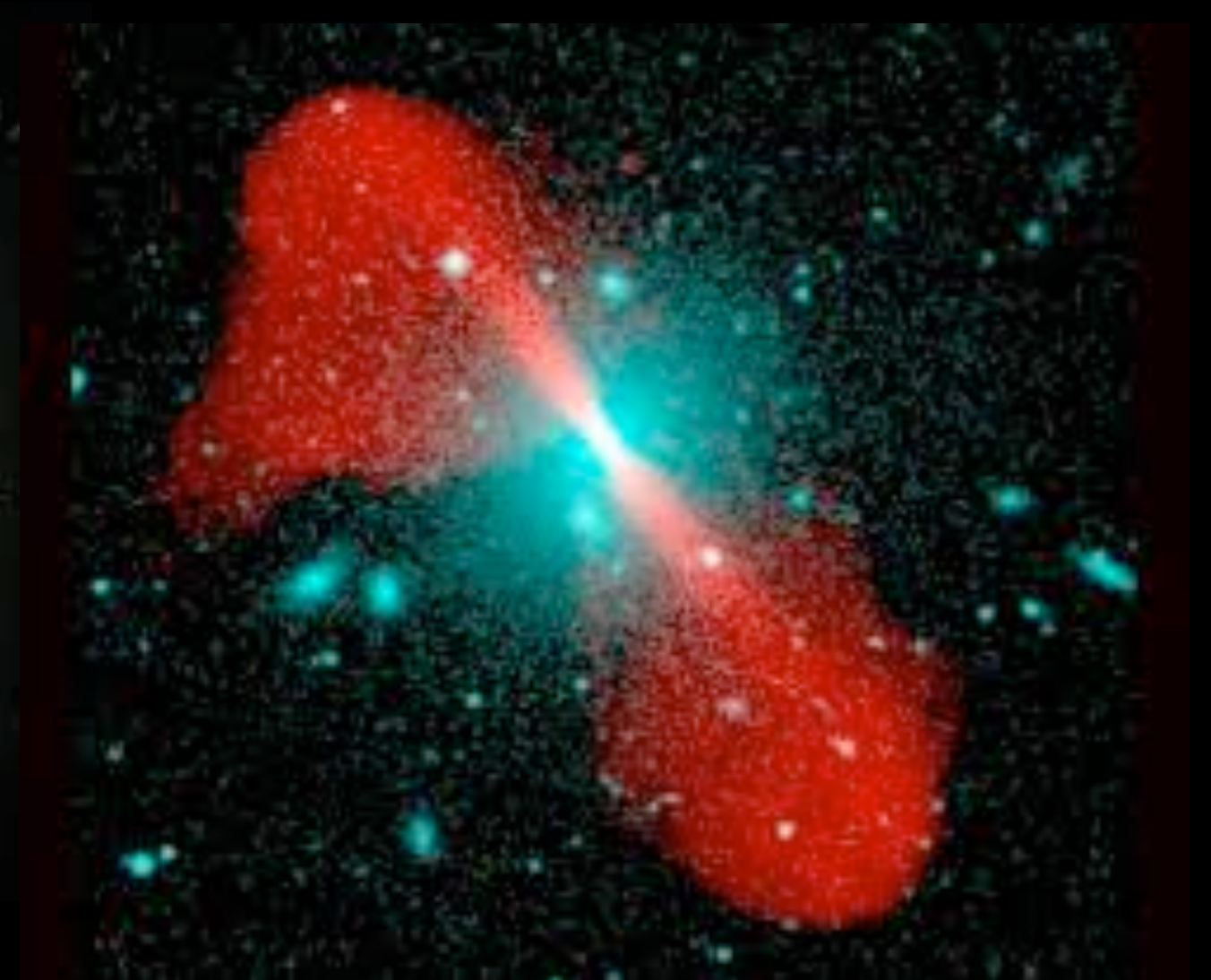




Galaxy formation



But, Galaxies do change!

Classic observational cosmology

Ann. Rev. Astron. Astrophys. 1988, 26: 561-630
 Copyright © 1988 by Annual Reviews Inc. All rights reserved

$$\frac{kc^2}{R^2} = H_0^2(2q_0 - 1) \equiv H_0^2(\Omega_0 - 1)$$

$$dl^2 = R^2(t) \left[\frac{dr^2}{1-kr^2} + r^2(d\theta^2 + \sin^2 \theta d\phi^2) \right]$$

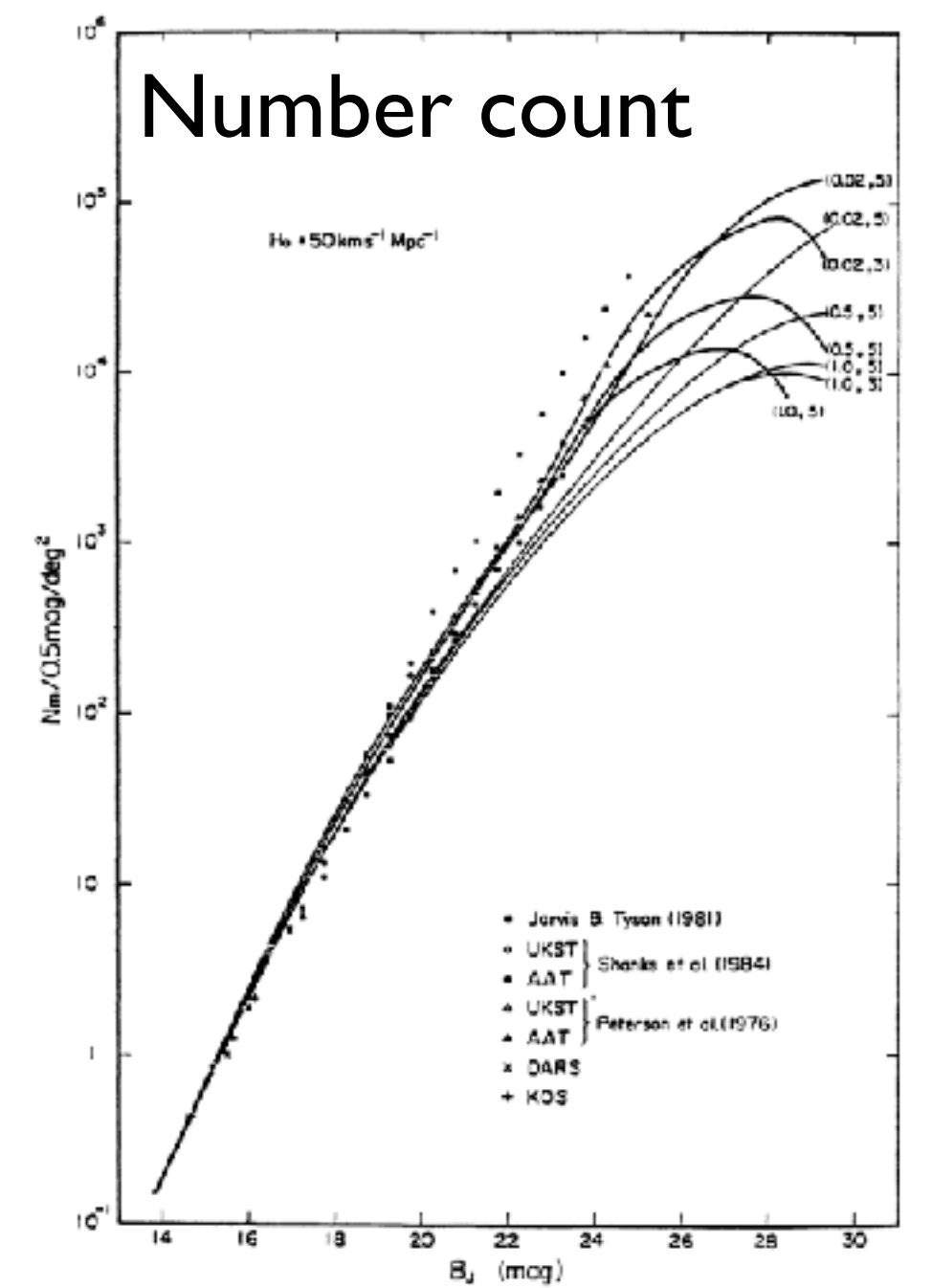


Figure 3 Comparison of predicted $A(m, q_0)$ functions with the observed differential counts (i.e. number at magnitude B_j in magnitude interval ± 0.25 mag) from various surveys. The four heavy lines to the left are for the marked values of q_0 and the redshift of galaxy formation, with luminosity evolution included. The four light lines to the right are the same but with zero luminosity evolution. The relations depend on H_0 only to set the time scale for the above luminosity evolution correction (from Mould & Tully 1982)

OBSERVATIONAL TESTS OF WORLD MODELS

Allan Sandage

Department of Physics and Astronomy, The Johns Hopkins University, Baltimore, Maryland 21218 and Space Telescope Science Institute, 3700 San Martin Drive, Baltimore, Maryland 21218¹

1. THE ELEMENTS OF PRACTICAL COSMOLOGY

The standard model of cosmology, based on what has come to be called the Friedmann-Lemaitre-Robertson-Walker (FLRW) model (hereinafter simply the Friedmann model), is now part of scientific culture. The most popular current version leads to the hot big bang (HBB) description of events near the beginning of the cosmic expansion, which has often been called a creation² moment at the beginning of physical time. In this review a prejudice in favor of the HRR lin contrast to cold beginnings discussed

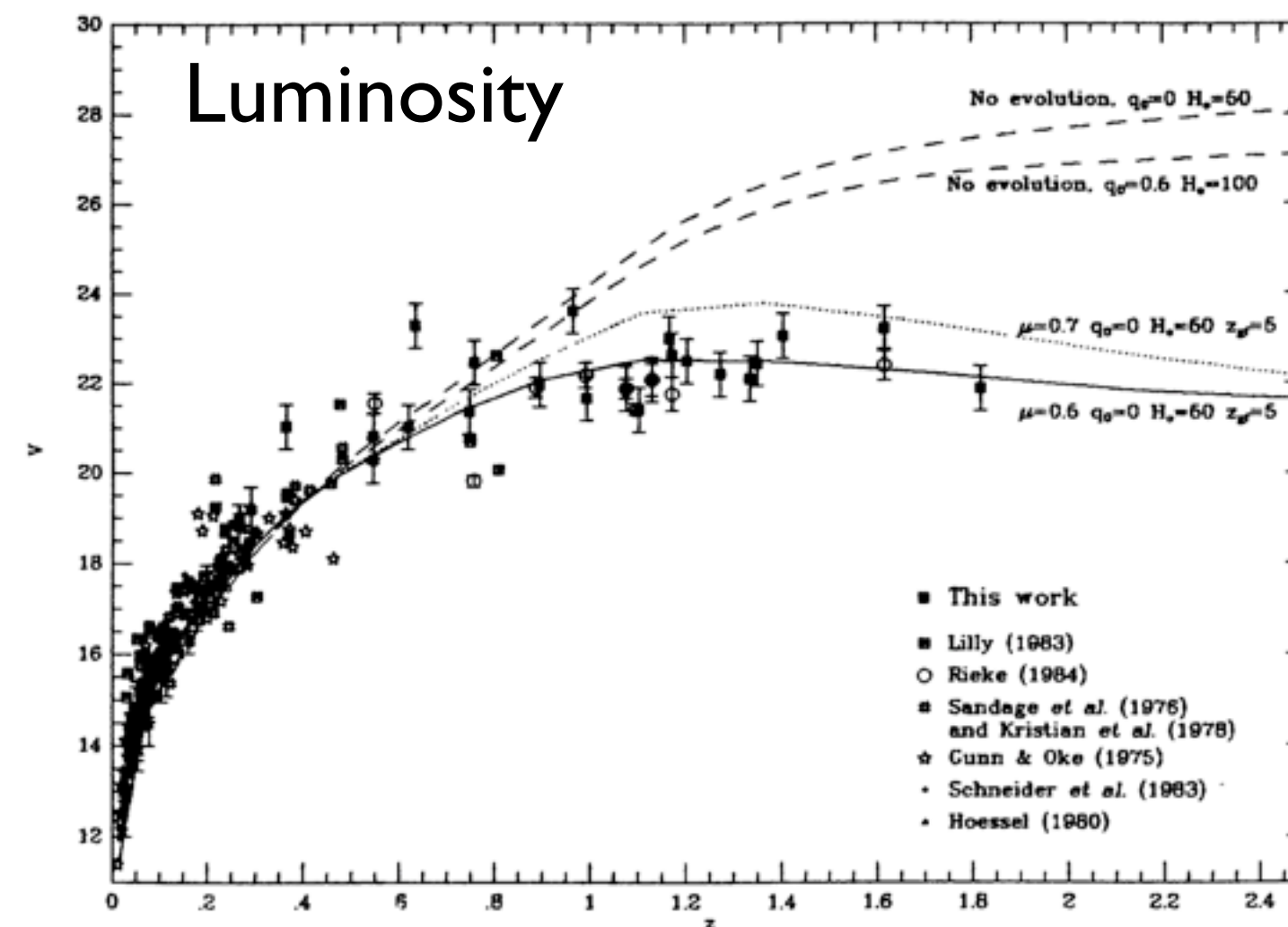


Figure 7 The $m(z)$ Hubble diagram (plotted differently from Figure 5, with V magnitude as ordinate and linearly with z) extended to large redshift. Predicted theoretical $m(z)$ lines for various q_0 values and evolutionary corrections (Bruzual models) are shown. The K corrections have been applied to the theoretical curves (from Djorgovski et al. 1985).

