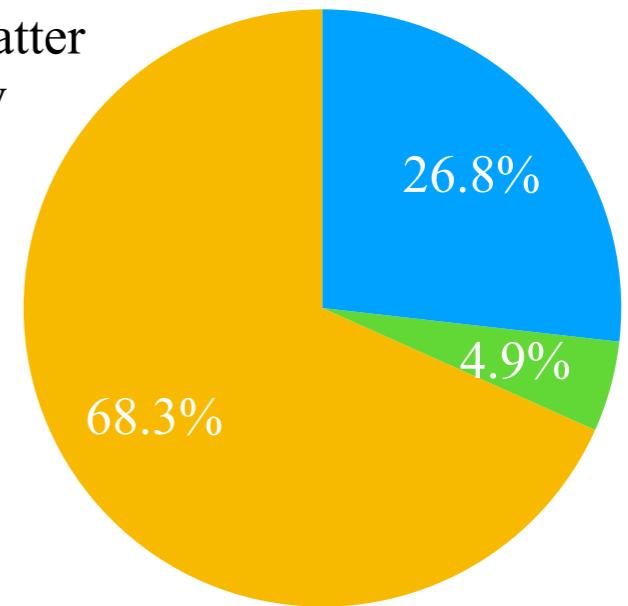
⇒ We need to extend the SM !

# Back ground: Dark matter

## The nature of dark matter

- 1, massive
- 2, no electric charge
- 3, stable

● Dark matter  
● Ordinary matter  
● Dark energy



**Strict restriction is imposed  
on the models including DM.**

XENON Collaboration  
[arXiv:1805.12562]

# Outline

- Back ground
- CxSM model definition
- Degenerate-scalar scenario
- EWPT in the degenerate-scalar scenario
- Numerical results
- Conclusion

# CxSM Model Definition

**Complex singlet extension of the SM (CxSM)** Barger et al, arXiv:0811.0393

$$V_0 = \frac{m^2}{2}|H|^2 + \frac{\lambda}{4}|H|^4 + \frac{\delta_2}{2}|H|^2|S|^2 + \frac{b_2}{2}|S|^2 + \frac{d_2}{4}|S|^4 + \left( a_1 S + \frac{b_1}{4}S^2 + \text{c.c.} \right)$$

Global U(1) and soft breaking terms (minimal set of S.B. operators to realize pNG DM)

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v + h \end{pmatrix}, \quad S = (v_S + s + i\chi)/\sqrt{2}$$

**DM**    (DM stability  $\leftrightarrow$  CP sym.)

Mass eigenstates

$$\begin{pmatrix} h \\ s \end{pmatrix} = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

Mass eigenvalues

**$h_1, h_2$**

$$\Lambda^2 \equiv \frac{d_2}{2}v_S^2 - \sqrt{2}\frac{a_1}{2v_S}$$

**DM**

$$m_{h_1, h_2}^2 = \frac{1}{2} \left( \frac{\lambda}{2}v^2 + \Lambda^2 \mp \sqrt{\left( \frac{\lambda}{2}v^2 - \Lambda^2 \right)^2 + 4 \left( \frac{\delta_2}{2}vv_S \right)^2} \right)$$

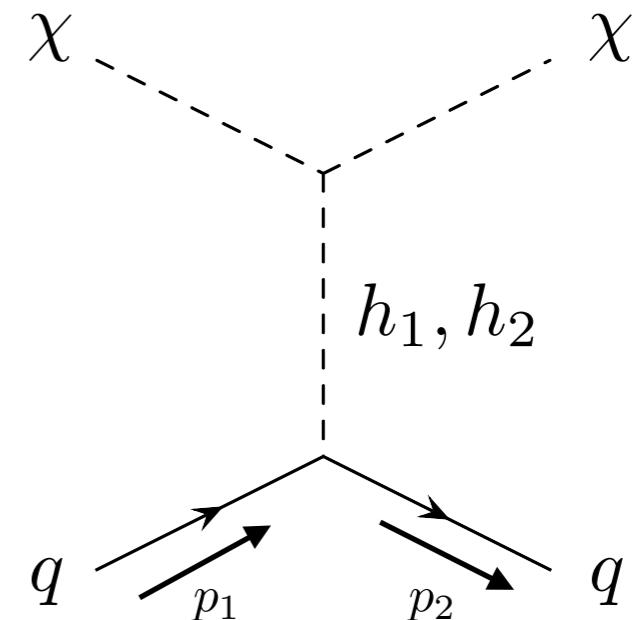
$$m_\chi^2 = -b_1 - \sqrt{2}\frac{a_1}{v_S}$$

# Degenerate-Scalar Scenario

Abe, Cho, Mawatari arXiv:2101.04887

$$i\mathcal{M}_{h_1} = -i \frac{m_f}{vv_S} \frac{m_{h_1}^2 + \frac{\sqrt{2}a_1}{v_S}}{t - m_{h_1}^2} \sin \alpha \cos \alpha \bar{u}(p_3) u(p_1),$$

$$i\mathcal{M}_{h_2} = +i \frac{m_f}{vv_S} \frac{m_{h_2}^2 + \frac{\sqrt{2}a_1}{v_S}}{t - m_{h_2}^2} \sin \alpha \cos \alpha \bar{u}(p_3) u(p_1),$$



$$\begin{aligned} i(\mathcal{M}_{h_1} + \mathcal{M}_{h_2}) &= i \frac{m_f}{vv_S} \left( -\frac{m_{h_1}^2 + \frac{\sqrt{2}a_1}{v_S}}{t - m_{h_1}^2} + \frac{m_{h_2}^2 + \frac{\sqrt{2}a_1}{v_S}}{t - m_{h_2}^2} \right) \sin \alpha \cos \alpha \bar{u}(p_3) u(p_1) \\ &\simeq i \frac{m_f}{vv_S} \sin \alpha \cos \alpha \bar{u}(p_3) u(p_1) \\ &\quad \times \left\{ \left( \frac{\sqrt{2}a_1}{v_S} + t \right) \left( \frac{1}{m_{h_1}^2} - \frac{1}{m_{h_2}^2} \right) + \frac{\sqrt{2}a_1}{v_S} t \left( \frac{1}{m_{h_1}^4} - \frac{1}{m_{h_2}^4} \right) \right\} @ t \rightarrow 0 \end{aligned}$$

$$\simeq i \frac{m_f}{vv_S} \sin \alpha \cos \alpha \bar{u}(p_3) u(p_1) \frac{\sqrt{2}a_1}{v_S} \left( \frac{1}{m_{h_1}^2} - \frac{1}{m_{h_2}^2} \right)$$

$$\simeq 0 \quad (m_{h_1} \sim m_{h_2})$$

# EWPT in the degenerate-scalar scenario

Strong 1st order phase transition  
(SFOEWPT)



$$\frac{v_c}{T_c} \gtrsim 1$$

$T_c$  : critical temperature  
 $v_c$  : higgs vev at  $T_c$

[Two calculation schemes on the scalar potential (**gauge independent**)]

**HT potential**  $V^{\text{HT}}(\varphi, \varphi_S; T) = V_0(\varphi, \varphi_S) + \frac{1}{2} (\Sigma_H \varphi^2 + \Sigma_S \varphi_S^2) T^2$

**PRM scheme**  $\frac{\partial V_{\text{eff}}(\varphi, \xi)}{\partial \xi} = -C(\varphi, \xi) \frac{\partial V_{\text{eff}}(\varphi, \xi)}{\partial \varphi}$

M. J. Ramsey-Musolf, JHEP 07 (2011), 029.

$$V_0 \left( 0, v_{S, \text{tree}}^{\text{sym}} \right) + V_1 \left( 0, v_{S, \text{tree}}^{\text{sym}} ; T \right) = V_0 \left( v_{\text{tree}}, v_{S, \text{tree}} \right) + V_1 \left( v_{\text{tree}}, v_{S, \text{tree}} ; T \right)$$

$v_C$ ,  $v_{SC}$  and  $v_{SC}^{\text{sym}}$  are determined by the use of  $V^{HT}$

[Two resummation methods in evaluating one-loop effective potential (**gauge dependent**)]

$$V_{\text{eff}}(\varphi, \varphi_S; T) = V_0(\varphi, \varphi_S; T) + \sum_i n_i \left[ V_{\text{CW}}(\bar{m}_i^2) + \frac{T^4}{2\pi^2} I_{B,F} \left( \frac{\bar{m}_i^2}{T^2} \right) \right]$$

**Parwani scheme** Replace  $\bar{m}^2$  with thermally corrected field depending masses  $\bar{M}^2$

**AE scheme**  $V_{\text{daisy}}(\varphi, \varphi_S; T) = \sum_{\substack{i=h_{1,2}, \chi \\ W_L, Z_L, \gamma_L}} -n_i \frac{T}{12\pi} \left[ (\bar{M}_i^2)^{3/2} - (\bar{m}_i^2)^{3/2} \right]$

# EWPT in the degenerate-scalar scenario

Parametrize the two scalar fields using radial coordinates as

$$\varphi = z \cos \gamma, \varphi_S = z \sin \gamma + v_S^{\text{sym}}$$

HT potential

$$V^{\text{HT}}(z, \gamma; T) = c_0 + c_1 z + (c_2 + c'_2 T^2) z^2 - c_3 z^3 + c_4 z^4$$

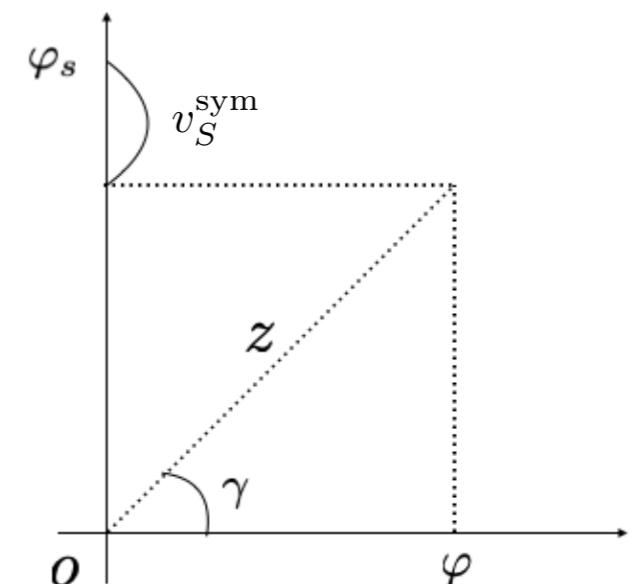
In the case of first-order EWPT

$$T_C \simeq \sqrt{\frac{1}{2\Sigma_H} \left( -m^2 - \frac{(v_{SC}^{\text{sym}})^2}{2} \delta_2 \right)}, \quad v_C = \lim_{T \nearrow T_C} v(T)$$

$$v_{SC} = \lim_{T \nearrow T_C} v_S(T)$$

$$v_{SC}^{\text{sym}} = \lim_{T \searrow T_C} v_S(T)$$

$$v_C \simeq \sqrt{\frac{2\delta_2 (v_{SC}^{\text{sym}})^2}{\lambda}} \left( 1 - \frac{v_{SC}}{v_{SC}^{\text{sym}}} \right)$$



