

Self-Resonant Dark Matter

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Chung-Ang University, SeongSik Kim, Bin Zhu, and Hyun Min Lee arXiv 2108.06278

Motivation - SIDM model

There are some mismatches between Λ CDM prediction, and observations.

Self-Interacting Dark Matter (SIDM) model could resolve these mismatches with an appropriate DM scattering cross-section.



Core Cusp Problem

Reprinted From S. Tulin, Hai-Bo Yu (2017), arXiv:1705.02358v2 [hep-ph]



Missing Satellites Problem

Reprinted From A. A. Klypin, A. V. Kravtsov, O. Valenzuela, and F. Prada, Astrophys. J. 522, 82 (1999), astro-ph/9901240.

Appropriated DM Self–Scattering Cross–section is $\sigma/m \sim O(1)$ cm²/g

Sommerfeld Enhancement

Sommerfeld Enhancement : One of the cross-section enhancing effects

When?

Long-Range Attractive Interaction Between Colliding Species

Non-relativistic collision

How?

By Enhancing 2 Particle-wavefunction Amplitude Near the Origin

Why?

The Enhancement is so Sizable, affecting DM relic abundance significantly.



$$\sigma = |A|^2 \sigma_0$$

σ₀: cross-section calculated with ordinary approach
σ: Corrected Cross-section
A: Sommerfeld Factor (Enhanced Amplitude ratio)

Bethe-Salpeter Equation

Connect Enhancement Condition to QFT Condition

A Common Situation for $2\rightarrow 2$ process



High-Order Term suppressed by Perturbative Coupling

Momentum in denominator ignored since we consider a non-relativistic limit

Bethe-Salpeter Equation

Connect Enhancement Condition to QFT Condition

Light mediator breaks perturbativity!



High-Order Term Enhanced!

Even momentum is included in the denominator there is still a possibility of breaking perturbativity

Bethe-Salpeter Equation

Methods for calculating Sommerfeld Factor



- $G_i(q)$ (i = 1,2) : particle *i* propagator with momentum transfer q. $\frac{i}{q^2 - m_i^2}$,
- $i\tilde{\Gamma}$: 4 point function via single mediator exchange. $q = p_{f2} - p_{i1}$

· $i\Gamma$: total 4 point function

$$i\Gamma(p_{i1}, p_{i2}; p_{f1}, p_{f2}) = -\int \frac{d^4k}{(2\pi)^4} \tilde{\Gamma}(p_{i1}, p_{i2}; k, p_{i1} + p_{i2} - k)G_2(k)G_1(p_{i1} + p_{i2} - k)\Gamma(k, p_{i1} + p_{i2} - k; p_{f1}, p_{f2})$$

The Final Equation is the Schrodinger Equation with Yukawa potential

(scalar particle model)

$$\left(\frac{\overrightarrow{p}^2}{2\mu} - E\right)\widetilde{\chi}(Q) + \int \frac{d^3q'}{(2\pi)^3} V(\overrightarrow{p} - \overrightarrow{q}')\widetilde{\chi}(q') = 0$$

New Way to Induce Non-Perturbativity

Collision Mediated by MASSIVE Own



denominator =
$$q^2 - m_{med}^2 = (E_{f2} - E_{i1})^2 - (\overrightarrow{p}_{f2} - \overrightarrow{p}_{i1})^2 - m_1^2$$

Simple Model for Precise Calculation

Collision Mediated by MASSIVE Own

$$\mathscr{L} = |\partial_{\mu}\phi_{1}|^{2} - m_{1}^{2}\phi_{1}^{2} + \frac{1}{2}(\partial_{\mu}\phi_{2})^{2} - \frac{1}{2}m_{2}^{2}\phi_{2}^{2} - 2gm_{1}\phi_{2}|\phi_{1}|^{2}$$

 m_1 in coupling introduced to make g dimensionless.

2 DM (ϕ_1 , ϕ_2) model, One Real Scalar and One Complex Scalar

We consider $(m_1 <) m_2 < 2m_1$ case.

Q. Why $m_2 < 2m_1$?



Significant Decay Allowed in $m_2 \ge 2m_1$ model, Leading No Particles to interact

Equation From New Mechanism

The Same Bethe-Salpeter Equation applied for this mechanism

With Different equation outcome



Schrodinger-like Equation with Long-Range potential-like term

Yukawa limit Calculation

Numerical Sommerfeld Factor for s-wave cases



FIG. 2: Sommerfeld factor for s-wave elastic scattering, $\phi_1\phi_2 \rightarrow \phi_1\phi_2$, as a function of the relative velocity $v_{\rm rel}$. We chose $\alpha = g^2/(4\pi) = 0.1$ and $\Delta = 10^{-2}, 10^{-3}, 10^{-4}, 10^{-5}$, for $m_2 = 2m_1(1 - \Delta)$, in the lines from bottom to top.

$$\Delta \equiv 1 - \frac{m_2}{2m_1} \quad \text{Mass-difference parameter}$$

We call Yukawa limit because exponential suppression of potential remains.

- Enhancement becomes larger as $|m_2 2m_1|$ smaller.
- There is no absolute mass dependence in Enhancement.

Yukawa limit Result

Sommerfeld-Enhanced cross-section. s-wave cases



FIG. 3: Self-scattering cross section per dark matter mass for s-wave elastic scattering, $\phi_1\phi_2 \rightarrow \phi_1\phi_2$, as a function of $\langle v_{\rm rel} \rangle$. We chose $\Delta = 10^{-5}, 2 \times 10^{-5}$ and $m_1 = 7, 3 \,\text{GeV}$ in orange solid and dashed lines, respectively. We took $\alpha = g^2/(4\pi) = 0.1$ and $m_{\rm eff} = 2m_1(1+b)$.

M. Kaplinghat, S. Tulin and H. B. Yu, (2016), Phys. Rev. Lett. 116 no.4, 041302, arXiv:1508.03339

$$\Delta \equiv 1 - \frac{m_2}{2m_1}$$

Experimental Observations

- Red : THINGS dwarf galaxies
- Green : clusters
- Blue : LSB galaxies

Huge DM Mass requires a smaller Delta for larger non-perturbativity.

Summary

- Dark Matter Scattering Cross-sections play important role in comparing observation and theory.
- Sommerfeld Effect, which is traditionally considered due to light mediator, enhances cross-section, especially lower velocity.
- We present a new Sommerfeld enhancement mechanism without a light mediator. Instead, the particle itself became a propagator.
- We Calculated Sommerfeld Factor and Enhanced Cross-section with a simple model.