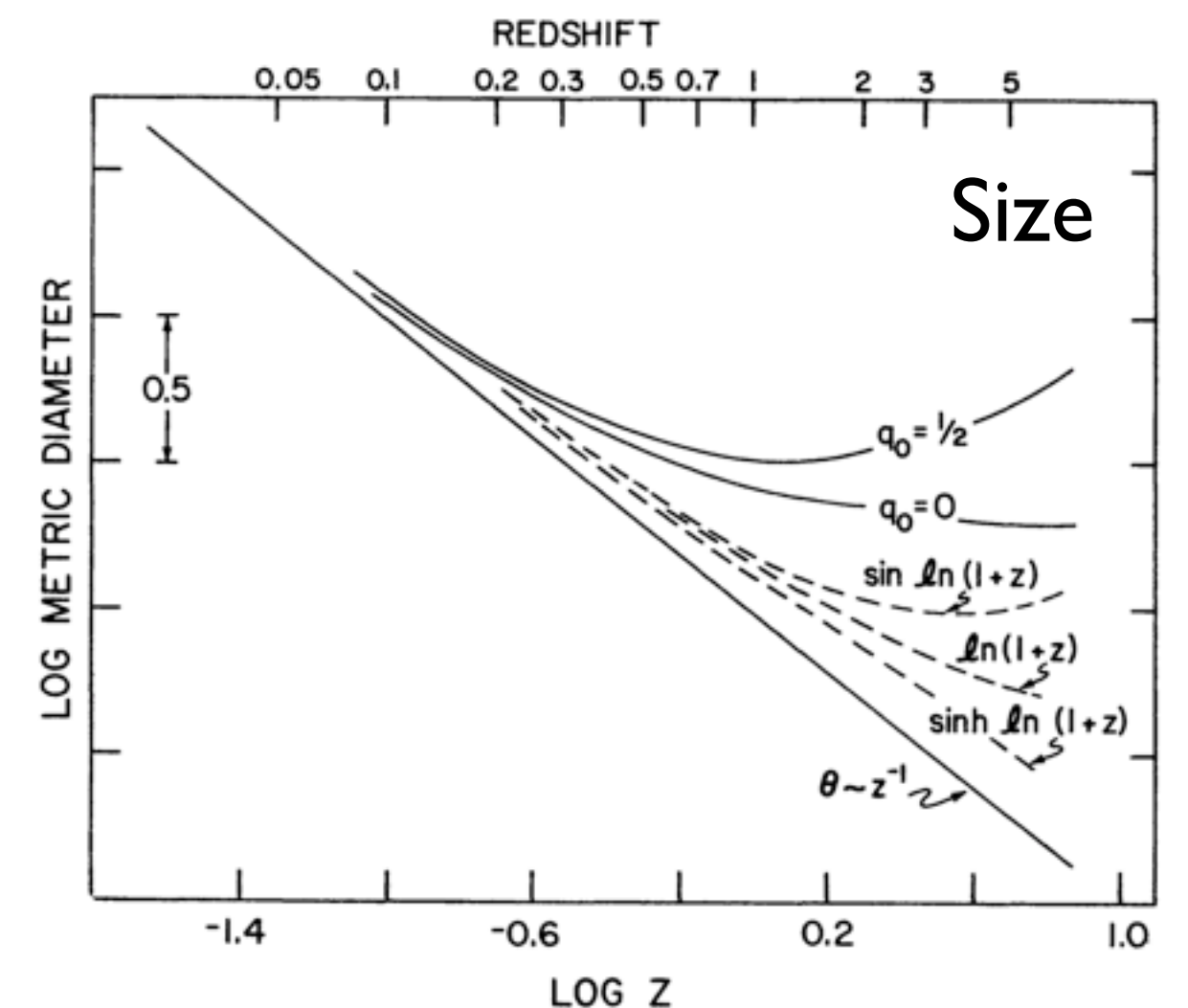


Figure 11 Theoretical angular size-redshift relations for $q_0 = 0$ and $q_0 = 1/2$ standard Friedmann models, and for three types of geometry for the tired-light speculation. The Euclidean $\theta \sim z^{-1}$ intuitive dependence is shown for purposes of comparison.

Galaxy challenge

Disc galaxies



Spiral Galaxy M101  HUBBLESITE.org



25000 ly



NGC 891

Credit: Hewholooks <https://commons.wikimedia.org/wiki/File:NGC891HunterWilson.jpg>

Galaxy challenge

Elliptical galaxies

$M_{\text{star}} \sim 1e12$, $M_{\text{BH}} \sim 6.5e9 M_{\text{sun}}$

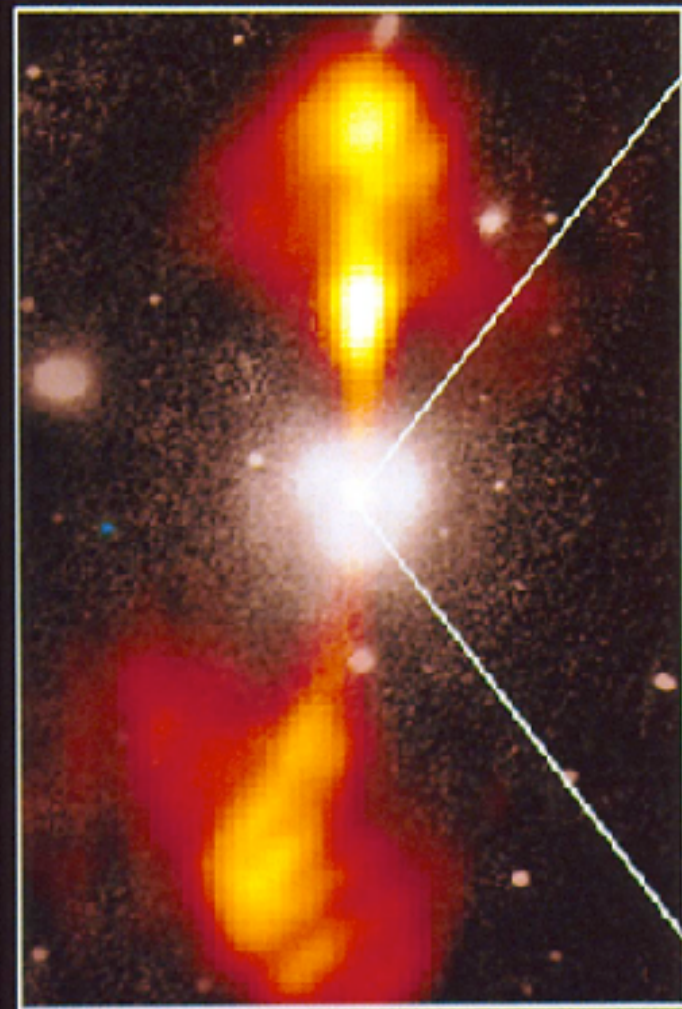


NASA/JPL/IPAC

Core of Galaxy NGC 4261

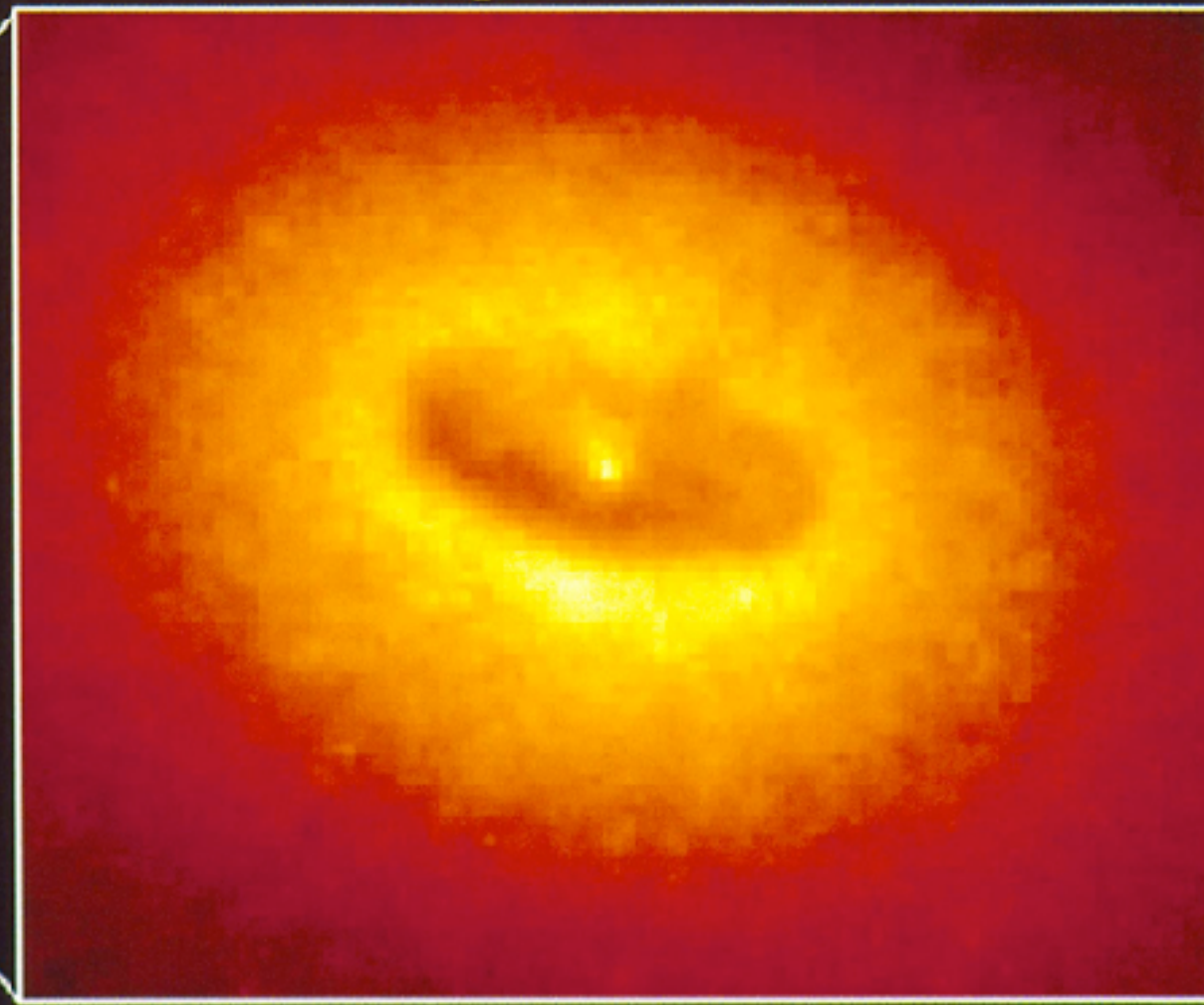
Hubble Space Telescope
Wide Field / Planetary Camera

Ground-Based Optical/Radio Image



380 Arc Seconds
88,000 LIGHT-YEARS

HST Image of a Gas and Dust Disk



17 Arc Seconds
400 LIGHT-YEARS

April 5



April 6



April 10



April 11



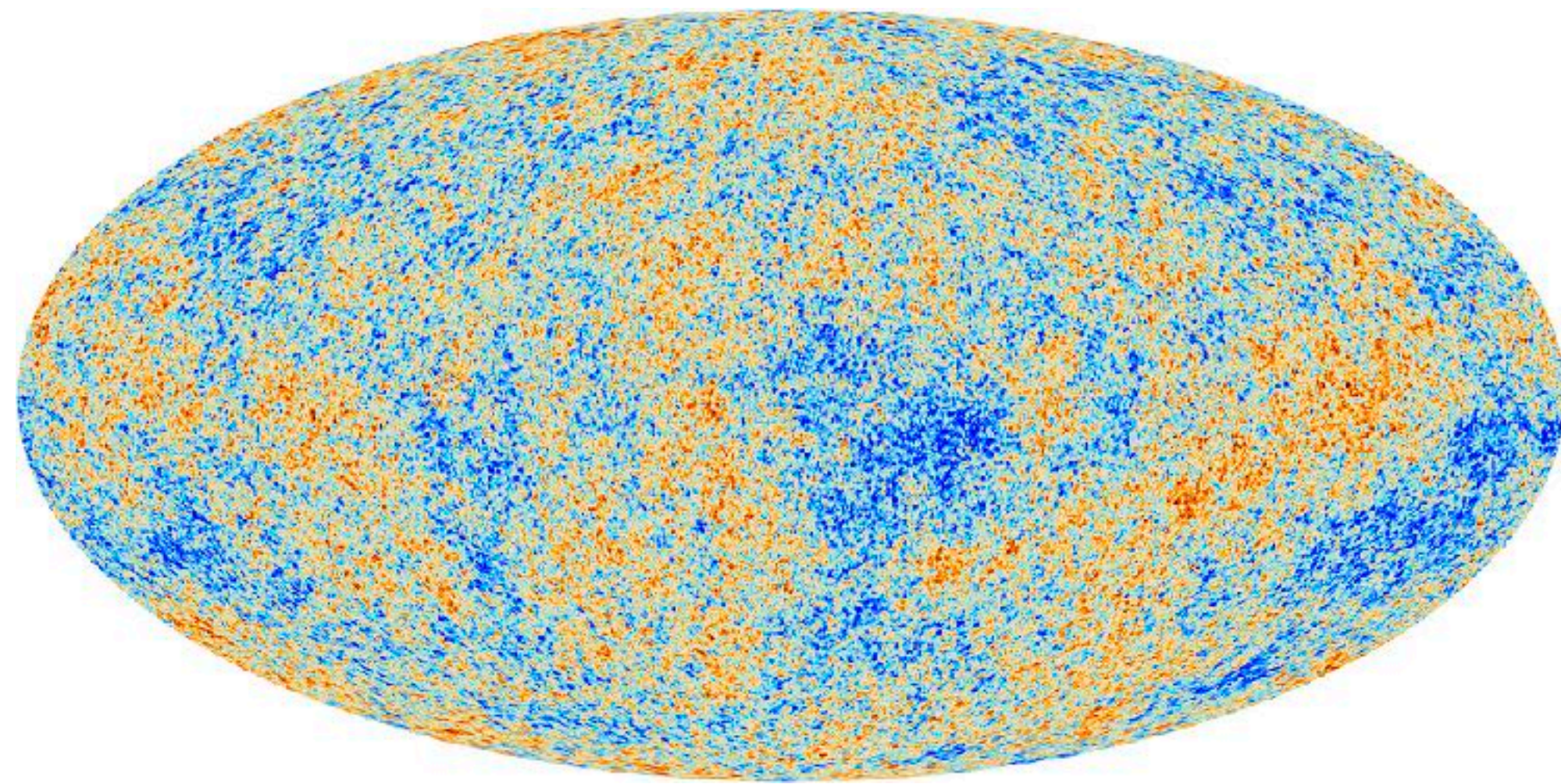
Observations from the Event Horizon Telescope of the supermassive black hole at the center of the elliptical galaxy M87, for four different days. [EHT Collaboration et al 2019]