Condition of SFOEWPT

$$\frac{v_c}{T_c} > 1$$

# EWPT in the degenerate-scalar scenario

Tree level potential  $V_0 = \frac{m^2}{2}|H|^2 + \frac{\lambda}{4}|H|^4 + \frac{\delta_2}{2}|H|^2|S|^2 + \frac{b_2}{2}|S|^2 + \frac{d_2}{4}|S|^4 + \left(a_1 S + \frac{b_1}{4}S^2 + \text{c.c.}\right)$

$$T_C \simeq \sqrt{\frac{1}{2\Sigma_H} \left( -m^2 - \frac{(v_{SC}^{\text{sym}})^2}{2} \delta_2 \right)},$$

$$v_C \simeq \sqrt{\frac{2\delta_2 (v_{SC}^{\text{sym}})^2}{\lambda}} \left( 1 - \frac{v_{SC}}{v_{SC}^{\text{sym}}} \right)$$

Condition of SFOEWPT

$$\frac{v_c}{T_c} > 1$$

About  $T_C$

$T_C \rightarrow$  small,  $\delta_2 \rightarrow$  positive and sizable

$$\delta_2 = \frac{2}{vv_S} (m_{h_1}^2 - m_{h_2}^2) \sin \alpha \cos \alpha$$

$v_S \rightarrow$  small,  $\alpha \rightarrow$  the maximal mixing  $\frac{\pi}{4}$

# EWPT in the degenerate-scalar scenario

$$V_0 = \frac{m^2}{2}|H|^2 + \frac{\lambda}{4}|H|^4 + \frac{\delta_2}{2}|H|^2|S|^2 + \frac{b_2}{2}|S|^2 + \frac{d_2}{4}|S|^4 + \left( a_1 S + \frac{b_1}{4}S^2 + \text{c.c.} \right)$$

$$T_C \simeq \sqrt{\frac{1}{2\Sigma_H} \left( -m^2 - \frac{(v_{SC}^{\text{sym}})^2}{2} \delta_2 \right)},$$

$$v_C \simeq \sqrt{\frac{2\delta_2(v_{SC}^{\text{sym}})^2}{\lambda} \left( 1 - \frac{v_{SC}}{v_{SC}^{\text{sym}}} \right)}$$

Condition of SFOEWPT

$$\frac{v_c}{T_c} > 1$$

About  $v_C$

$v_C \rightarrow$  large with an amplification factor  $(v_{SC}^{\text{sym}})^2 (1 - v_{SC}/v_{SC}^{\text{sym}})$

$$(v_{SC}^{\text{sym}})^3 + Av_{SC}^{\text{sym}} + B = 0$$

$$A = 2(b_1 + b_2 + 2\Sigma_S)/d_2$$

$$B = 4\sqrt{2}a_1/d_2$$

$v_{SC}^{\text{sym}}$  is scaled by  $1/\sqrt{d_2}$

$\therefore d_2 \rightarrow$  small

$$d_2 = \frac{2}{v_S^2} \left[ m_{h_1}^2 + (m_{h_2}^2 - m_{h_1}^2) \cos^2 \alpha + \frac{\sqrt{2}a_1}{v_S} \right] \simeq \frac{2}{v_S^2} \left[ m_{h_1}^2 + \frac{\sqrt{2}a_1}{v_S} \right] \quad a_1 < 0$$

(1) large  $\delta_2$  with a positive sign i.e.,  $|\alpha| \simeq \frac{\pi}{4}$  and  $v_S < 1$  GeV

(2) small  $d_2$  i.e.,  $a_1 < 0$  with its moderate absolute value

# Numerical results

Two benchmark points

the varying parameter

Inputs	$v$ [GeV]	$m_{h_1}$ [GeV]	$m_{h_2}$ [GeV]	$\alpha$ [rad]	$a_1$ [GeV $^3$ ]	$v_S$ [GeV]	$m_\chi$ [GeV]
BP1	246.22	125	124	$\pi/4$	-6576.17	0.6	62.5
BP2	246.22	125	126	$-\pi/4$	-6682.25	0.6	62.5

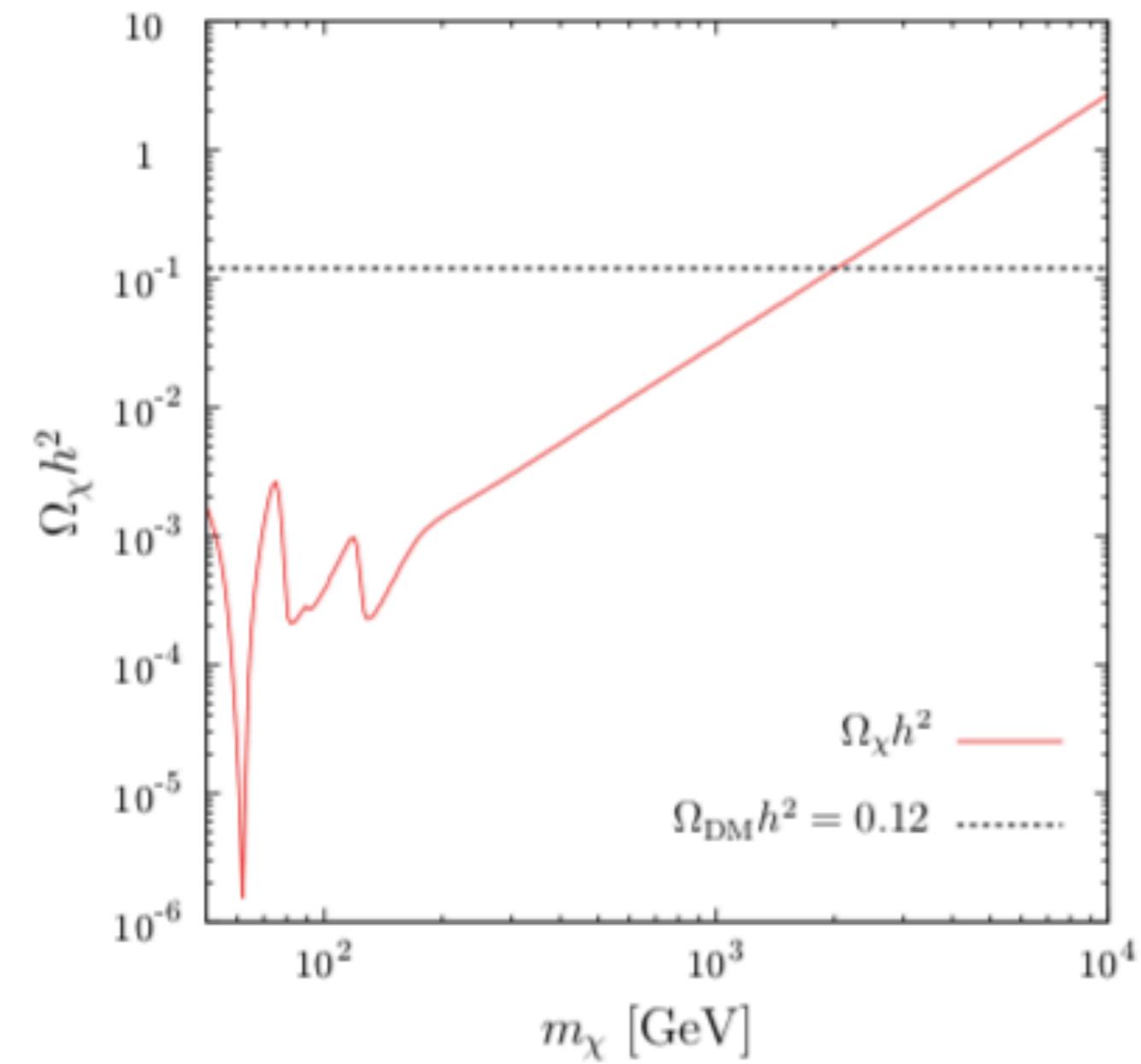
  

Outputs	$m^2$ [GeV $^2$ ]	$b_1$ [GeV $^2$ ]	$b_2$ [GeV $^2$ ]	$\lambda$	$a_1$ [GeV $^3$ ]	$d_2$	$\delta_2$
BP1	$-(124.5)^2$	$-(107.7)^2$	$-(178.0)^2$	0.511	-6576.17	1.77	1.69
BP2	$-(125.5)^2$	$-(108.8)^2$	$-(178.4)^2$	0.520	-6682.25	1.70	1.59

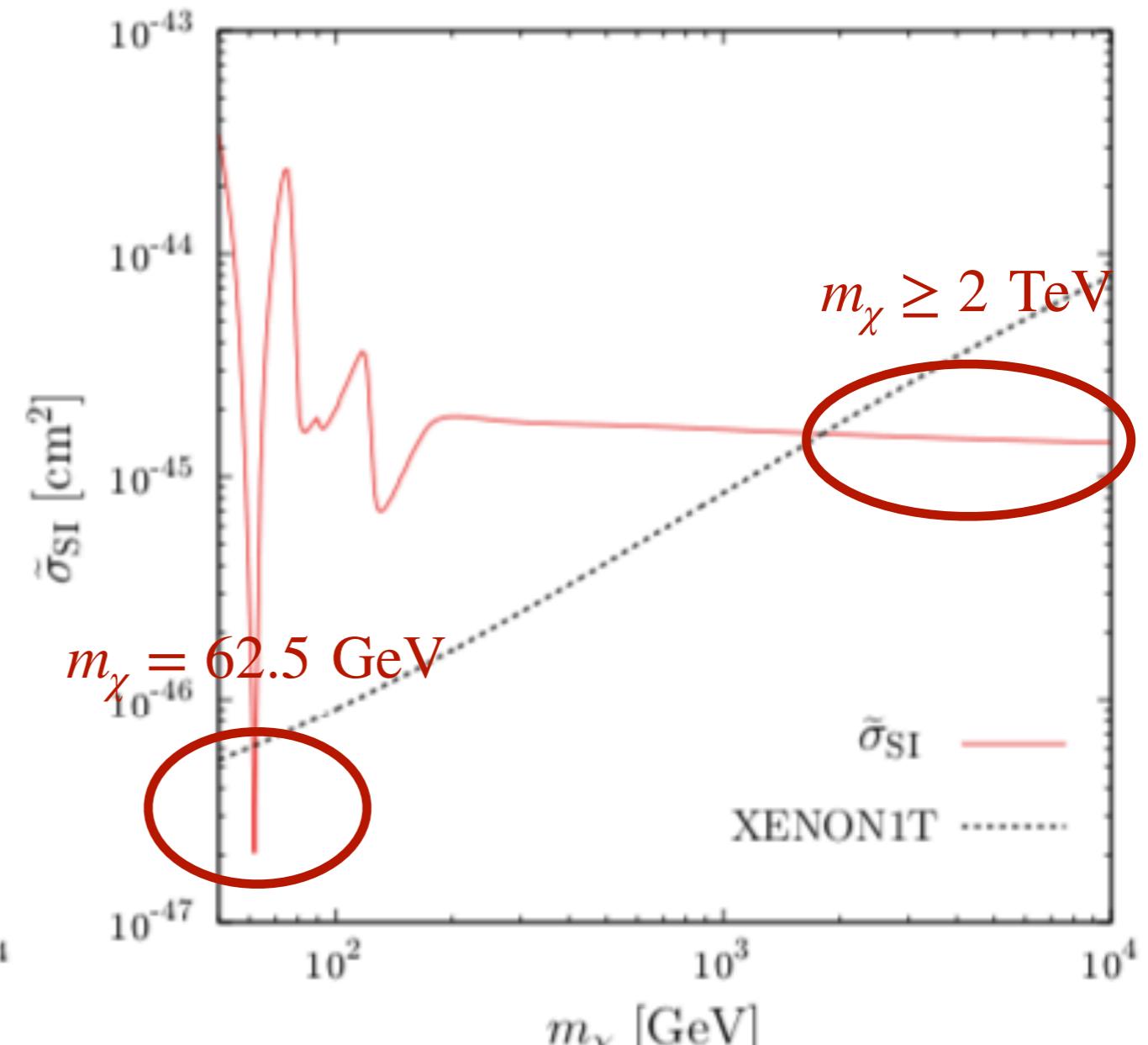
Calculate the DM relic density  $\Omega_\chi h^2$  and SI cross section with the nucleons  $\sigma_{\text{SI}}$  in BP1.