Understanding galaxies

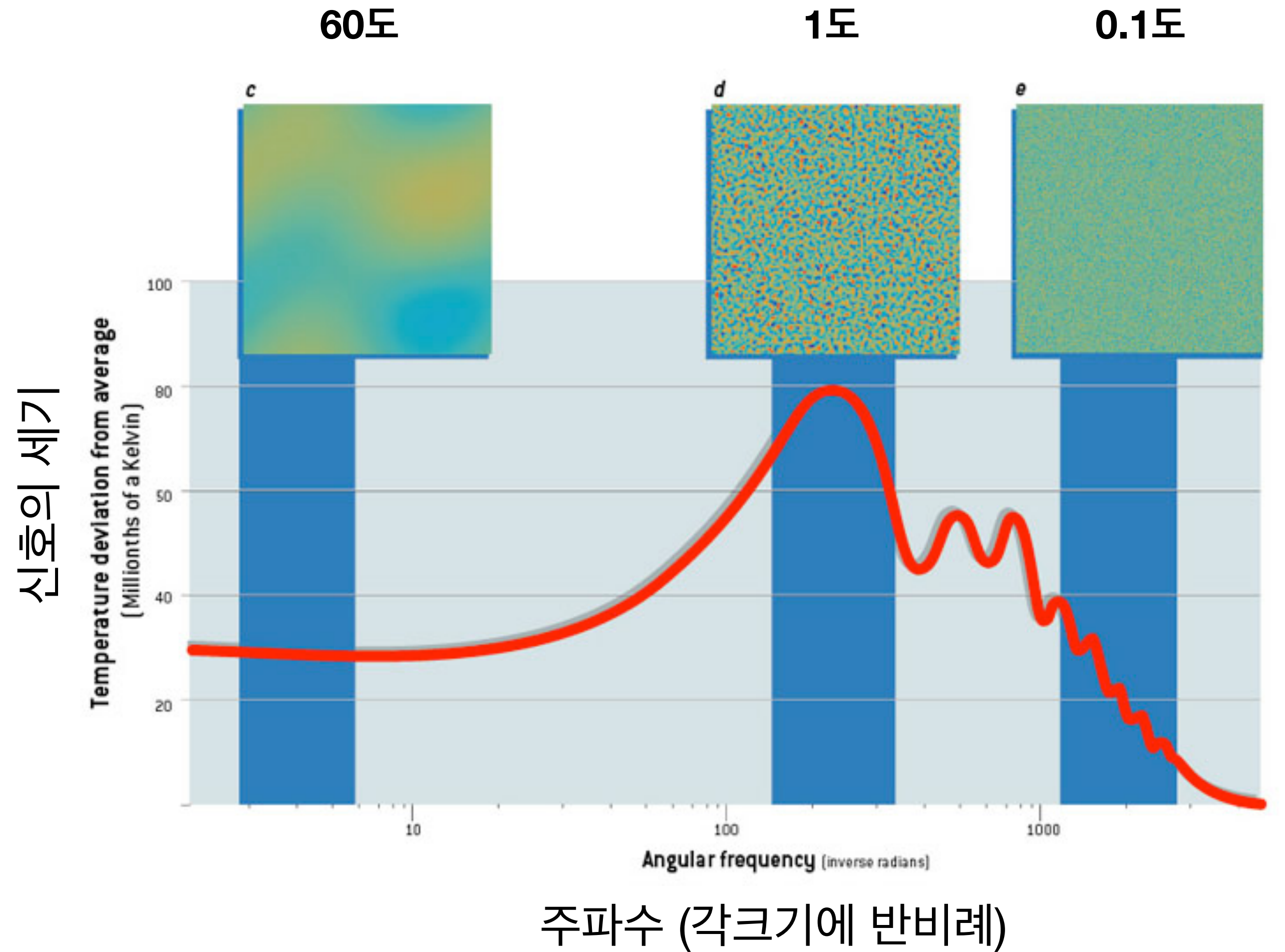
- Cosmology
 - Cosmic expansion history
 - Dark matter
 - Energy budget
- Baryon physics

CMBA

Power spectrum



360 도



CMBA

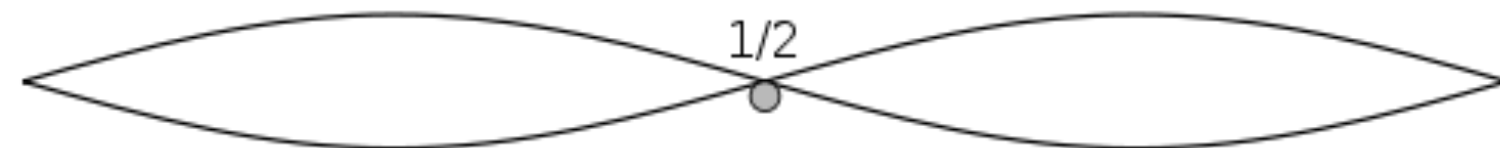
Sudden expansion (Inflation), producing density fluctuation



기본음 440Hz A4



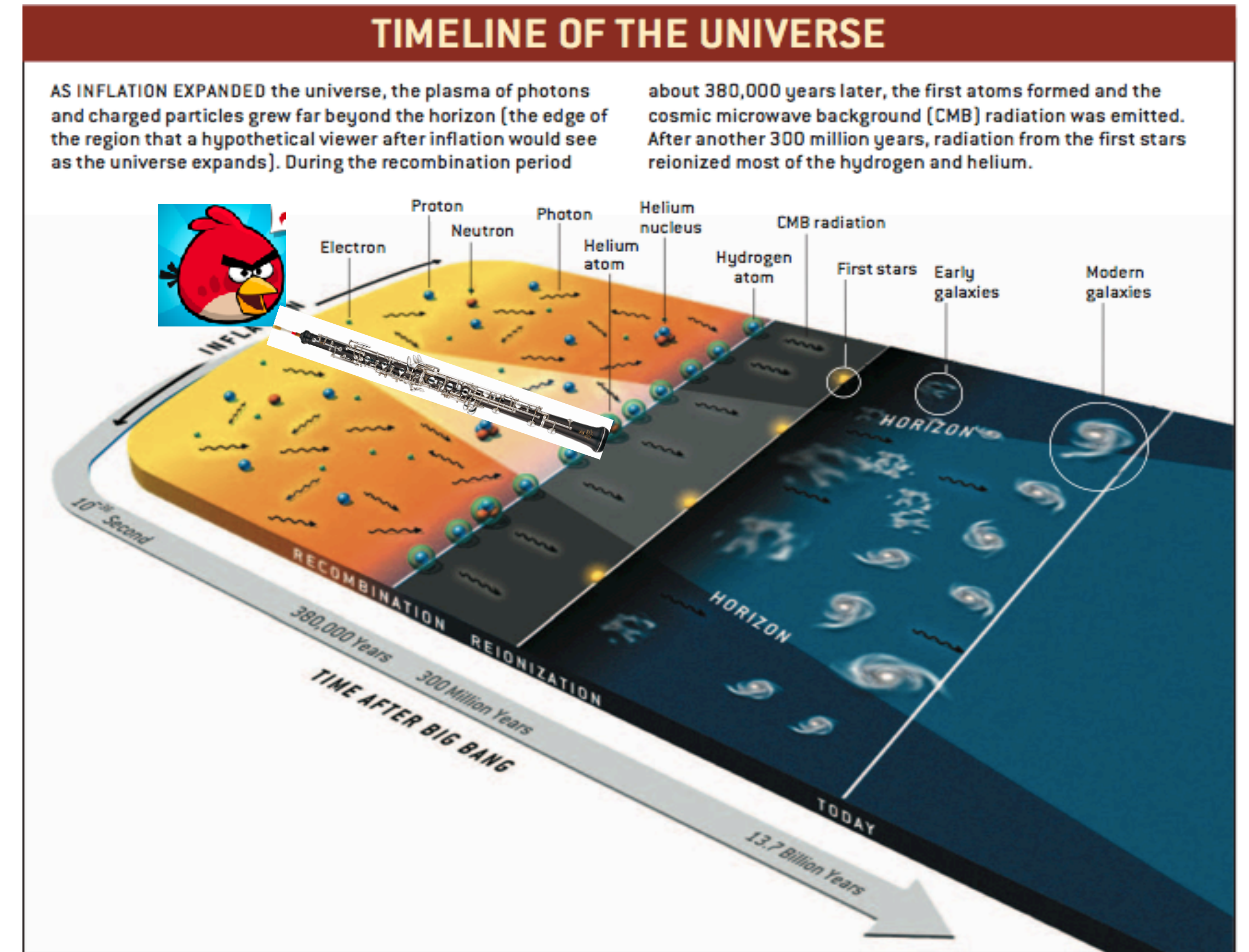
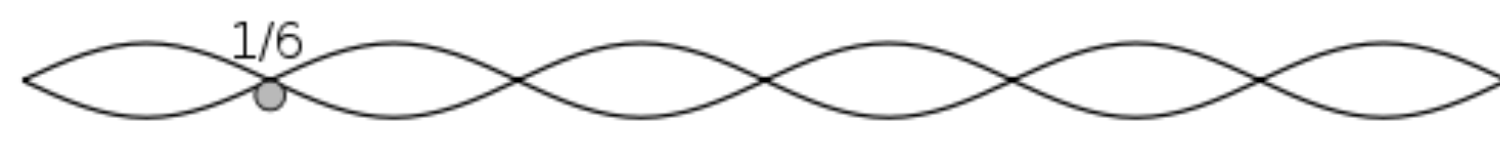
제1 배음 880Hz A5



제2 배음 1320Hz E6



제3 배음 1760Hz A6

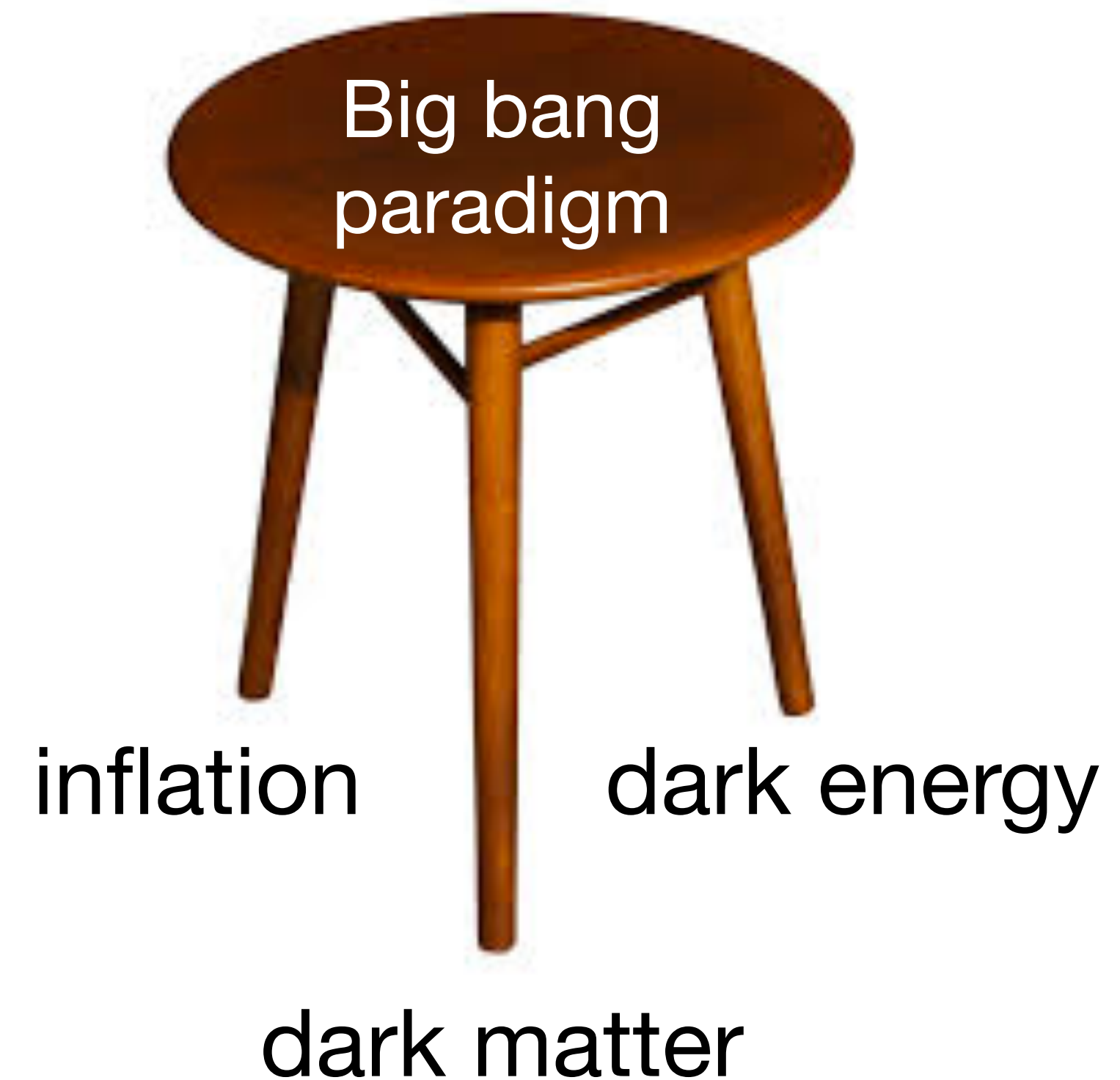


Concordance model

Summary

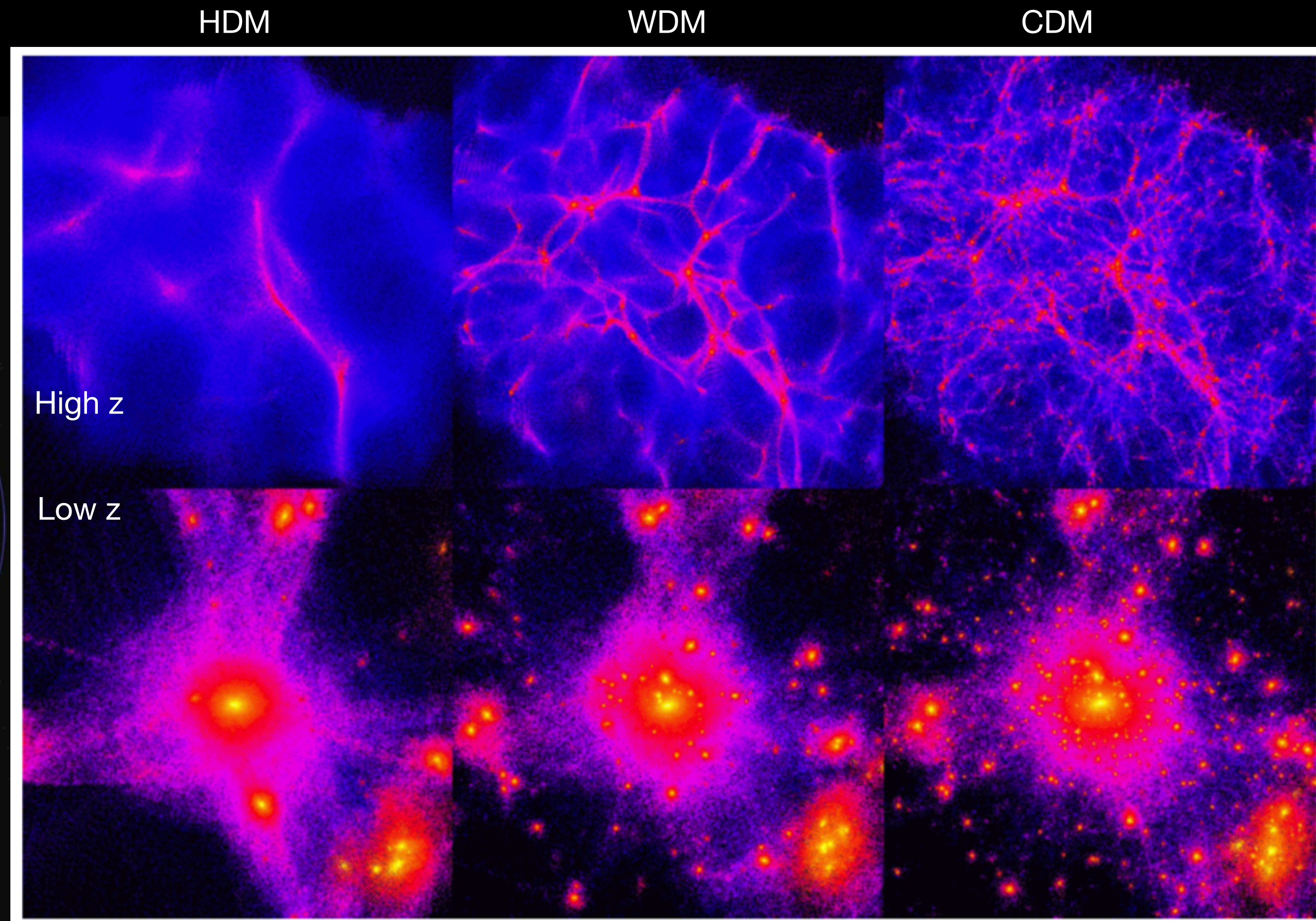
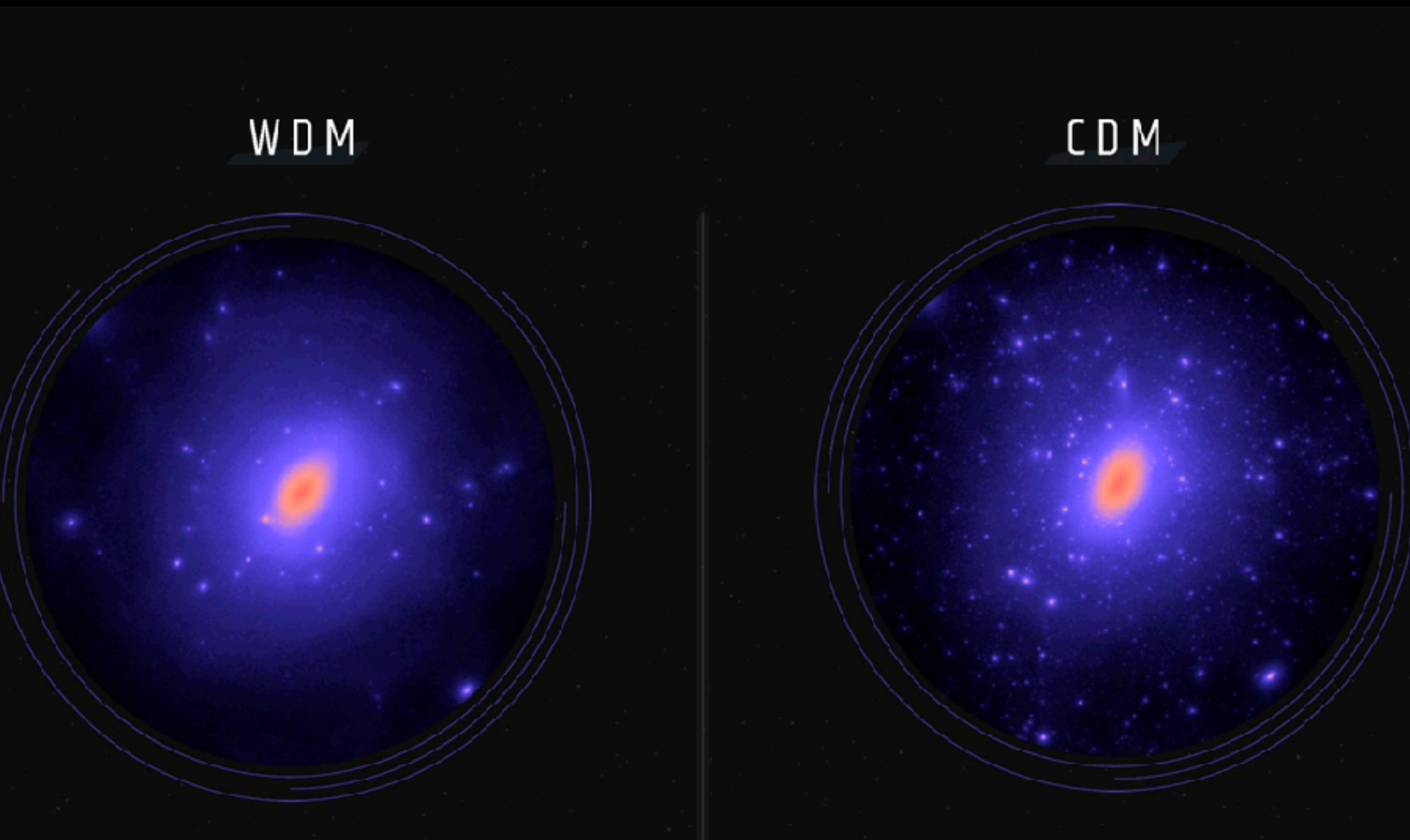
- $\Omega_{\text{bar}} \sim 0.04$
- $\Omega_{\text{DM}} \sim 0.24$
- $\Omega_{\Lambda} \sim 0.7$
- $\Omega_{\text{tot}} \sim 1.0$
- Space geometry is flat
- Age of Universe ~ 13.7 Gyr
- Primordial density fluctuation

$$\frac{\delta\rho}{\langle\rho\rangle} \sim 10^{-5}$$



Cold dark matter

More substructures. Small structures form first.



Kolchin 2017 / Simulations by V. Robles, T. Kelley, and B. Bazek+

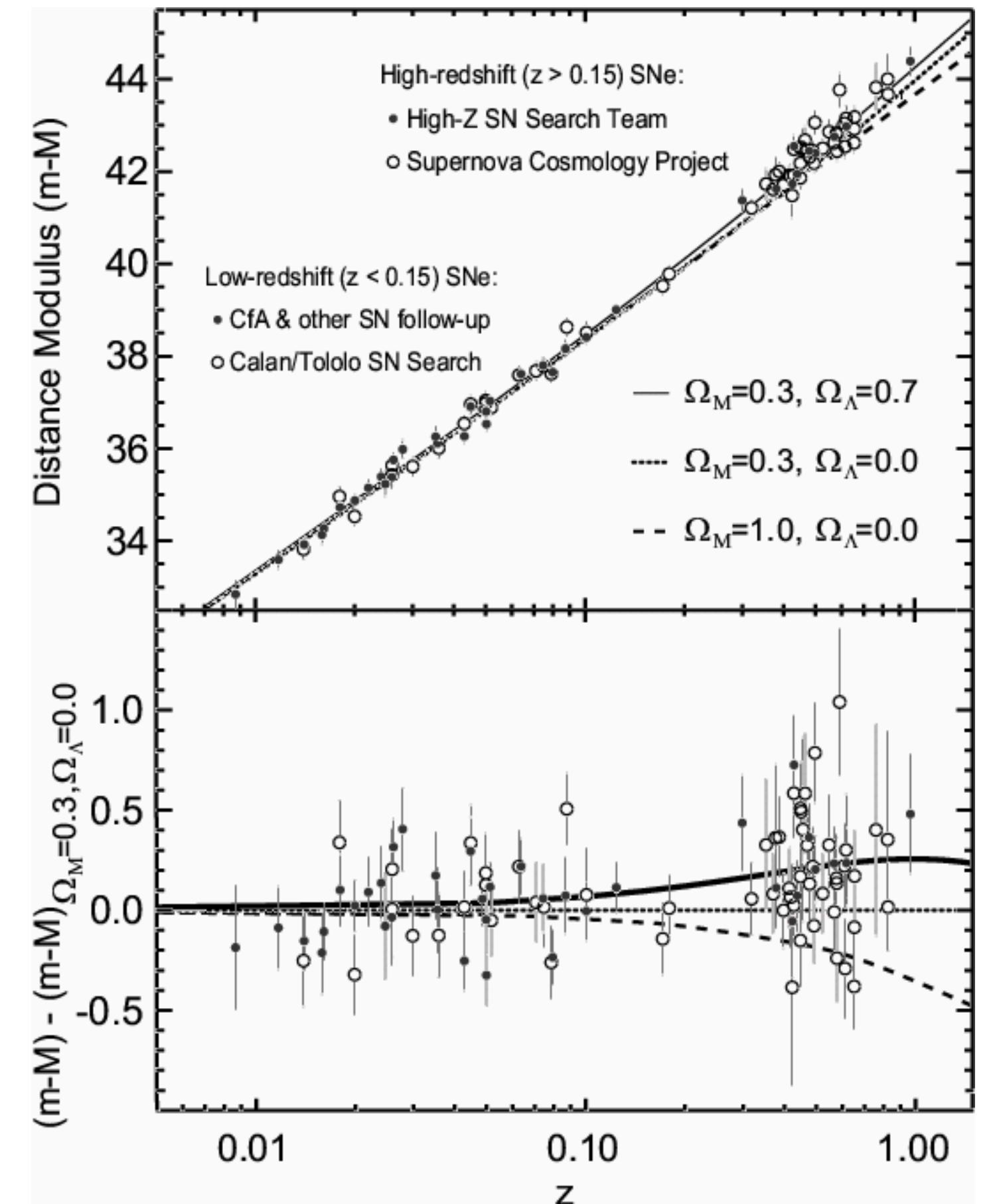
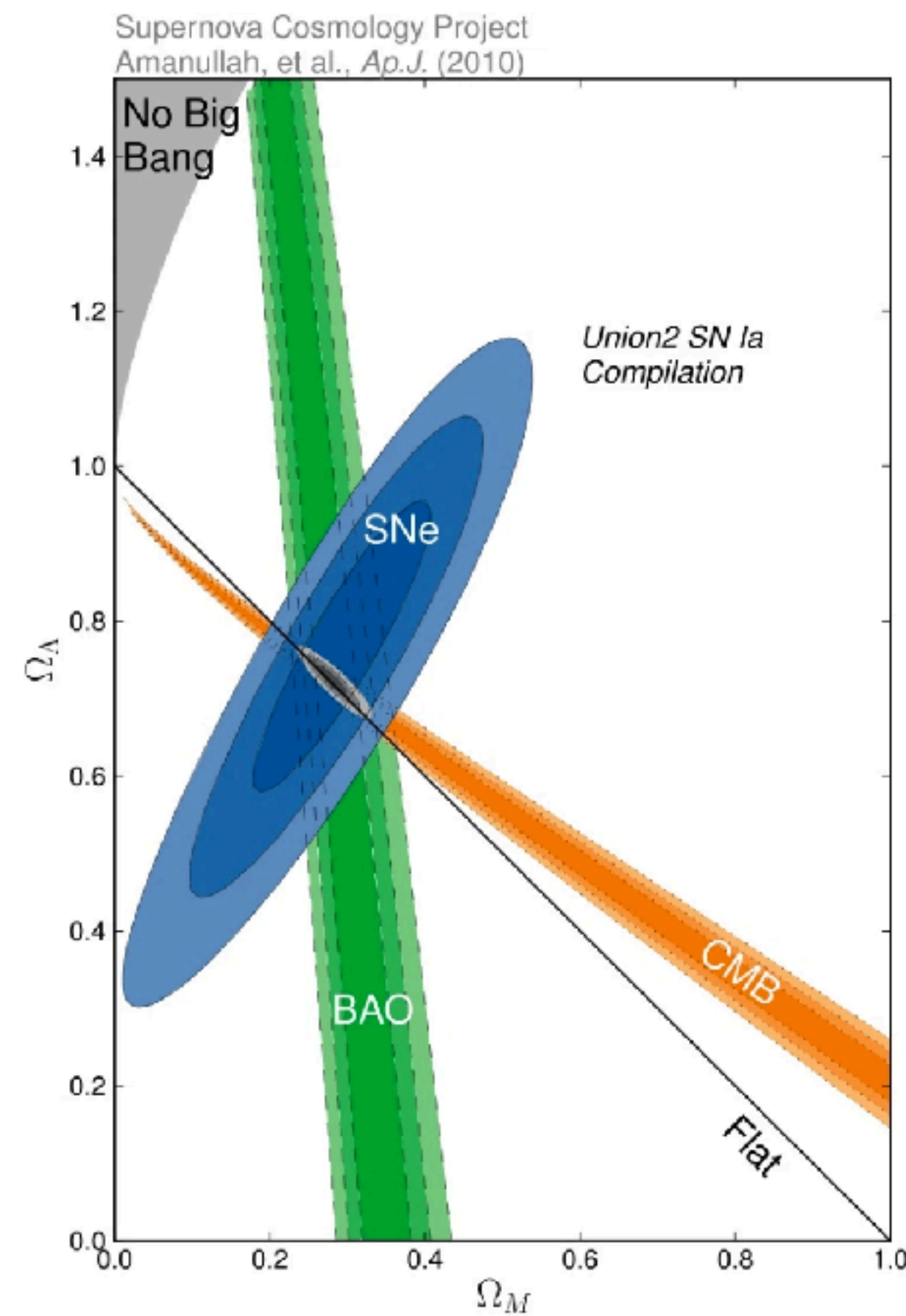
Credit: Bullock & Boylan-Kolchin 2017/ Simulations by V. Robles, T. Kelley, & B. Bazek+

Courtesy ITC @ University of Zurich

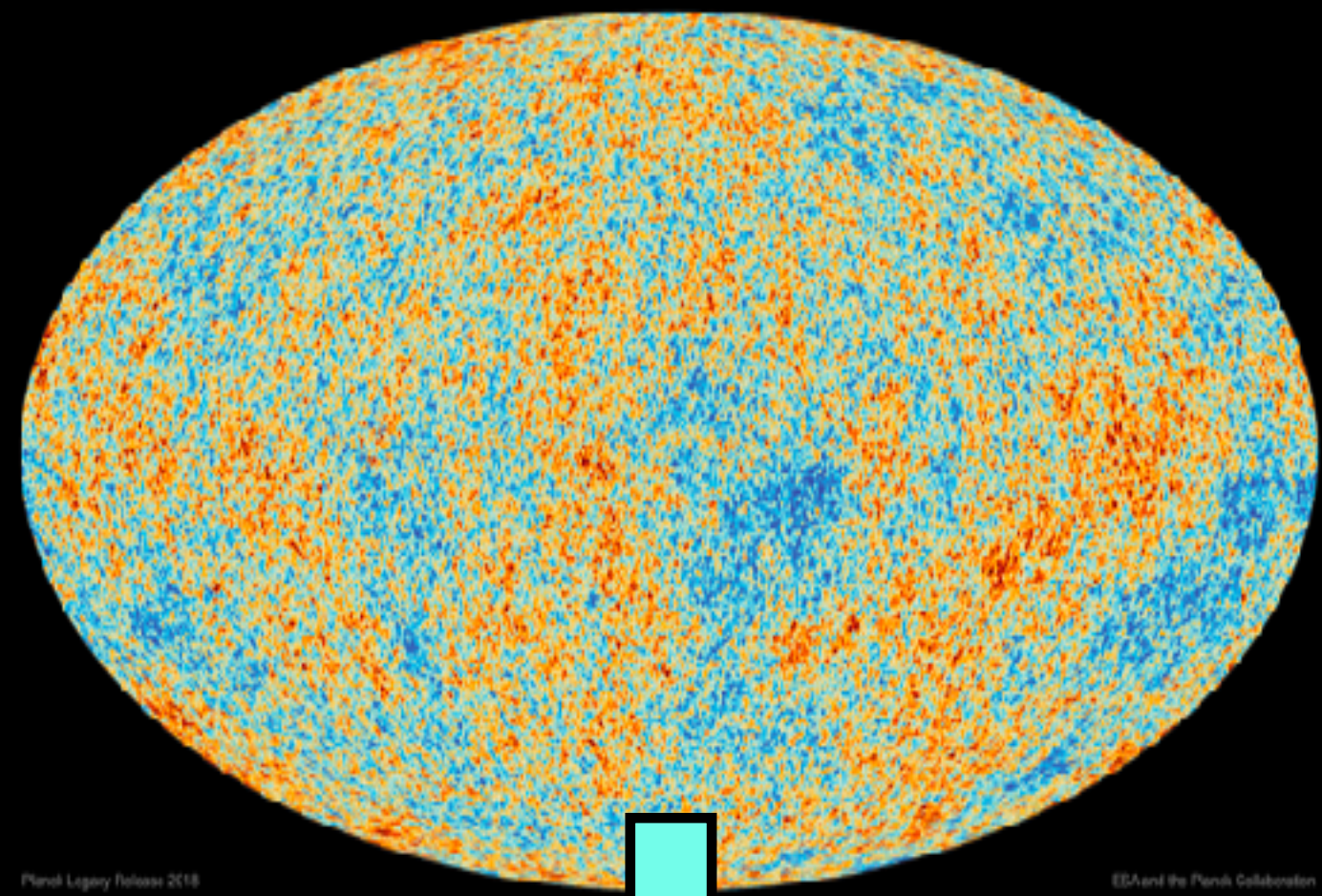
Dark energy

Accelerating expansion

- $\Omega_{\Lambda} \sim 0.7$
- Making the Universe geometry flat
- And accelerating expansion?
- Is supernova cosmology robust?