(For the moment,  $m_\chi$  is treated as the varying parameter.)

# Numerical results

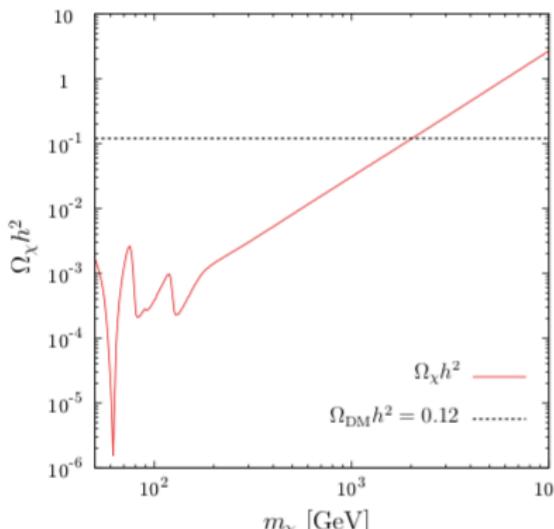


DM relic density  $\Omega_\chi h^2$

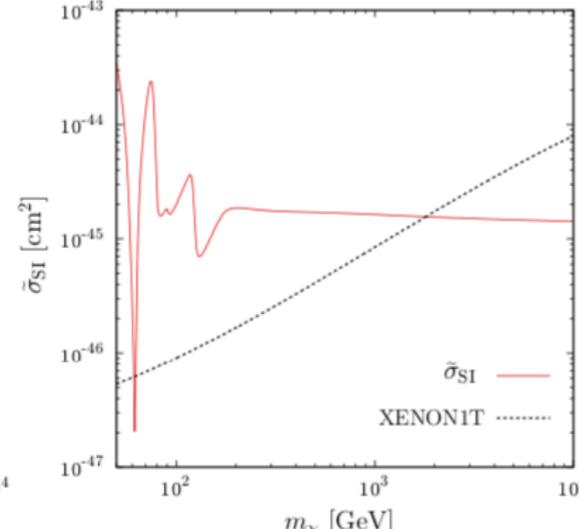


SI scattering cross section  $\tilde{\sigma}_{\text{SI}}$

# Numerical results



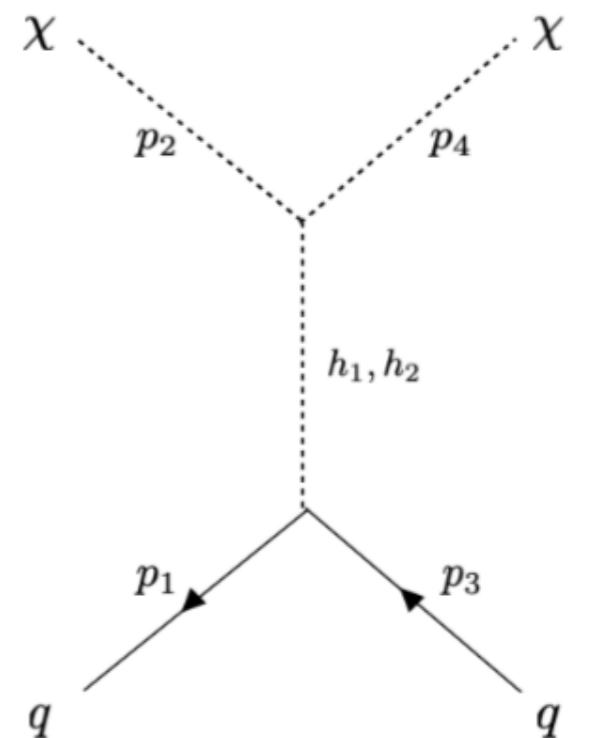
DM relic density  $\Omega_\chi h^2$



SI scattering cross section  $\sigma_{\text{SI}}$

$$\sigma_{\text{SI}} \propto \sin^2 \alpha \cos^2 \alpha \left( \frac{1}{m_{h_1}^2} - \frac{1}{m_{h_2}^2} \right)^2 \frac{a_1^2}{v_S^4} = \frac{\delta_2^2 v^2}{4m_{h_1}^4 m_{h_2}^4} \frac{a_1^2}{v_S^2}$$

$\delta_2 = \frac{2}{vv_S} (m_{h_1}^2 - m_{h_2}^2) \sin \alpha \cos \alpha$



The scattering of  
dark matter  $\chi$  and quark  $q$

Strong 1st EWPT  
 $\delta_2 \rightarrow$  large  
 $v_S \rightarrow$  small  
 (less than 1 GeV)

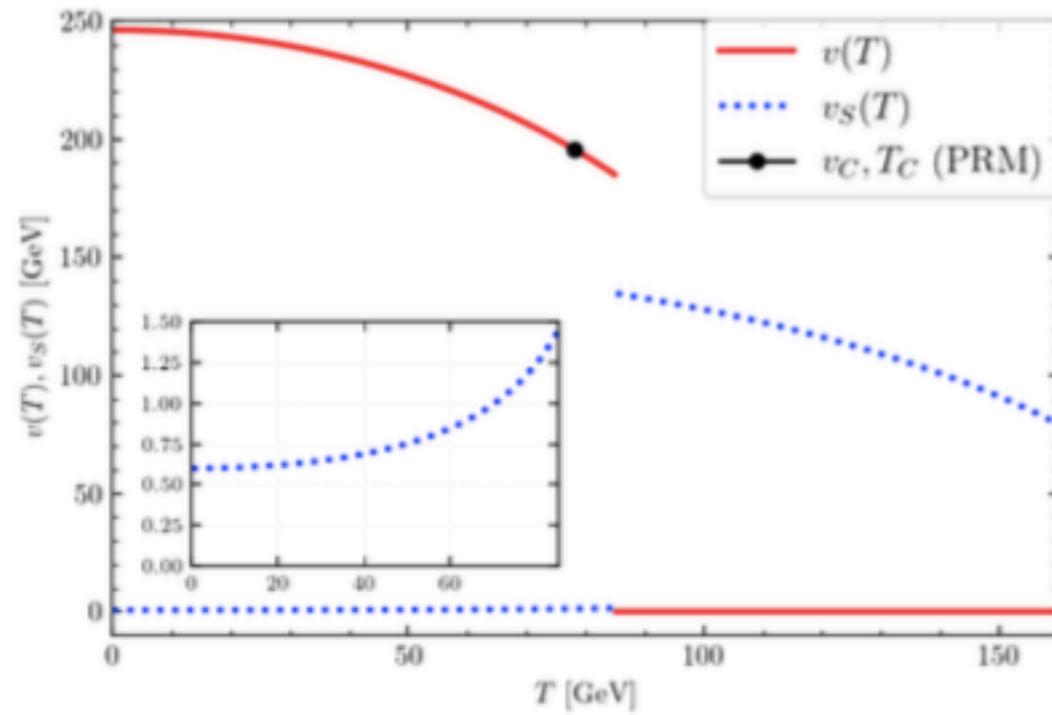
The core of the cancellation mechanism in the degenerate-scalar scenario:

The suppression of  $\delta_2$  owing to  $m_{h_1} \simeq m_{h_2}$  with moderate values of  $v_S$ .

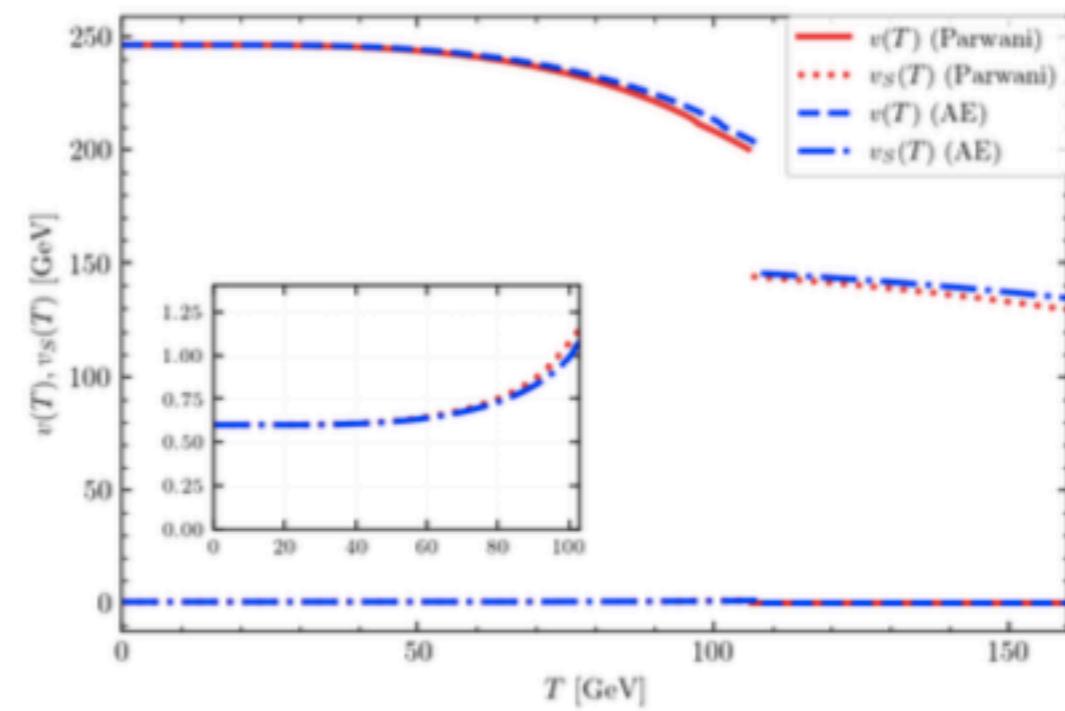
The conditions for the strong 1st EWPT is incompatible with the suppression mechanism.