Perlmutter & Schmidt (2003)



Structure formation

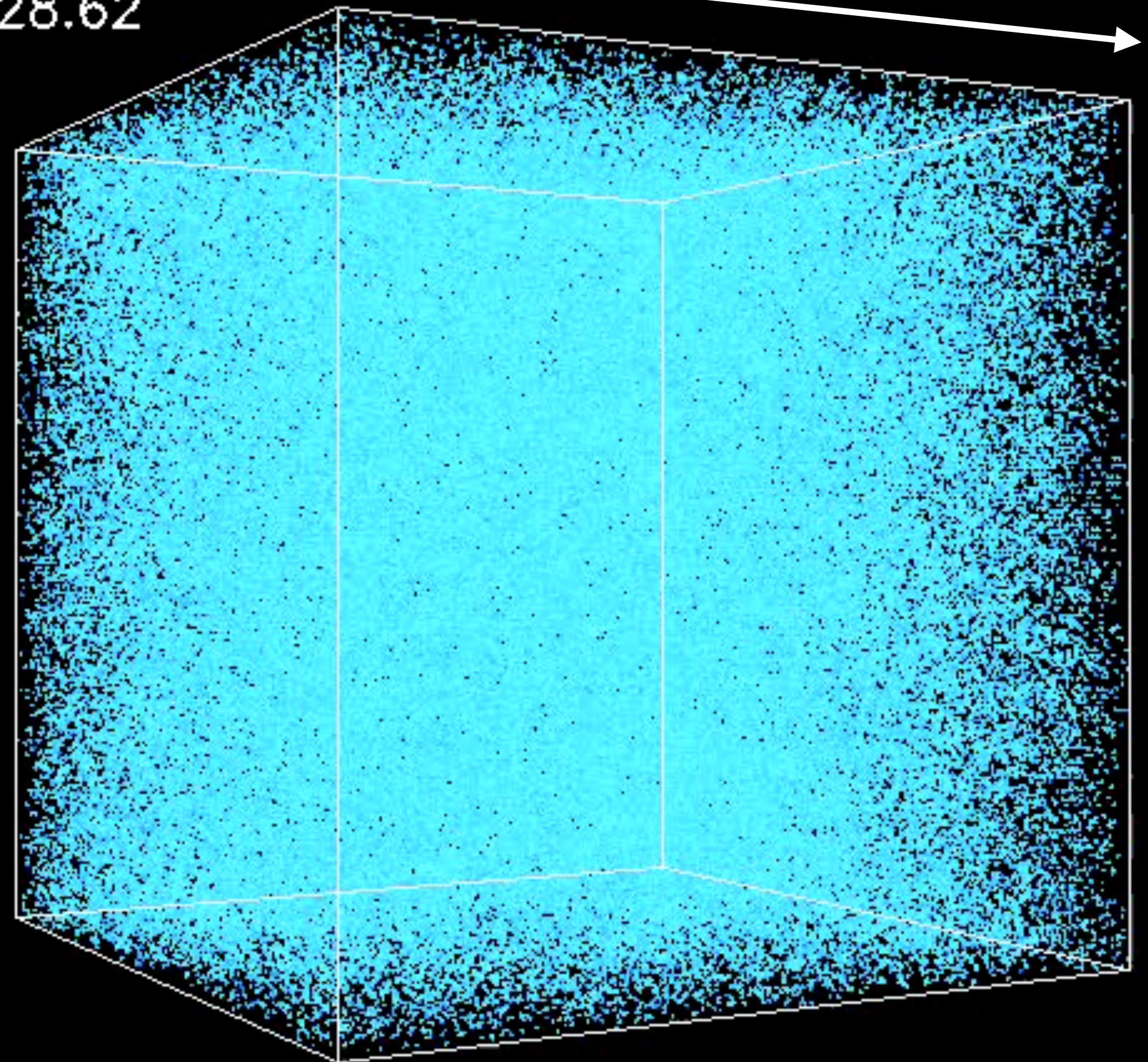
Primordial density fluctuation



누리온

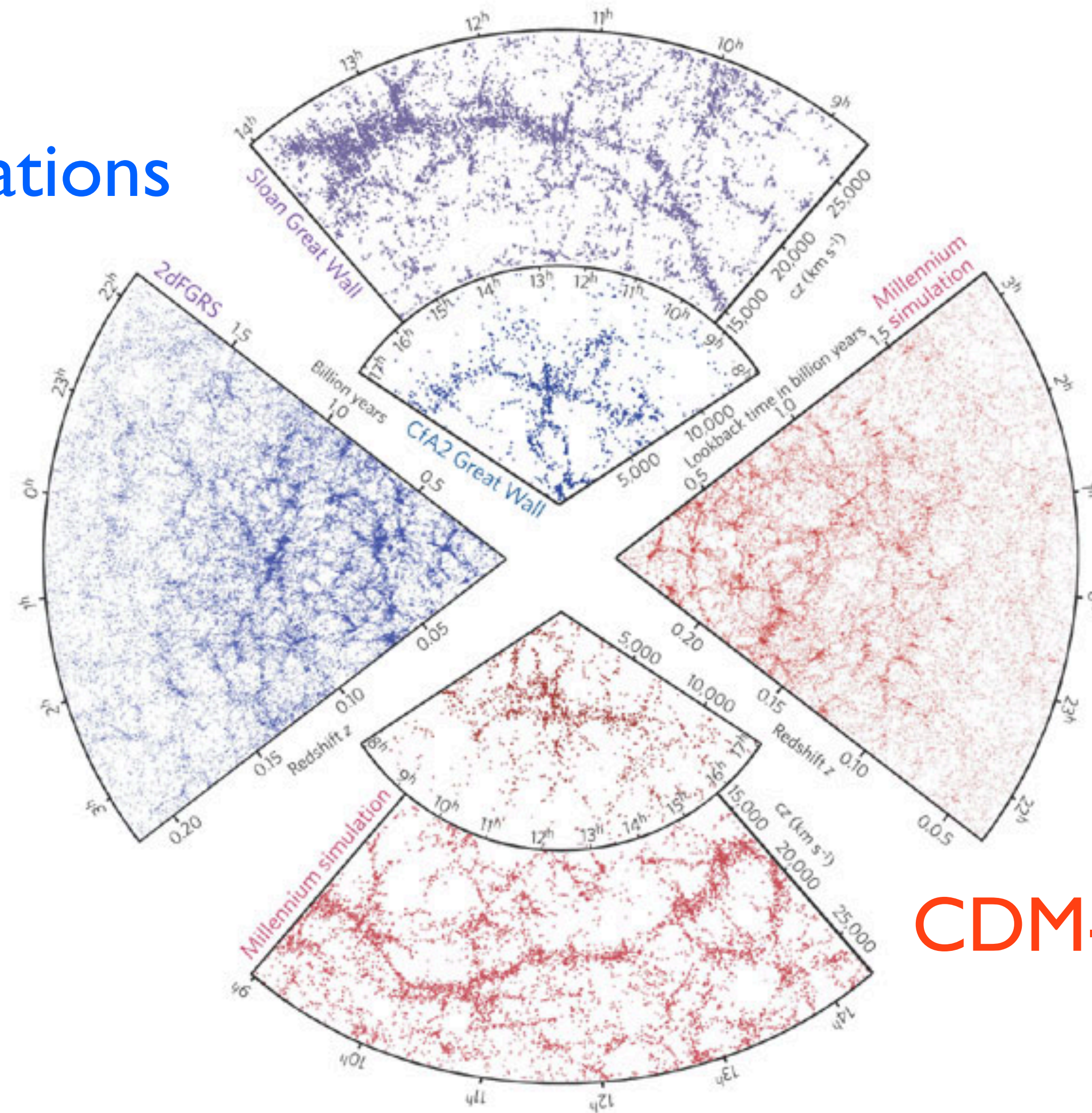
$Z=28.62$

1억 4천만 광년



Credit: A. Kravtsov & A. Klypin

Observations

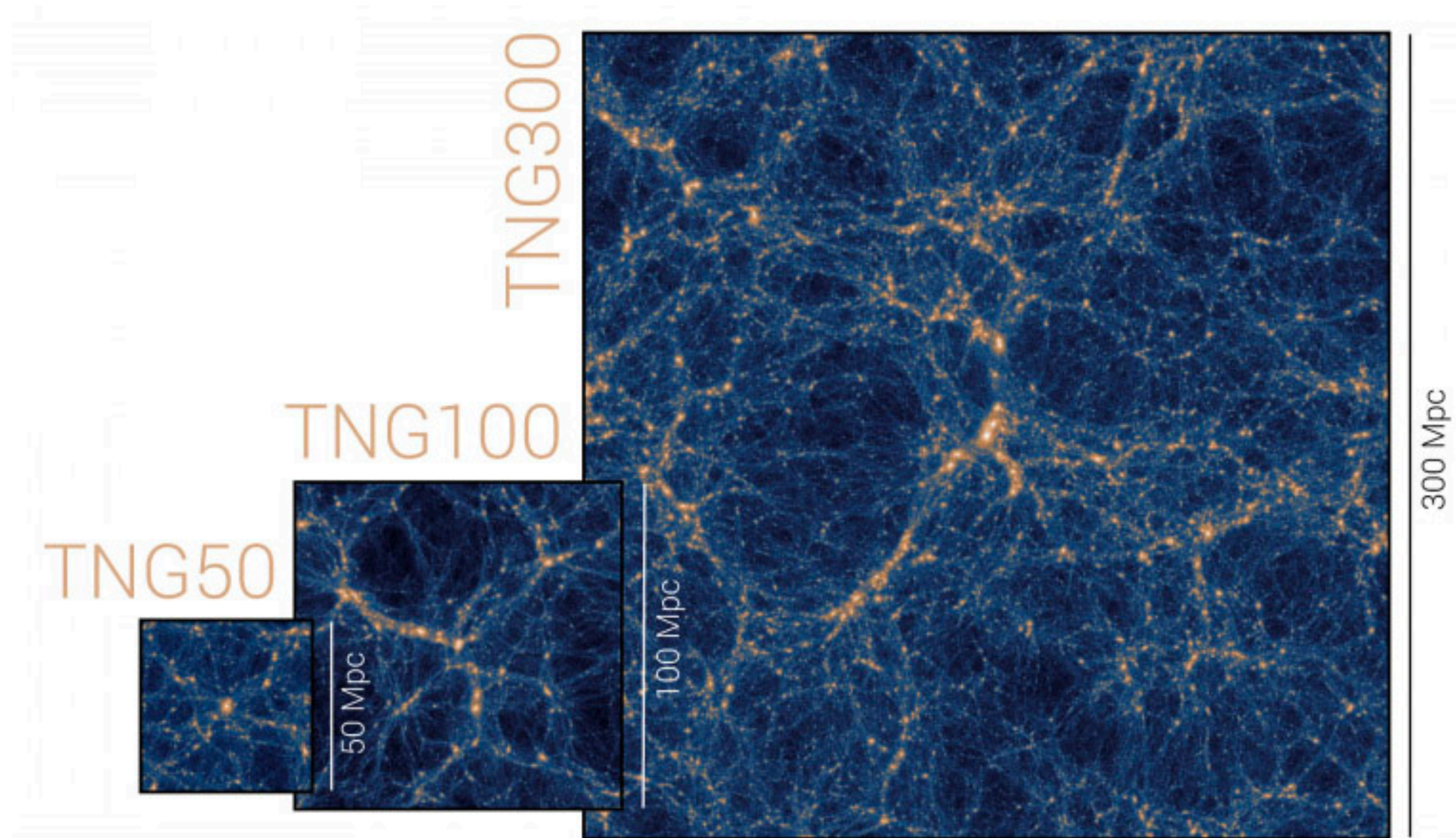


CDM-only simulation

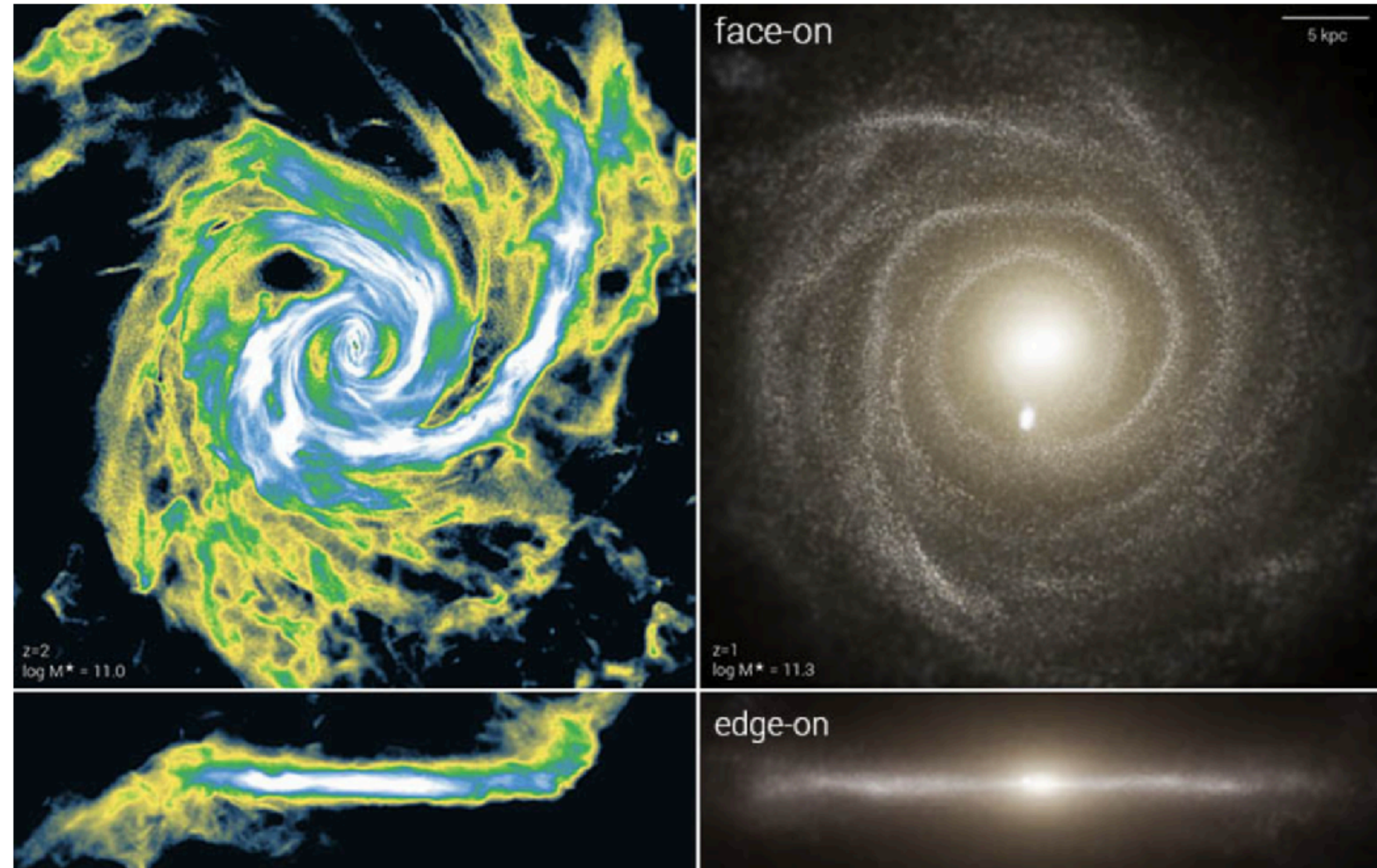
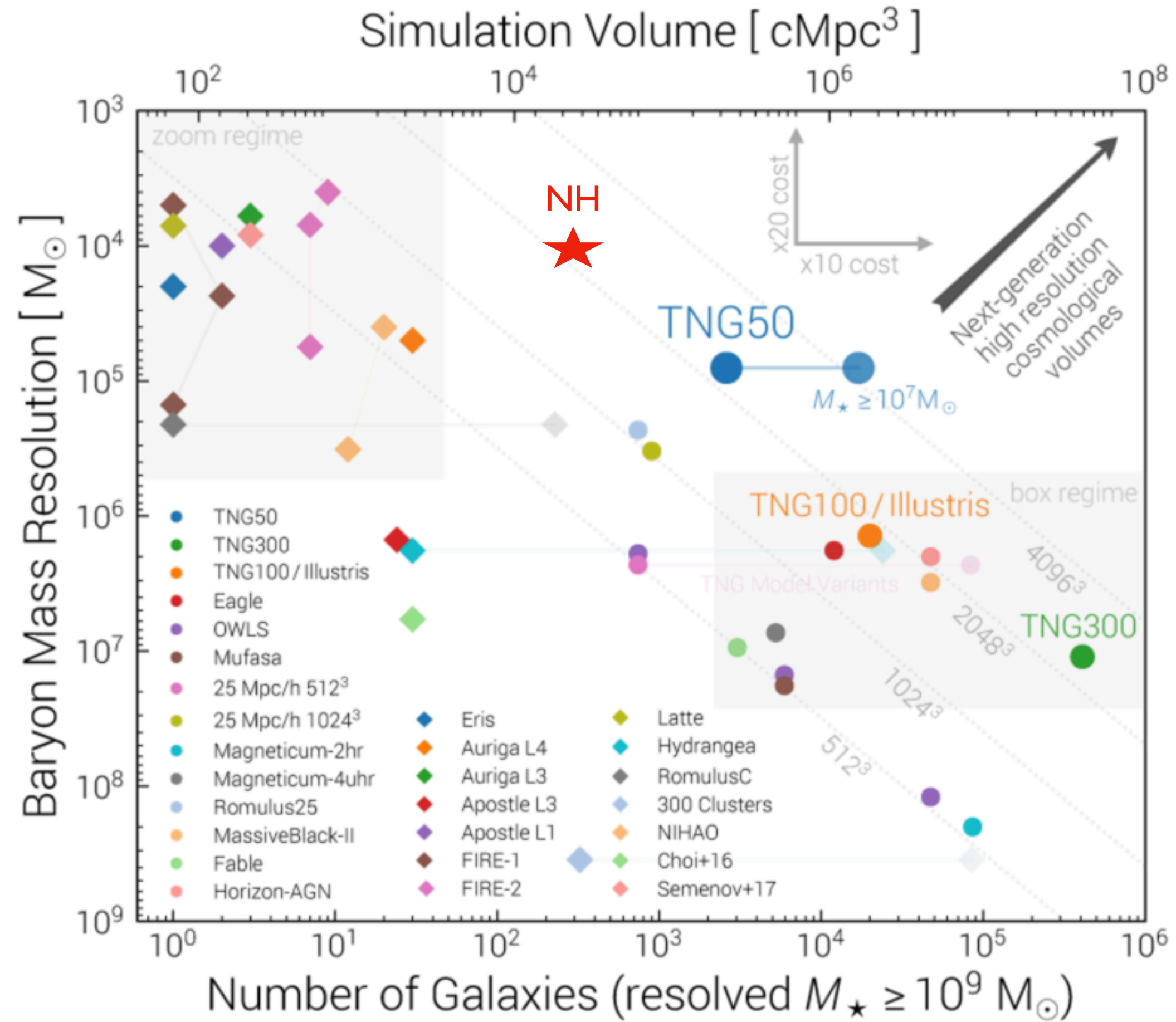
Cosmological hydrodynamic simulations

Of galaxy formation

- Gravity + **hydrodynamics**
- + **Astro-baryon** physics
 - Gas cooling, star formation
 - Black hole physics
 - Chemical evolution
- 2014: $dx \sim 1$ kpc
 - EAGLE (SPH)
 - Horizon-AGN (AMR)
 - Illustris (moving mesh)
- ~ 2020
 - IllustrisTNG50: $dx \sim 0.3$
 - NewHorizon: $dx \sim 0.04$

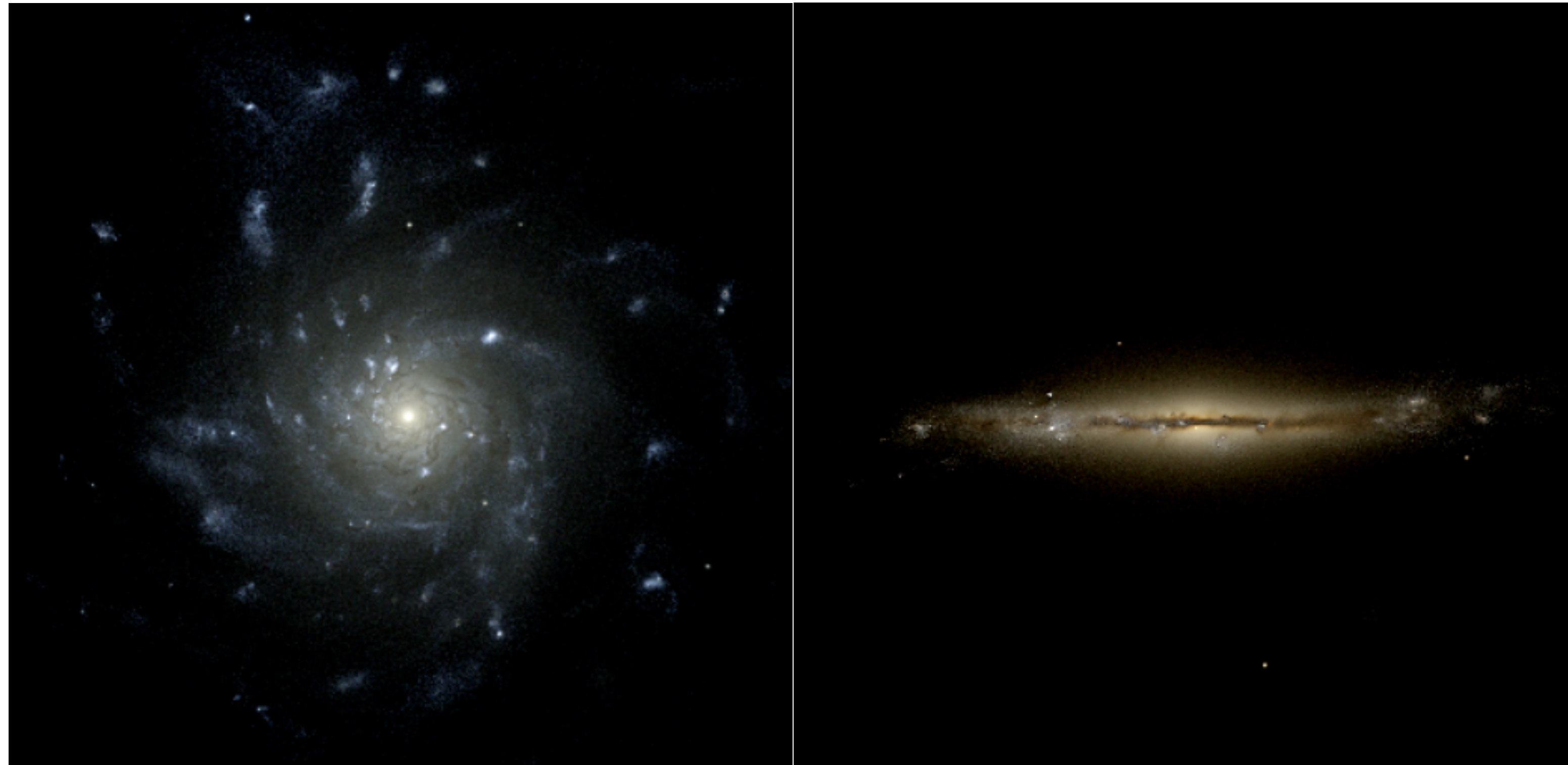
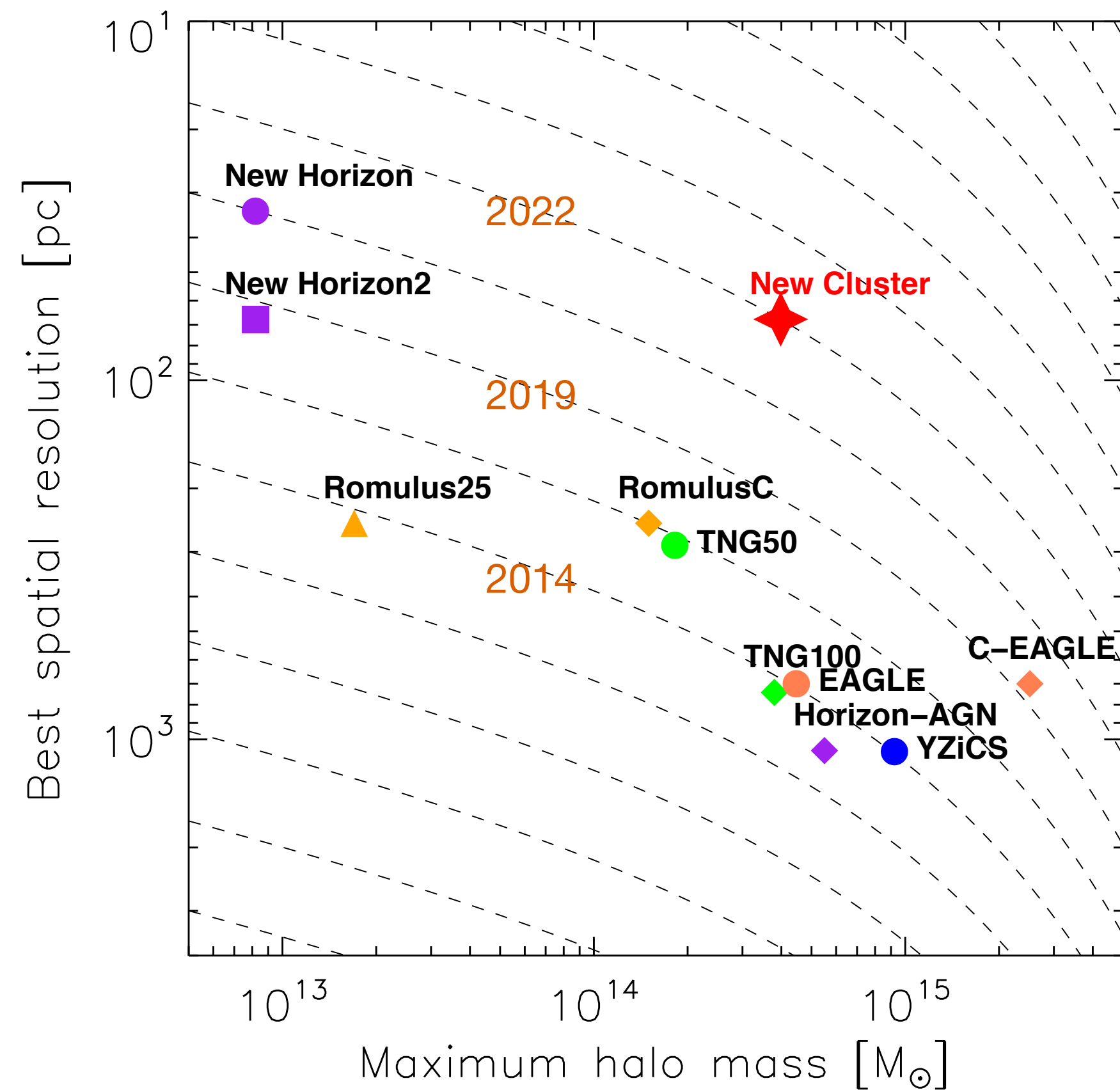


TNG & NewHorizon



Credit: <https://www.tng-project.org/about/>

TNG & NewHorizon



NewHorizon

슈퍼 컴퓨팅

NewHorizon simulation
IAP-Yonsei-Oxford coll.
4800 cores
80 Mhr (2017-2020)
6천만 광년