# Numerical results

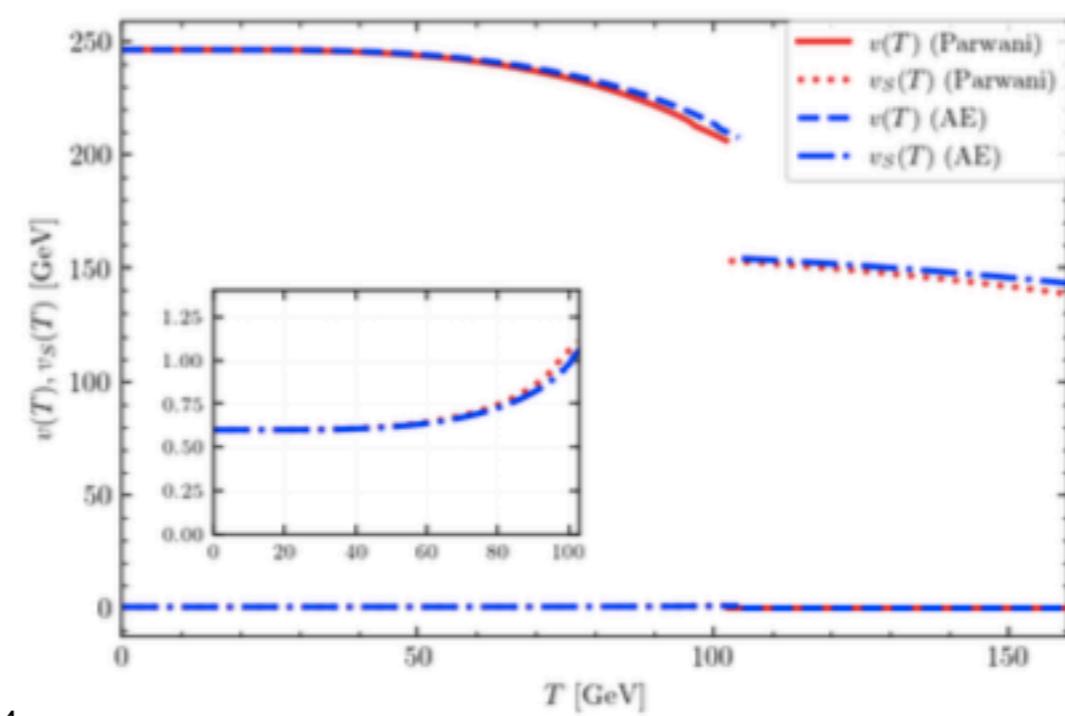
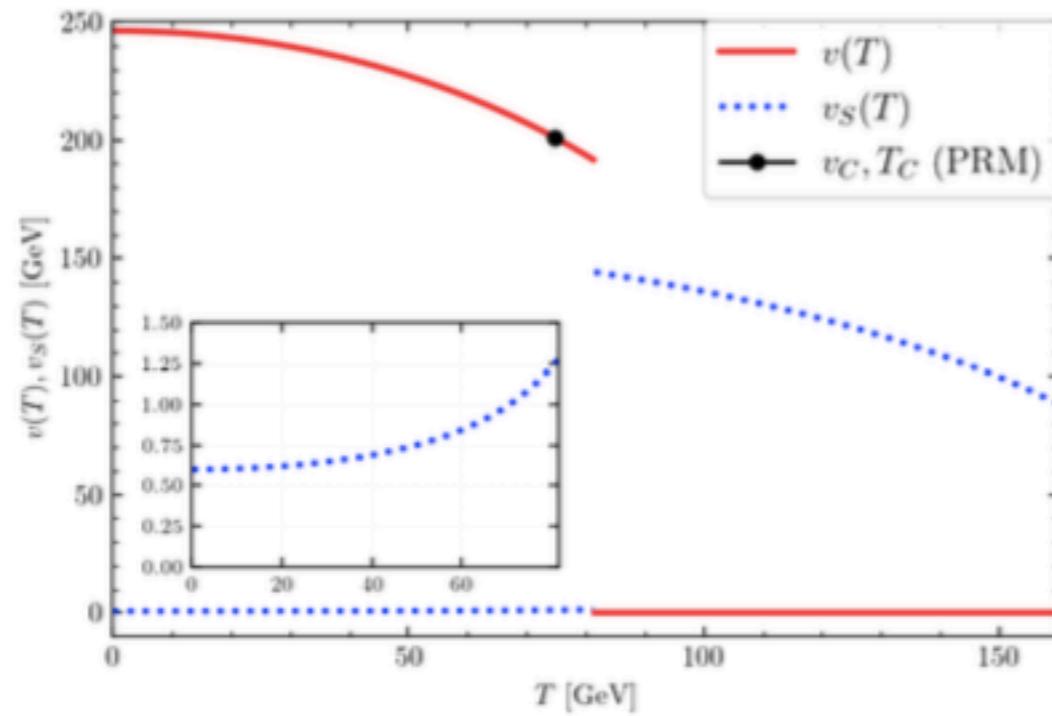
HT/PRM



Parwani/AE



BP2

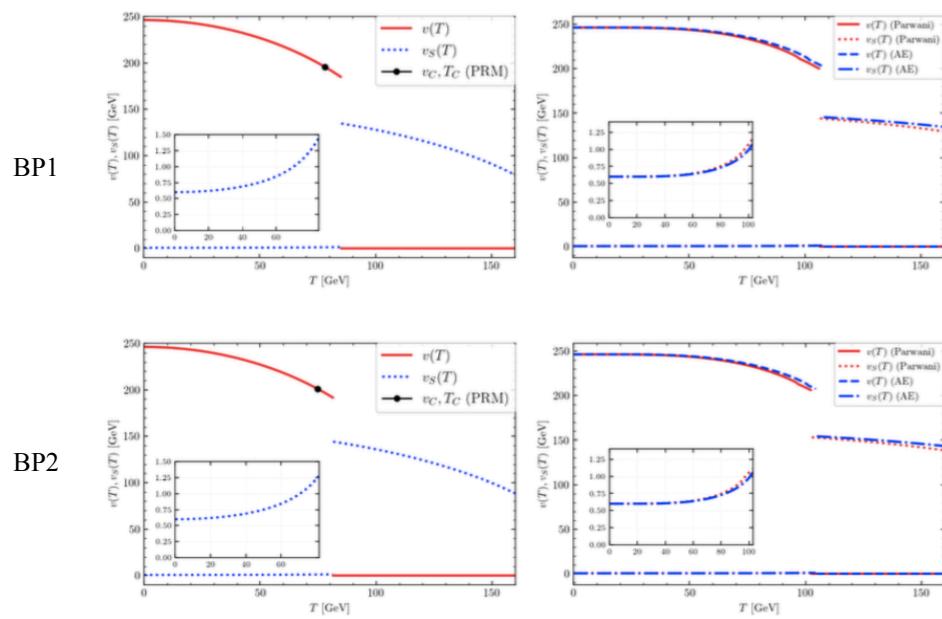


# Numerical results

HT/PRM

Parwani/AE

Ex) BP1



BP1				
Scheme	HT	PRM	Parwani	AE
$v_C/T_C$	$\frac{184.4}{85.3} = 2.2$	$\frac{195.6}{78.2} = 2.5$	$\frac{201.5}{106.8} = 1.9$	$\frac{202.7}{107.8} = 1.9$
$v_{SC}$ [GeV]	1.5	1.2	1.2	1.2
$v_{SC}^{\text{sym}}$ [GeV]	134.6	137.3	144.8	145.3

Strong 1st PT !

The consequences found in BP1 all apply to BP2 as well.

Strong first-order EWPT in the degenerate-scalar scenario is possible  
in the both cases  $m_{h_1} > m_{h_2}$  and  $m_{h_1} < m_{h_2}$ .

# Summary

We adopted CxSM as a model to explain dark matter, and discussed it from the view point of the strong 1st order phase transition necessary to explain baryon asymmetry.

We analytically showed that the suppression of  $\sigma_{\text{SI}}$  driven by the smallness of  $\delta_2$ , which could be realized by a ratio of the mass difference of two scalars and the singlet VEV  $v_S$ , conflicts with one of the necessary conditions for the strong first-order EWPT.

Our numerical analysis also confirms that  $\sigma_{\text{SI}}$  is not suppressed by the degenerated scalar masses. Nonetheless, the allowed regions are still present at around  $m_\chi = 62.5$  GeV and 2 TeV.

We analyzed EWPT in the viable DM regions by four different calculation schemes: HT, PRM, Parwani, and AE and all the calculations indicate the strong first-order EWPT.

# Back up

# CxSM Model Definition

The general scalar potential

$$V = \frac{m^2}{2}|H|^2 + \frac{\lambda}{4}|H|^4 + \frac{\delta_2}{2}|H|^2|S|^2 + \frac{b_2}{2}|S|^2 + \frac{d_2}{4}|S|^4 \\ + \left( a_1 S + \frac{\delta_1}{4}|H|^2 S + \frac{\delta_3}{4}|H|^2 S^2 + \frac{b_1}{4}S^2 + \frac{c_1}{6}S^3 + \frac{c_2}{6}S|S|^2 + \frac{d_1}{8}S^4 + \frac{d_3}{8}S^2|S|^2 + \text{c.c.} \right)$$

The minimization condition

Mixing angle  $\alpha$

$$-m^2 = \frac{\lambda}{2}v^2 + \frac{\delta_2}{2}v_S^2, \\ -b_2 = \frac{\delta_2}{2}v^2 + \frac{d_2}{2}v_S^2 + b_1 + 2\sqrt{2}\frac{a_1}{v_S}$$

$$\tan 2\alpha = 2\frac{\frac{\delta_2}{2}vv_S}{\frac{\lambda}{2}v^2 - \Lambda^2}, \quad \cos 2\alpha = \frac{\frac{\lambda}{2}v^2 - \Lambda^2}{m_{h_1}^2 - m_{h_2}^2}$$

Mass eigenvalues  $m_{h_1, h_2}^2 = \frac{1}{2} \left( \frac{\lambda}{2}v^2 + \Lambda^2 \mp \frac{\frac{\lambda}{2}v^2 - \Lambda^2}{\cos 2\alpha} \right)$

$$= \frac{1}{2} \left( \frac{\lambda}{2}v^2 + \Lambda^2 \mp \sqrt{\left( \frac{\lambda}{2}v^2 - \Lambda^2 \right)^2 + 4 \left( \frac{\delta_2}{2}vv_S \right)^2} \right)$$

# CxSM Model Definition

Scalar trilinear interactions

$$\mathcal{L}_S = -\frac{1}{2v_S} \left\{ \left( m_{h_1}^2 + \frac{\sqrt{2}a_1}{v_S} \right) \sin \alpha h_1 \chi^2 + \left( m_{h_2}^2 + \frac{\sqrt{2}a_1}{v_S} \right) \cos \alpha h_2 \chi^2 \right\}$$

Yukawa interactions

$$\mathcal{L}_Y = -\frac{m_f}{v} \bar{f} f (h_1 \cos \alpha - h_2 \sin \alpha)$$

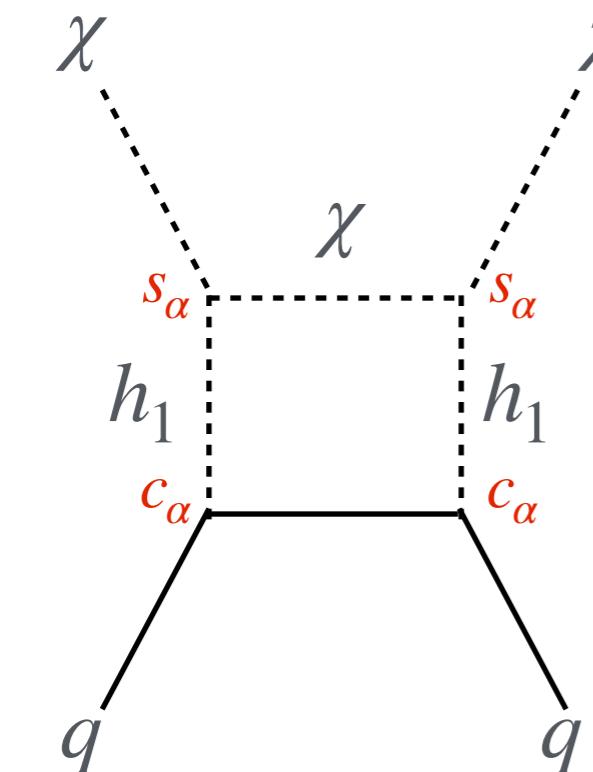
$$F(m_{h_1}) \cos^2 \alpha + F(m_{h_2}) \sin^2 \alpha \simeq F(m_{h_{\text{SM}}}) \text{ for } m_{h_1} \simeq m_{h_2} \simeq m_{h_{\text{SM}}}$$

# Degenerate-Scalar Scenario

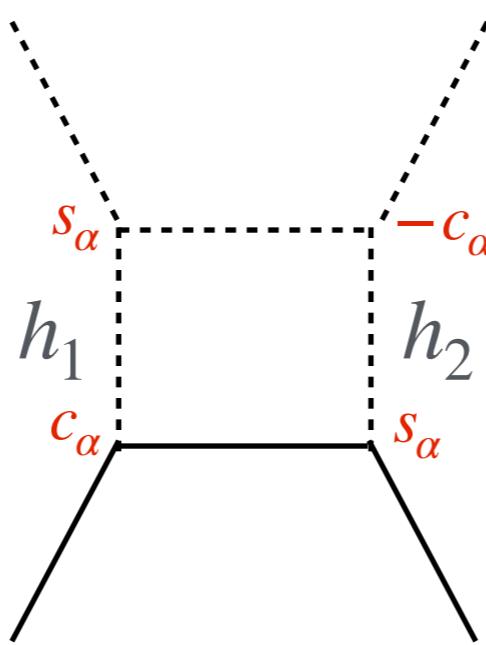
Degenerate scalar scenario@ one-loop

Azevedo et al., 1801.06105

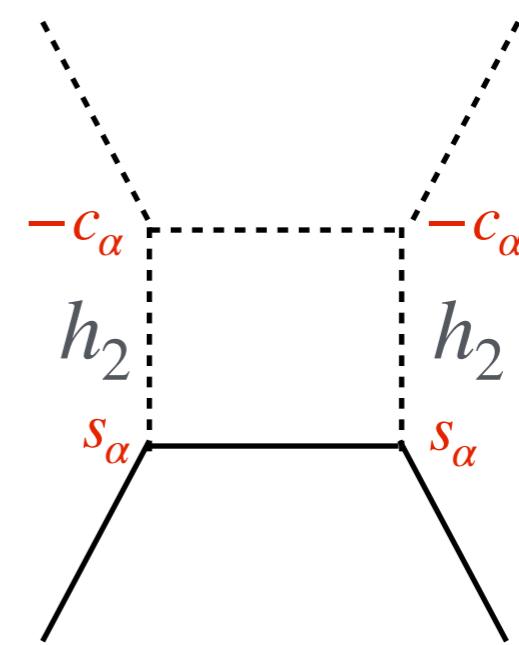
$$\sigma_{\chi N}^{\text{NLO}} = \sin 2\alpha \left( \frac{\mu_{\chi N} f_N m_N}{m_{h_1} m_{h_2}} \right)^2 \frac{m_{h_1}^2 - m_{h_2}^2}{v^3 v_S^3} \times \text{loop func.} \propto m_{h_1}^2 - m_{h_2}^2$$



$$(s_\alpha c_\alpha)^2 f(m_{h_1}, m_{h_1})$$



$$-2(s_\alpha c_\alpha)^2 f(m_{h_1}, m_{h_2})$$



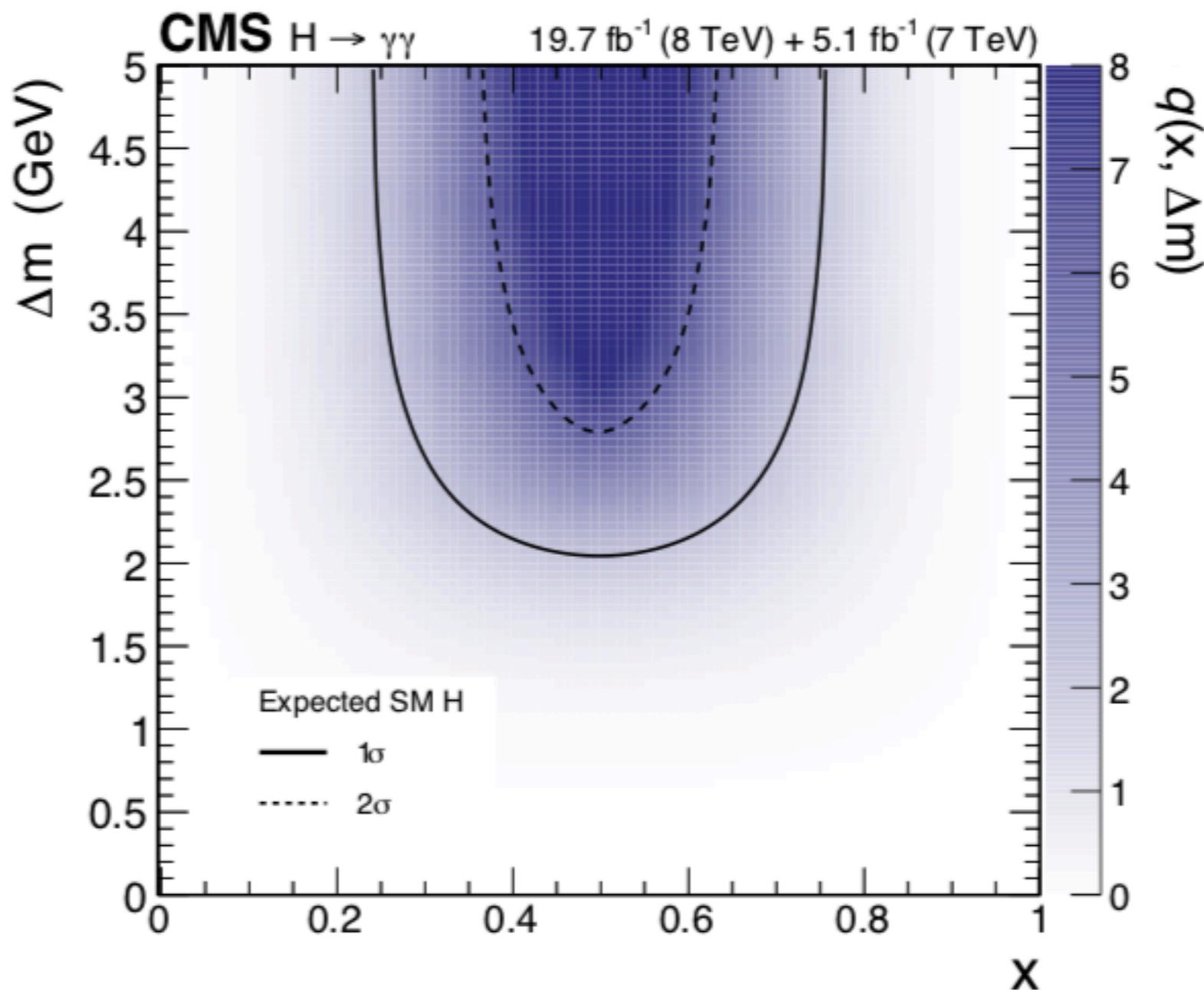
$$(-s_\alpha c_\alpha)^2 f(m_{h_2}, m_{h_2})$$

$$\text{Sum} = (s_\alpha c_\alpha)^2 (f(1,1) - f(1,2)) + (s_\alpha c_\alpha)^2 (f(2,2) - f(2,1)) \rightarrow 0 \text{ for } m_{h_1} \sim m_{h_2}$$

# Degenerate-Scalar Scenario

CMS collaboration, V. Khachatryan et al.,  
Eur. Phys. J. C 74 (2014) 3076, [1407.0558].