암흑 물질
고온 기체
별



5 cMpc

은하형성

Face-on view



Edge-on view



Galaxy challenge

Disc galaxies



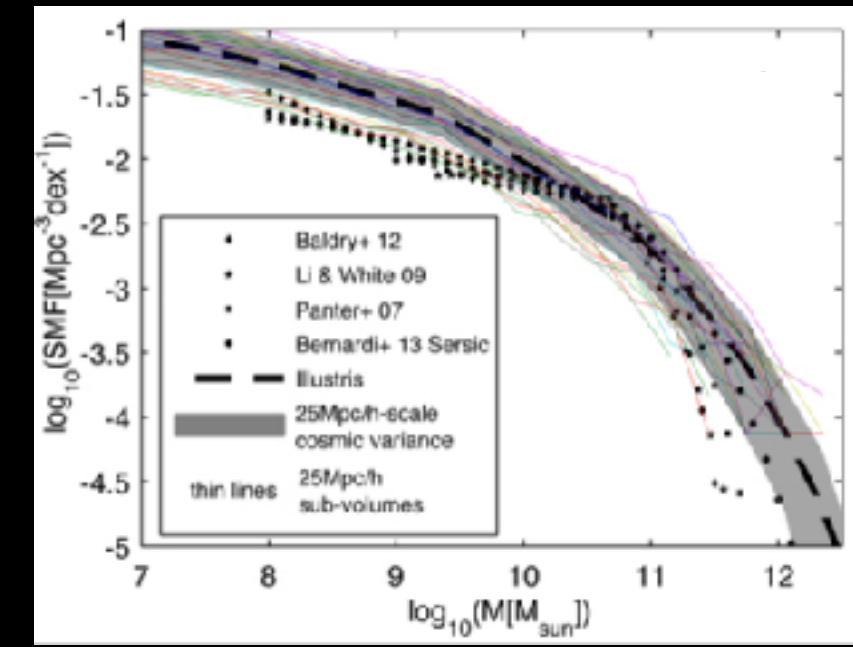
25000 ly



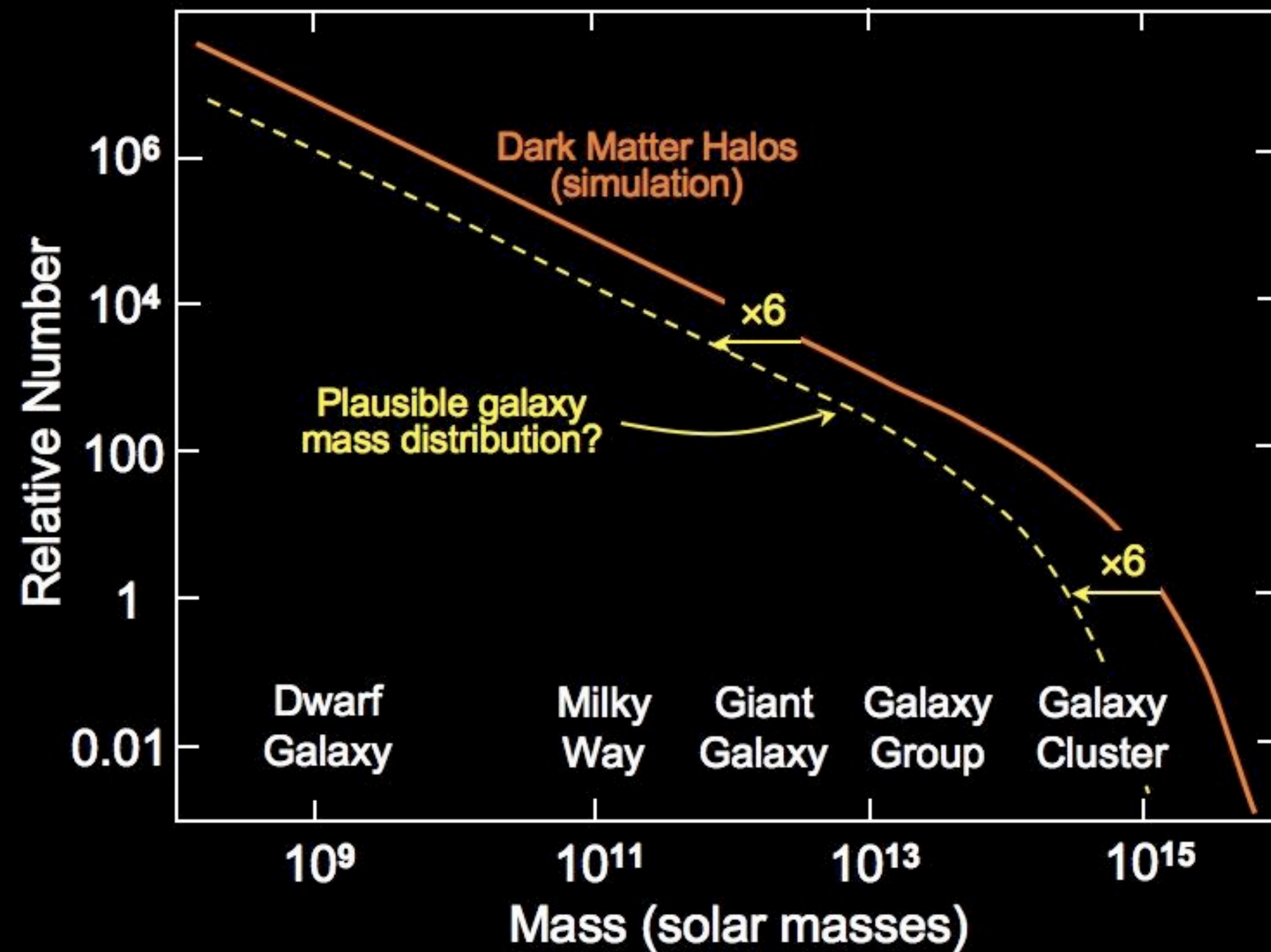
NGC 891

Shortcomings of models

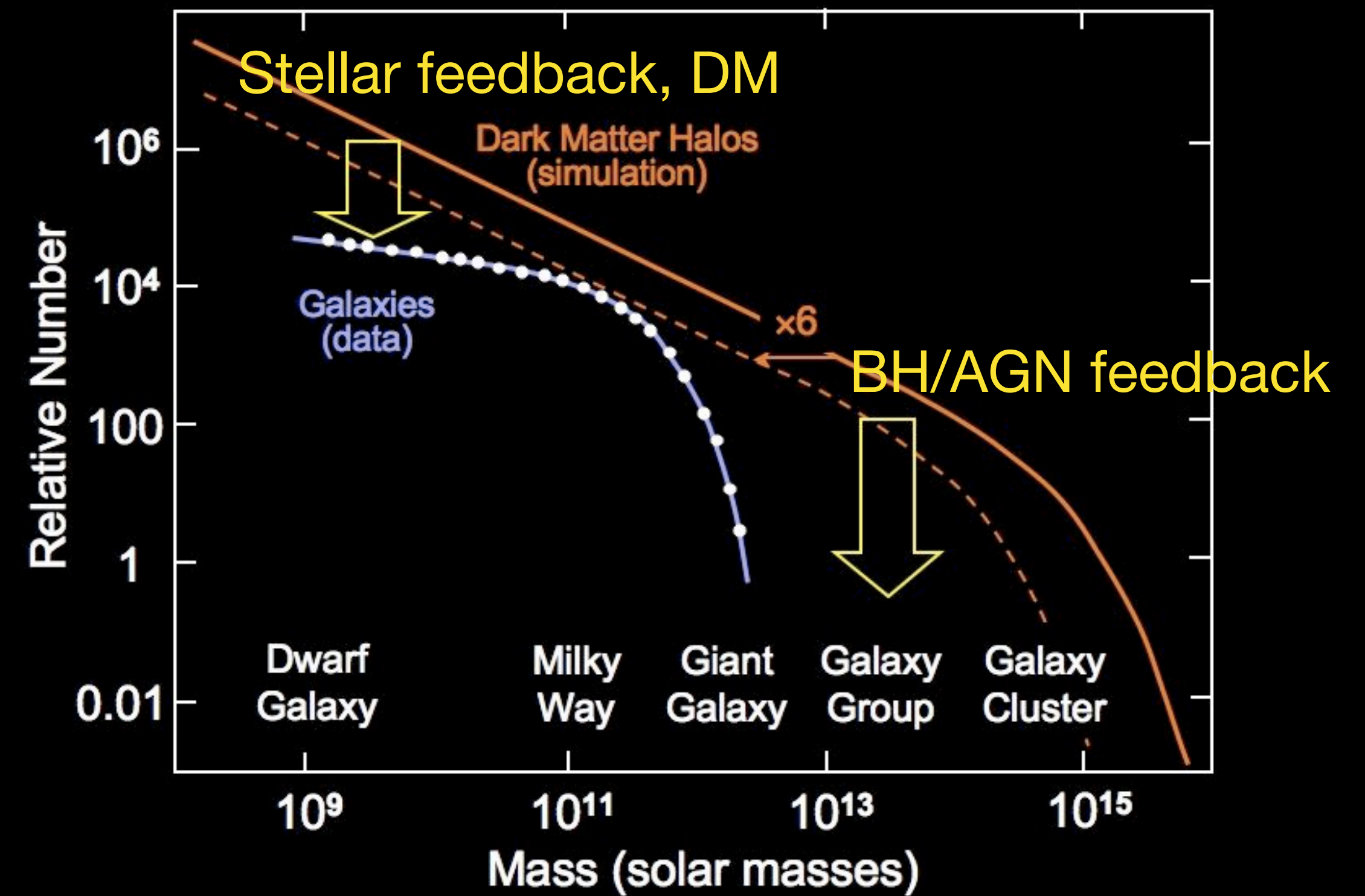
1) luminosity function problem



(Expected) Halo and Galaxy Mass Distributions



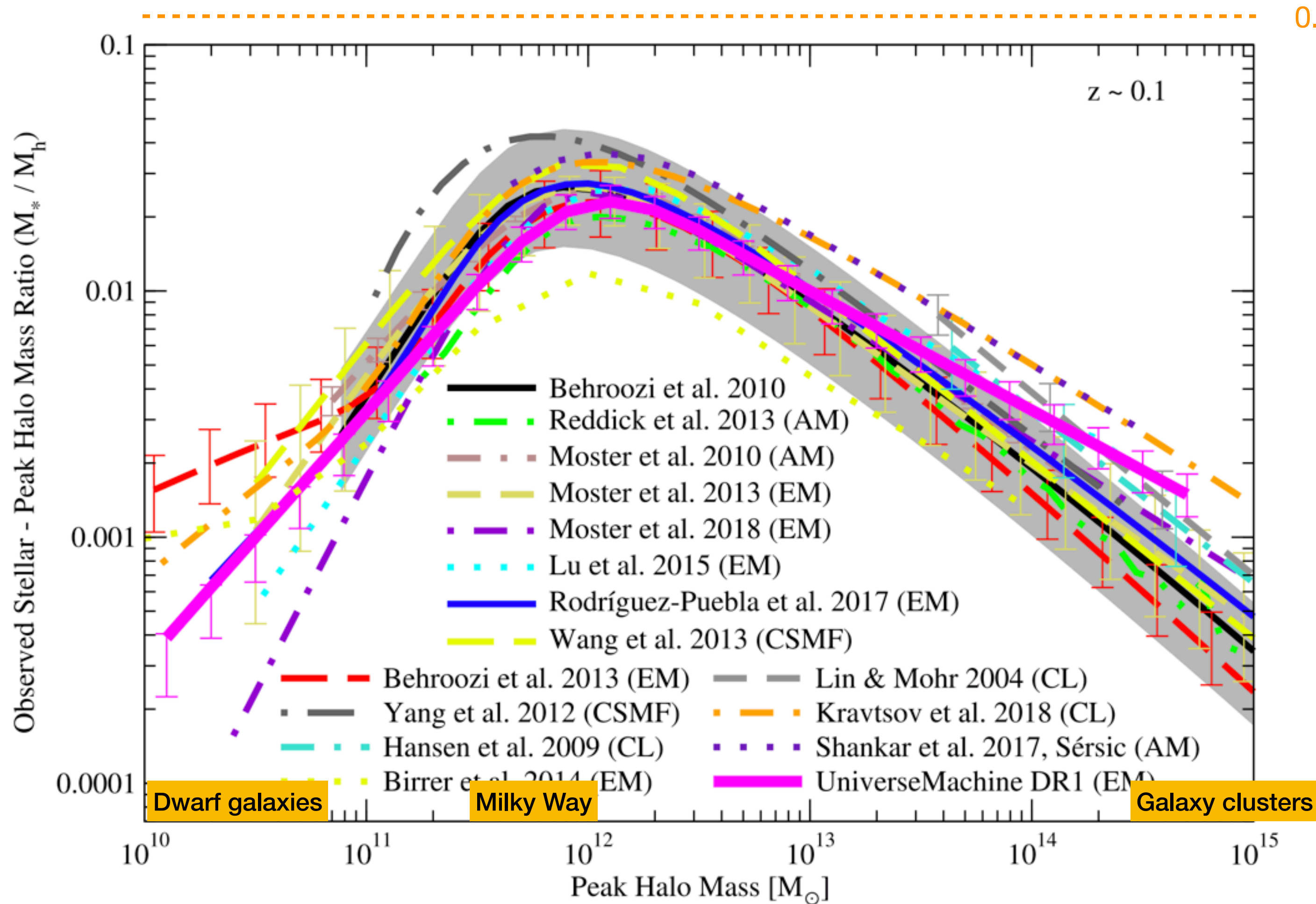
(Actual) Halo and Galaxy Mass Distributions