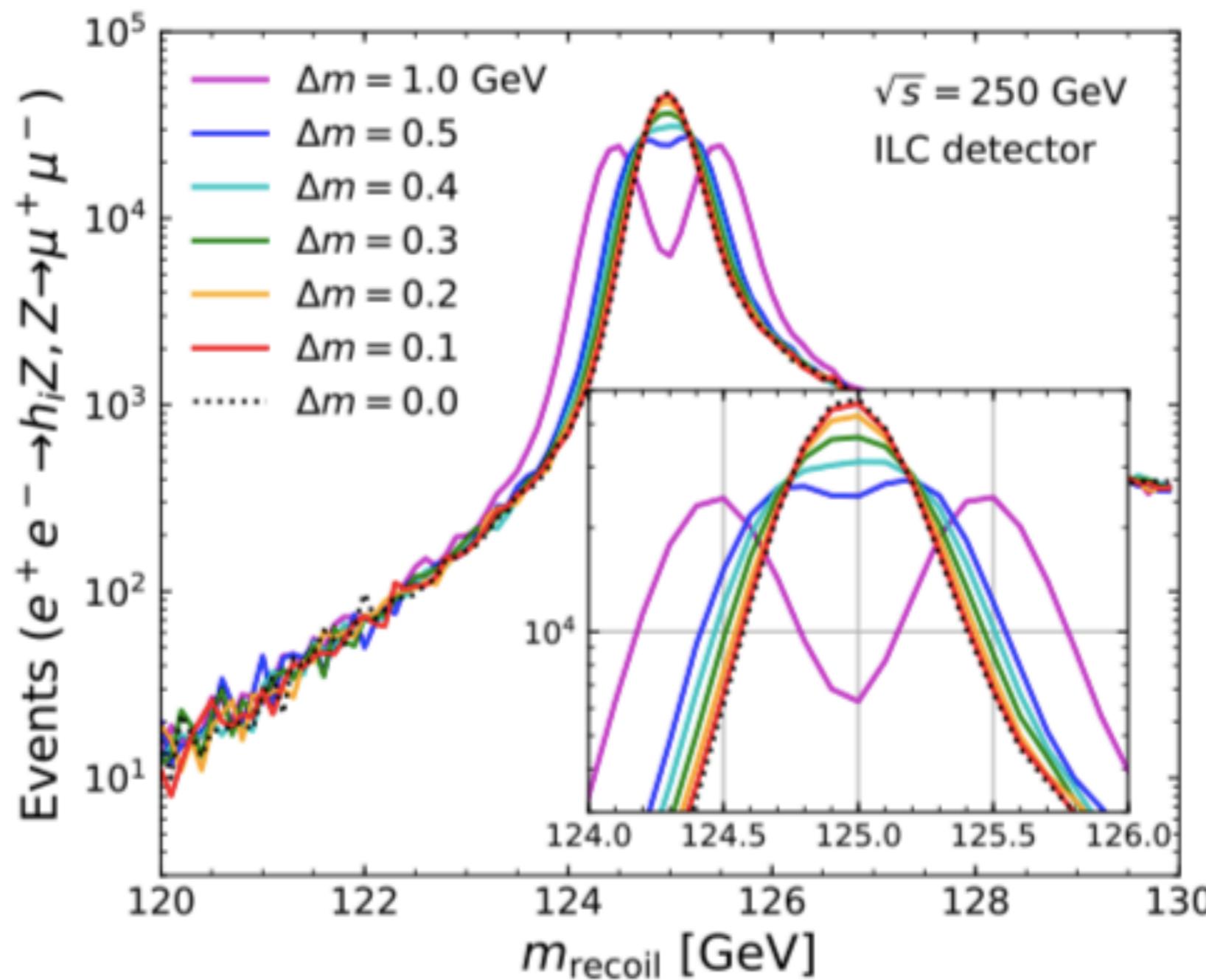
@ LHC



# Degenerate-Scalar Scenario

Sachiho Abe , Gi-Chol Cho, Kentarou Mawatari,  
arXiv:2101.04887

@ ILC



# EWPT in the degenerate-scalar scenario

**HT potential**  $V^{\text{HT}}(\varphi, \varphi_S; T) = V_0(\varphi, \varphi_S) + \frac{1}{2} (\Sigma_H \varphi^2 + \Sigma_S \varphi_S^2) T^2$  the gauge-invariant thermal masses

$$\Sigma_H = \frac{\lambda}{8} + \frac{\delta_2}{24} + \frac{3g_2^2 + g_1^2}{16} + \frac{y_t^2}{4}, \quad \Sigma_S = \frac{\delta_2 + d_2}{12}$$

**PRM scheme**  $\frac{\partial V_{\text{eff}}(\varphi, \xi)}{\partial \xi} = -C(\varphi, \xi) \frac{\partial V_{\text{eff}}(\varphi, \xi)}{\partial \varphi}$  the Nielsen-Fukuda-Kugo (NFK) identity

$$V_0 \left( 0, v_{S, \text{tree}}^{\text{sym}} \right) + V_1 \left( 0, v_{S, \text{tree}}^{\text{sym}} ; T \right) = V_0 \left( v_{\text{tree}}, v_{S, \text{tree}} \right) + V_1 \left( v_{\text{tree}}, v_{S, \text{tree}} ; T \right)$$

$v_C, v_{SC}$  and  $v_{SC}^{\text{sym}}$  are determined by the use of  $V^{HT}$

$$V_{\text{eff}}(\varphi, \varphi_S; T) = V_0(\varphi, \varphi_S; T) + \sum_i n_i \left[ V_{\text{CW}} \left( \bar{m}_i^2 \right) + \frac{T^4}{2\pi^2} I_{B,F} \left( \frac{\bar{m}_i^2}{T^2} \right) \right]$$

$$V_{\text{CW}} \left( \bar{m}_i^2 \right) = \frac{\bar{m}_i^4}{64\pi^2} \left( \ln \frac{\bar{m}_i^2}{\bar{\mu}^2} - c_i \right), \quad I_{B,F} \left( a^2 \right) = \int_0^\infty dx x^2 \ln \left( 1 \mp e^{-\sqrt{x^2 + a^2}} \right)$$

**Parwani scheme** Replace  $\bar{m}^2$  with thermally corrected field depending masses  $\bar{M}^2$

**AE scheme**  $V_{\text{daisy}}(\varphi, \varphi_S; T) = \sum_{\substack{i=h_1,2,\chi \\ W_L, Z_L, \gamma_L}} -n_i \frac{T}{12\pi} \left[ (\bar{M}_i^2)^{3/2} - (\bar{m}_i^2)^{3/2} \right]$

# EWPT in the degenerate-scalar scenario

	Gauge independence	Renormalization, so that tree-level relationships are also established at the one- loop level	One loop contribution
HT potential	○		✗
PRM scheme	○	✗	○
Parwani scheme	✗	○	○
AE scheme	✗	○	○

# EWPT in the degenerate-scalar scenario

$$V^{\text{HT}}(z, \gamma; T) = c_0 + c_1 z + (c_2 + c'_2 T^2) z^2 - c_3 z^3 + c_4 z^4$$

$$c_0 = \sqrt{2} a_1 v_s^A(T) + \frac{1}{4} (b_1 + b_2 + 2\Sigma_S T^2) (v_s^A(T))^2 + \frac{1}{16} (v_s^A(T))^4,$$

$$c_1 = \left( \sqrt{2} a_1 + \frac{1}{2} (b_1 + b_2 + 2\Sigma_S T^2) v_s^A(T) + \frac{1}{4} d_4 (v_s^A(T))^3 \right) \sin \gamma,$$

$$c_2 = \frac{1}{4} ((b_1 + b_2) \sin^2 \gamma + m^2 \cos^2 \gamma) + \frac{1}{8} (3d_2 \sin^2 \gamma + \delta_2 \cos^2 \gamma) (v_s^A(T))^2,$$

$$c'_2 = \frac{1}{2} (\Sigma_H \cos^2 \gamma + \Sigma_S \sin^2 \gamma),$$

$$c_3 = \frac{1}{4} \sin \gamma (d_2 \sin^2 \gamma + \delta_2 \cos^2 \gamma) v_s^A(T),$$

$$c_4 = \frac{1}{16} (d_2 \sin^4 \gamma + 2\delta_2 \sin^2 \gamma \cos^2 \gamma + \lambda \cos^4 \gamma),$$

$$T_C^2 = \frac{1}{2(\Sigma_H + \Sigma_S t_{\gamma_C}^2)} \left[ -m^2 - \frac{(v_{SC}^{\text{sym}})^2 \delta_2}{2} - \left\{ b_1 + b_2 + \left( \frac{3d_2}{2} - \frac{(\delta_2 + d_2 t_{\gamma_C}^2)^2}{\lambda + 2\delta_2 t_{\gamma_C}^2 + d_2 t_{\gamma_C}^4} \right) (v_{SC}^{\text{sym}})^2 \right\} t_{\gamma_C}^2 \right],$$

$$v_C = \frac{-2t_{\gamma_C} (v_{SC}^{\text{sym}})^2 (\delta_2 + d_2 t_{\gamma_C}^2)}{\lambda + 2\delta_2 t_{\gamma_C}^2 + d_2 t_{\gamma_C}^4}$$

$$\begin{aligned} t_{\gamma_C} &= \frac{\sin \gamma(T_C)}{\cos \gamma(T_C)} \\ &= \frac{v_{SC} - v_{SC}^{\text{sym}}}{v_C}, \\ v_C &= \lim_{T \nearrow T_C} v(T), \\ v_{SC} &= \lim_{T \nearrow T_C} v_S(T), \\ v_{SC}^{\text{sym}} &= \lim_{T \searrow T_C} v_S(T) \end{aligned}$$

# EWPT in the degenerate-scalar scenario

$$\delta_2 = \frac{2}{v v_S} (m_{h_1}^2 - m_{h_2}^2) \sin \alpha \cos \alpha$$

Invariant under the transformation  $m_{h_1}^2 - m_{h_2}^2 \rightarrow - (m_{h_1}^2 - m_{h_2}^2)$  and  $\alpha \rightarrow -\alpha$

$$d_2 = \frac{2}{v_S^2} \left[ m_{h_1}^2 + (m_{h_2}^2 - m_{h_1}^2) \cos^2 \alpha + \frac{\sqrt{2} a_1}{v_S} \right] \simeq \frac{2}{v_S^2} \left[ m_{h_1}^2 + \frac{\sqrt{2} a_1}{v_S} \right]$$

The sign of  $m_{h_1}^2 - m_{h_2}^2$  cannot be compensated by that of  $\alpha$

# EWPT in the degenerate-scalar scenario

The energy difference between the electroweak vacuum prescribed by  $(v, v_S)$  and the local vacuum on the  $\varphi_S$  axis specified by  $(0, v_S^{\text{sym}})$

$$\begin{aligned}\Delta E = & V_0(0, v_S^{\text{sym}}) - V_0(v, v_S) \\ = & \sqrt{2}a_1(v_S^{\text{sym}} - v_S) + \frac{1}{4}(b_1 + b_2)\left((v_S^{\text{sym}})^2 - v_S^2\right) + \frac{d_2}{16}\left((v_S^{\text{sym}})^4 - v_S^4\right) \\ & - \frac{m^2}{4}v^2 - \frac{\lambda}{16}v^4 - \frac{\delta_2}{8}v^2v_S^2\end{aligned}$$

$\Delta E$  could be negative for  $\delta_2 \gg 1$  and  $d_2 \ll 1$ .



$\delta_2$  and  $d_2$  have the upper and lower bound respectively.

Other conditions

Bounded from below

$$\lambda > 0, d_2 > 0, \lambda d_2 > \delta_2^2$$

Vacuum stability

$$\lambda \left( d_2 + \frac{2\sqrt{2}|a_1|}{v_S^3} \right) > \delta_2^2$$

Conditions from  
Perturbation Theory

$$\lambda \leq \frac{16}{3}\pi, d_2 \leq \frac{16}{3}\pi$$

# EWPT in the degenerate-scalar scenario

Phys. Rev. D 93, 065032 (2016)

Local minimum  $(v_h, v_S, v_A)$

→ It might be local min. also in  $S = v_S, h = v_h$  subspace

When the coeff. of  $A^2$  is negative,  $V_0(v_h, v_S, A)$  has min.

$$\frac{b_2 + b_1}{2} + \frac{\sqrt{2}c_2}{6}v_s + \frac{d_2}{4}v_s^2 + \frac{\delta_2}{4}v_h^2 < 0 \quad \boxed{m_A^2 + \frac{\sqrt{2}c_2}{6}(x - v_s) + \frac{d_2}{4}(v_s^2 - x^2) + \frac{\delta_2}{4}(v_h^2 - v^2) < 0}$$

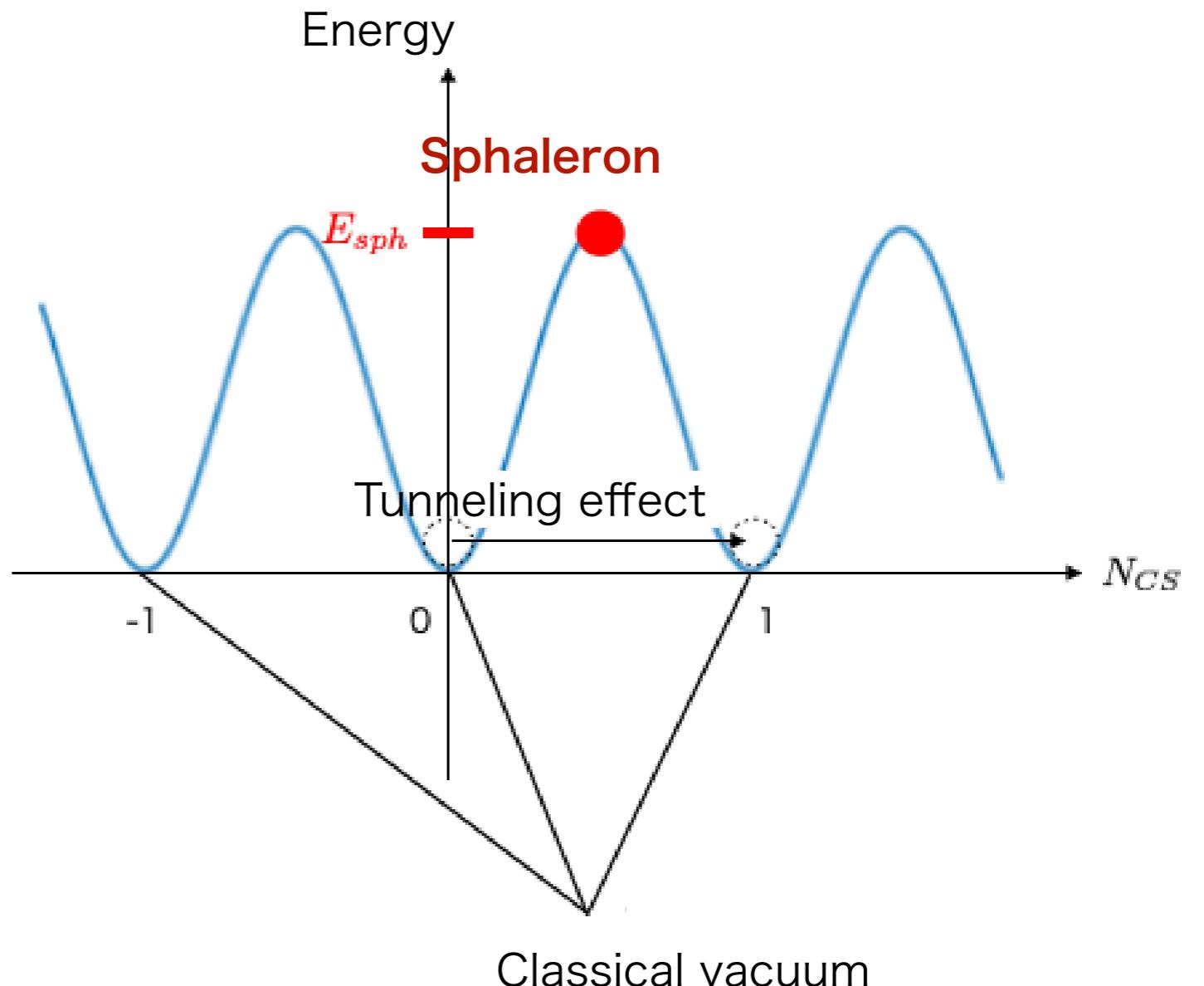
In this study, this inequality does not hold.

In  $T \neq 0$ , it is stable at  $A=0$  due to thermal contribution

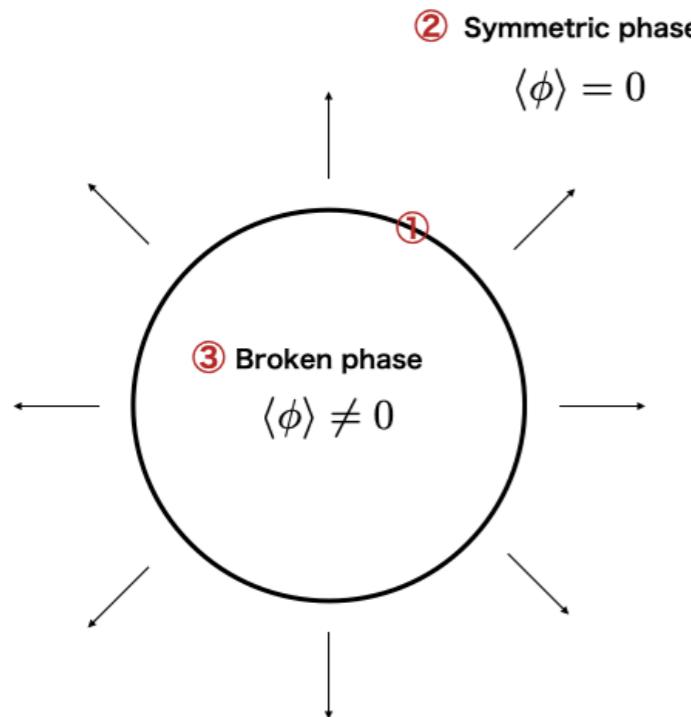
# Electroweak baryogenesis

Baryon number violation  
→ Sphaleron process

**Baryon number**  
quark : 1/3  
antiquark : -1/3  
lepton : 0  
boson : 0



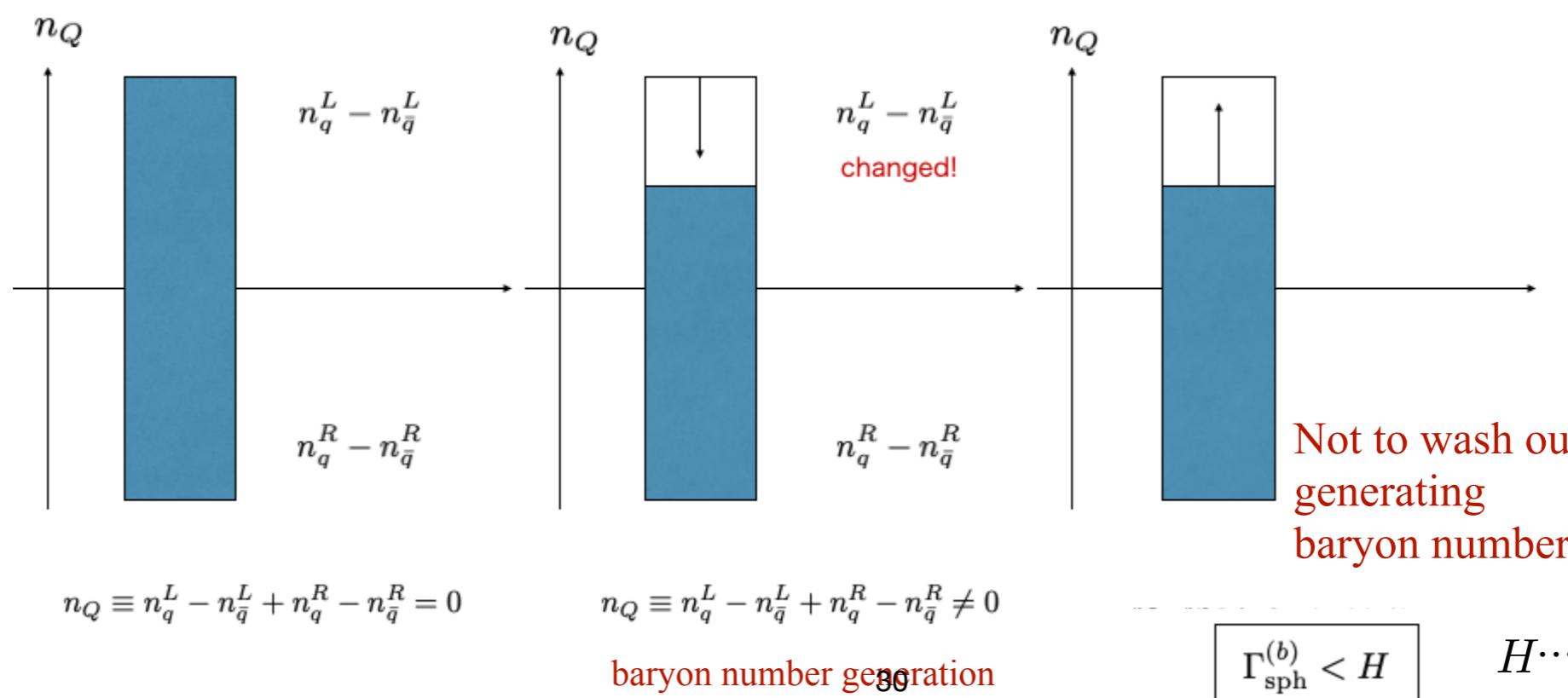
# Electroweak baryogenesis



Transmittance, Reflectance

Left-handed quark  $q^L$  = Right-handed antiquark  $\bar{q}^R$

Left-handed antiquark  $\bar{q}^L$  = Right-handed quark  $q^R$



# Electroweak baryogenesis

The change rate in the baryon number in the broken phase  $\Gamma_B^{(b)}(T)$

To generate baryon number

$$\Gamma_B^{(b)}(T) \quad \text{must be small}$$

$$\Gamma_B^{(b)}(T) \simeq (\text{pre}) \frac{\Gamma_{\text{sph}}^{(b)}}{T^3} \simeq (\text{pre}) e^{-E_{\text{sph}}/T}$$

$E_{\text{sph}}$  ..... sphaleron energy

Sphaleron rate/time/volume

$$\Gamma_{\text{sph}}^{(b)} \simeq T^4 e^{-E_{\text{sph}}/T}$$

$$E_{\text{sph}} \propto v(T)$$

Higgs vev must be large



$$\frac{v_c}{T_c} \gtrsim 1$$

# Electroweak baryogenesis

$$\Gamma_B^{(b)}(T) < H$$

$$\rightarrow \quad \Gamma_B^{(b)}(T) \simeq (\text{ pre }) e^{-E_{\text{sph}}/T} < H(T) \simeq 1.66 \sqrt{g_*} T^2 / m_P$$

$g_*$  ..... massless dof

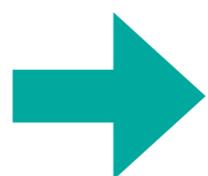
$m_P$ .....Plank mass

$$E_{\text{sph}} = 4\pi v \mathcal{E} / g_2 \rightarrow \quad g_2 \text{ ..... SU(2) gauge coupling constant}$$

$$\frac{v}{T} \geq \frac{g_2}{4\pi \mathcal{E}} (42.97 + \text{log corrections})$$

In the case of the SM

$$m_h = 125 \text{ GeV}, \mathcal{E} = 1.92(T = 0)$$



$$\frac{v}{T} \geq 1.16$$

# EWPT in the SM

M. Quiros, [arXiv:hep-ph/9901312 [hep-ph]]

Effective potential of the SM

$$\Gamma[\phi_c] = - \int d^4x V_{\text{eff}}(\phi_c)$$

- tree level potential
- zero-temperature one loop potential  
(the Coleman Weinberg Potential)
- finite-temperature one loop potential

$$V(\phi_c, T) = D(T^2 - T_o^2) \phi_c^2 - ET\phi_c^3 + \frac{\lambda(T)}{4}\phi_c^4$$

$$D = \frac{2m_W^2 + m_Z^2 + 2m_t^2}{8v^2}$$

$$E = \frac{2m_W^3 + m_Z^3}{4\pi v^3}$$

$$T_o^2 = \frac{m_h^2 - 8Bv^2}{4D}$$

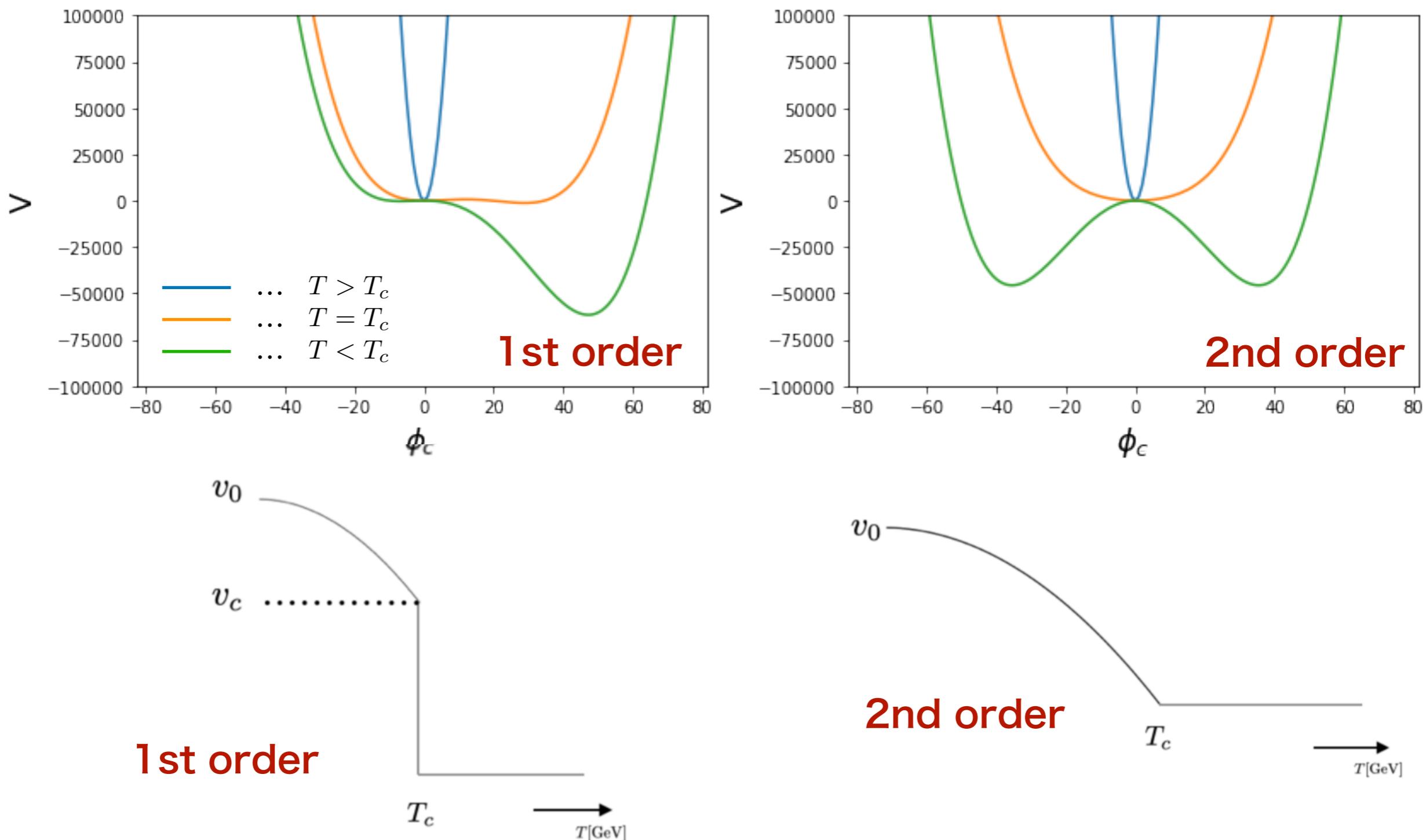
$$B = \frac{3}{64\pi^2 v^4} (2m_W^4 + m_Z^4 - 4m_t^4)$$

$$\lambda(T) = \lambda - \frac{3}{16\pi^2 v^4} \left( 2m_W^4 \log \frac{m_W^2}{A_B T^2} + m_Z^4 \log \frac{m_Z^2}{A_B T^2} - 4m_t^4 \log \frac{m_t^2}{A_F T^2} \right)$$

Higgs field

$$H = \begin{pmatrix} \chi_1 + i\chi_2 \\ \frac{\phi_c + h + i\chi_3}{\sqrt{2}} \end{pmatrix} \quad \begin{array}{l} \phi_c \dots\dots \text{real background field} \\ \chi_a \ (a = 1, 2, 3) \dots\dots \text{goldstone bosons} \end{array}$$

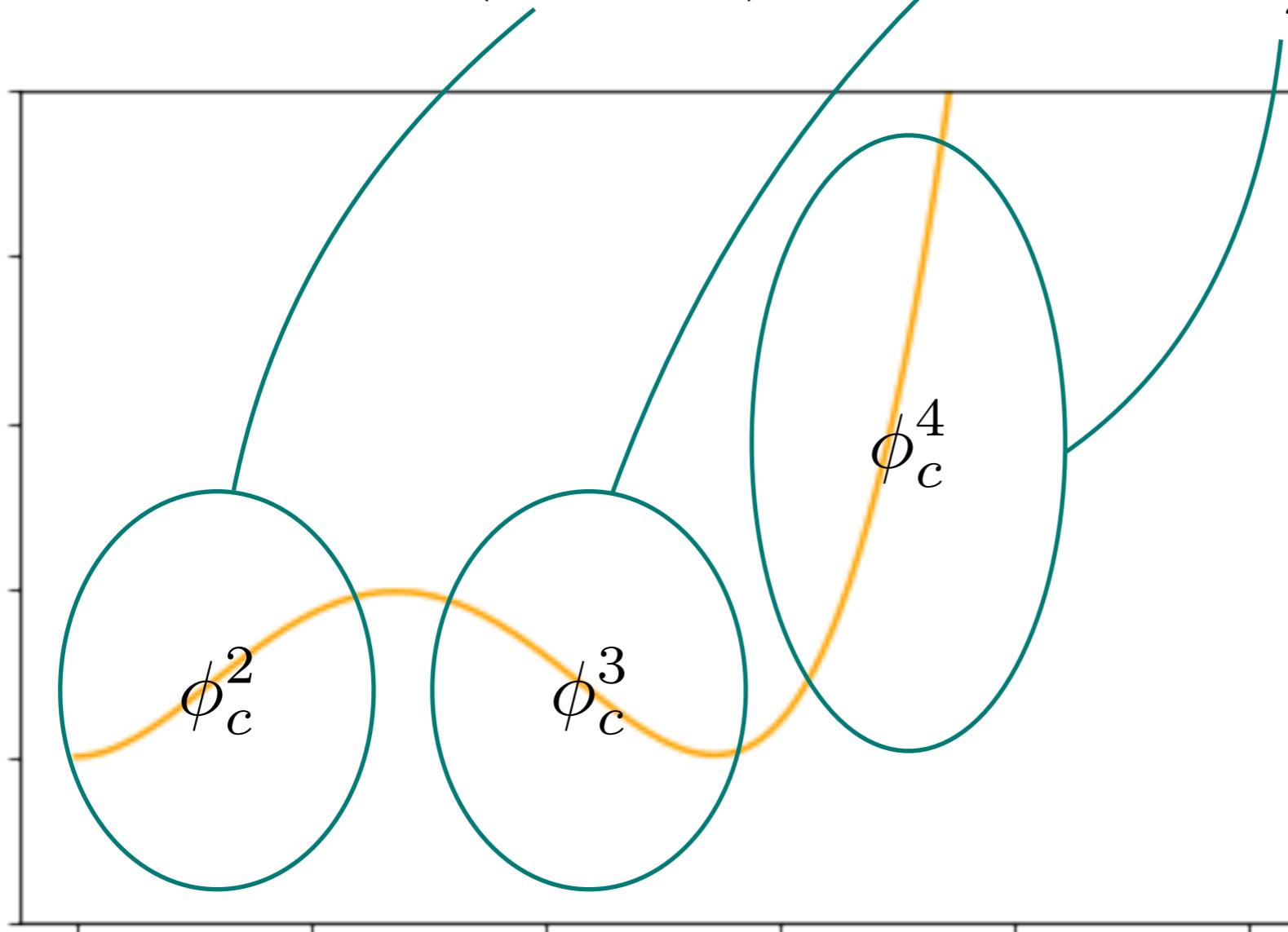
# EWPT in the SM



$-ET\phi_c^3$  from finite-temperature boson loop causes a 1st order PT.  
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# EWPT in the SM

$$V(\phi_c, T) = D(T^2 - T_o^2) \phi_c^2 - ET\phi_c^3 + \frac{\lambda(T)}{4} \phi_c^4$$



$v(T)$  makes  
discontinuous  
transition.  
(1st order PT)



A barrier is needed  
between the origin,  
and  $v(T)$



$\phi_c^3$  contributes.

# EWPT in the SM

In the SM, SFOEWPT condition

$$\frac{v_c}{T_c} = \frac{2E}{\lambda(T_c)} \gtrsim 1$$



$$m_h \lesssim 64 \text{ GeV}$$

**Conflict with observation at LHC → We need to extend the SM!**

# Numerical results

We use a public code micrOMEGAs to calculate  $\Omega_\chi h^2$  and  $\sigma_{\text{SI}}$ .

The value of  $\Omega_\chi h^2$  should not exceed the observed value

$$\Omega_{\text{DM}} h^2 = 0.1200 \pm 0.0012$$

In the case of  $m_\chi = 30$  GeV, for instance, the maximum value is  $\sigma_{\text{SI}} \simeq 4.1 \times 10^{-47}$  cm<sup>2</sup> under the assumption  $\Omega_\chi = \Omega_{\text{DM}}$ .

In cases that  $\Omega_\chi < \Omega_{\text{DM}}$ , we scale  $\sigma_{\text{SI}}$  as

$$\tilde{\sigma}_{\text{SI}} = \left( \frac{\Omega_\chi}{\Omega_{\text{DM}}} \right) \sigma_{\text{SI}}$$

# Future work

**Main topic: About the feasibility of CxSM when CP symmetry is broken.**

## 1. Spontaneous CP violation

$$V_0 = \frac{m^2}{2}|H|^2 + \frac{\lambda}{4}|H|^4 + \frac{\delta_2}{2}|H|^2|S|^2 + \frac{b_2}{2}|S|^2 + \frac{d_2}{4}|S|^4 + \left( \textcircled{a_1} S + \textcircled{\frac{b_1}{4}} S^2 + \text{c.c.} \right)$$

Investigate the feasibility of SFOEWPT

Introduce complex phase

## 2. Explicit CP violation

Introduce such a dimension-five operator

(coeff.)  $\bar{t}_L \gamma_5 t_R S + h.c.$

There is a phase in the (coeff.) that cannot be removed by the field redefinition, and it contributes to the baryon number generation.