Shortcomings of models

1) luminosity function problem

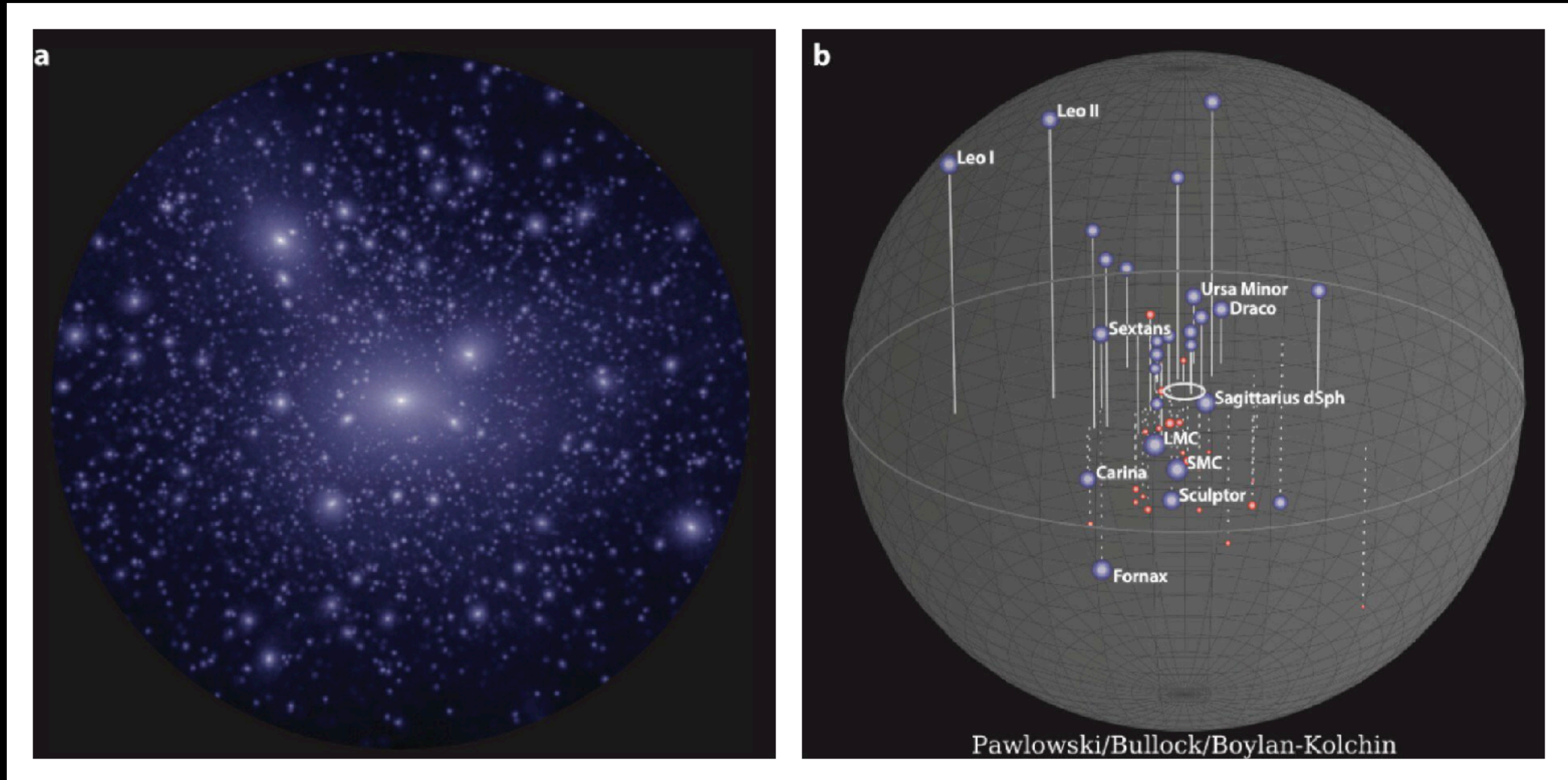
$$\frac{M_*}{M_{\text{halo}}}$$



Behroozi+ 2019

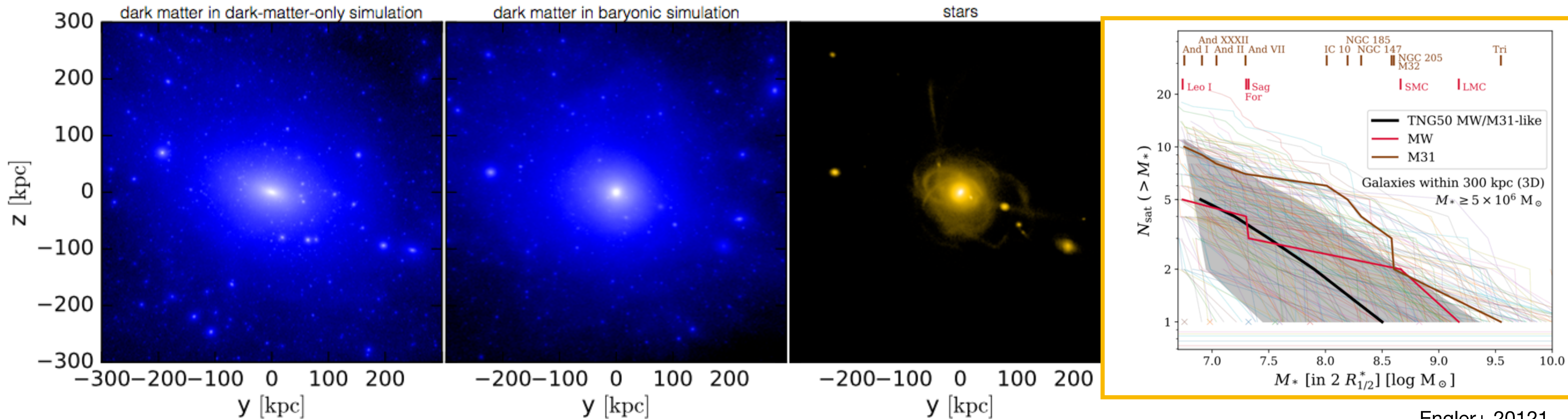
Shortcomings of models

2) missing satellite problem



Shortcomings of models

2) missing satellite problem



Engler+ 20121

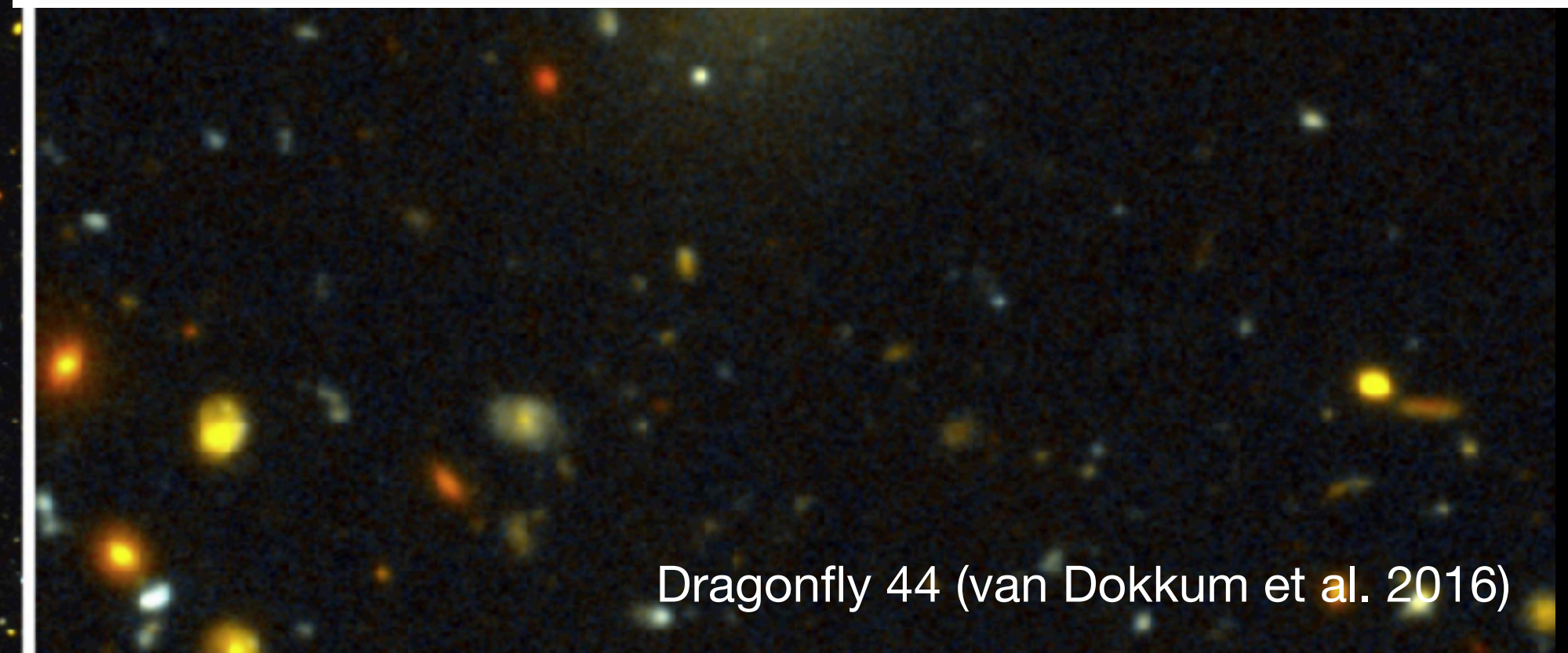
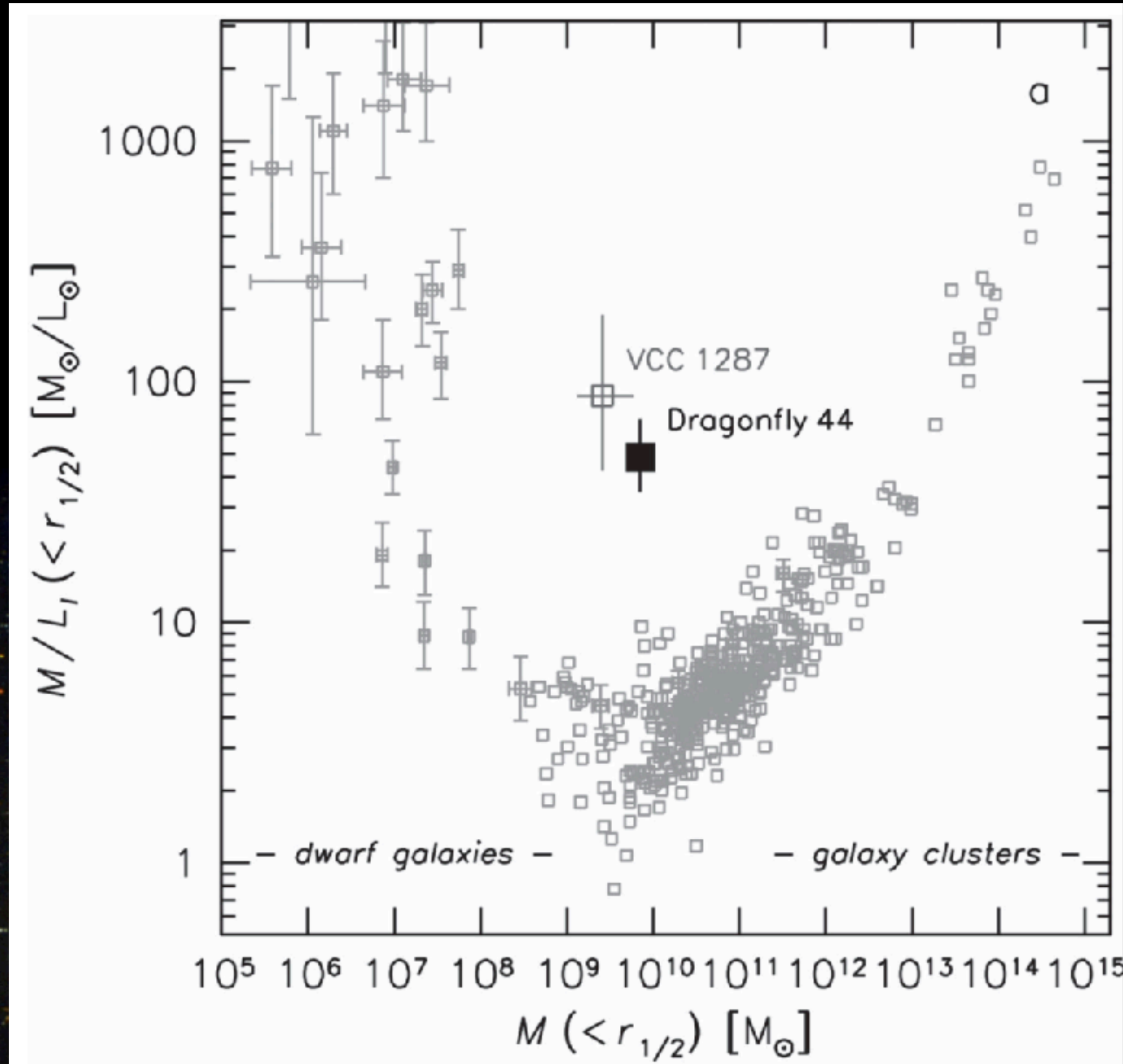
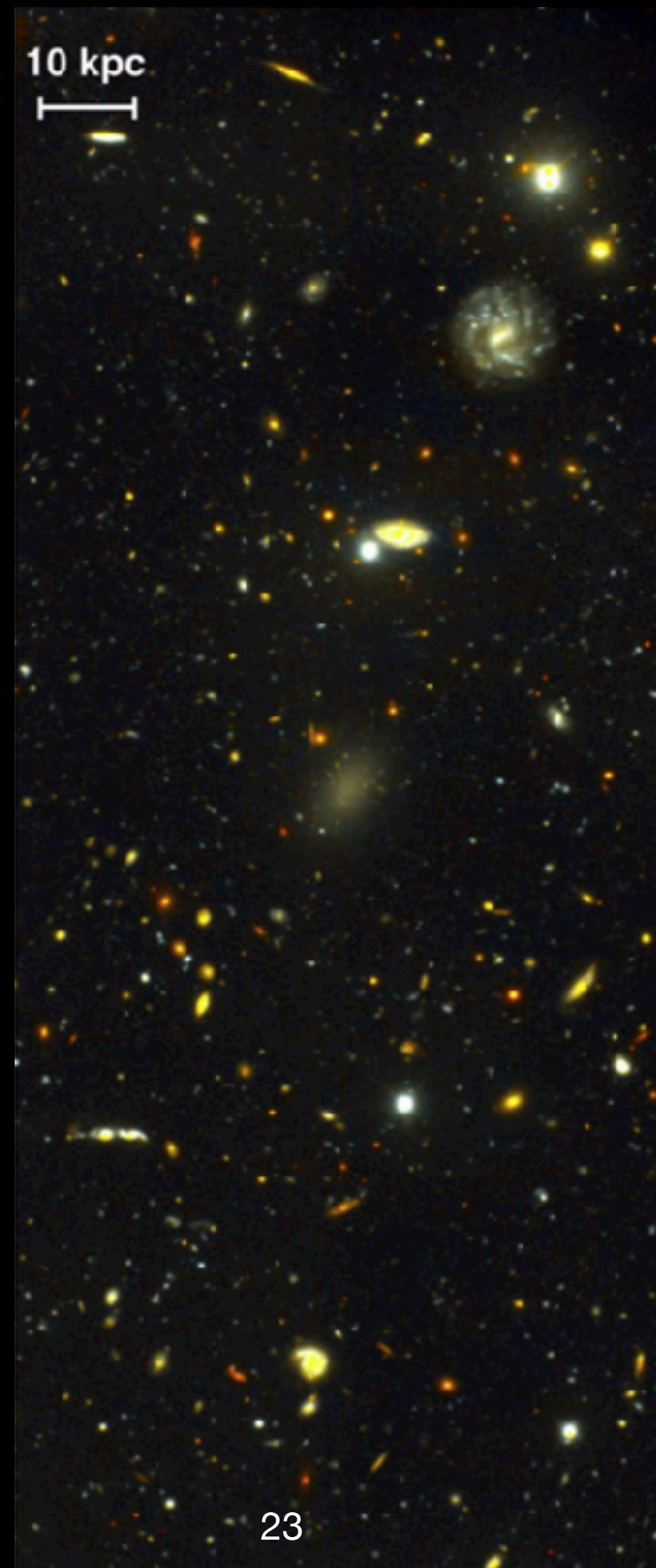
Figure 1. Projected mass surface densities of Milky Way-mass host in the Latte simulation at $z = 0$: dark matter in the dark-matter-only simulation (left); dark matter (middle) and stars (right) in the baryonic simulation. Color scale is logarithmic, spanning $10^4 - 10^9 M_{\odot} \text{ kpc}^{-2}$, same in all panels. Compared with dark-matter-only, the baryonic simulation contains *significantly* ($\approx 10\times$) fewer subhalos at fixed $V_{\text{circ,max}}$. Of these subhalos, only 9 host a satellite galaxy with $M_{\text{star}} > 3 \times 10^5 M_{\odot}$.

Wetzel+ 2016

Shortcomings of models

3) ultra-diffuse galaxies

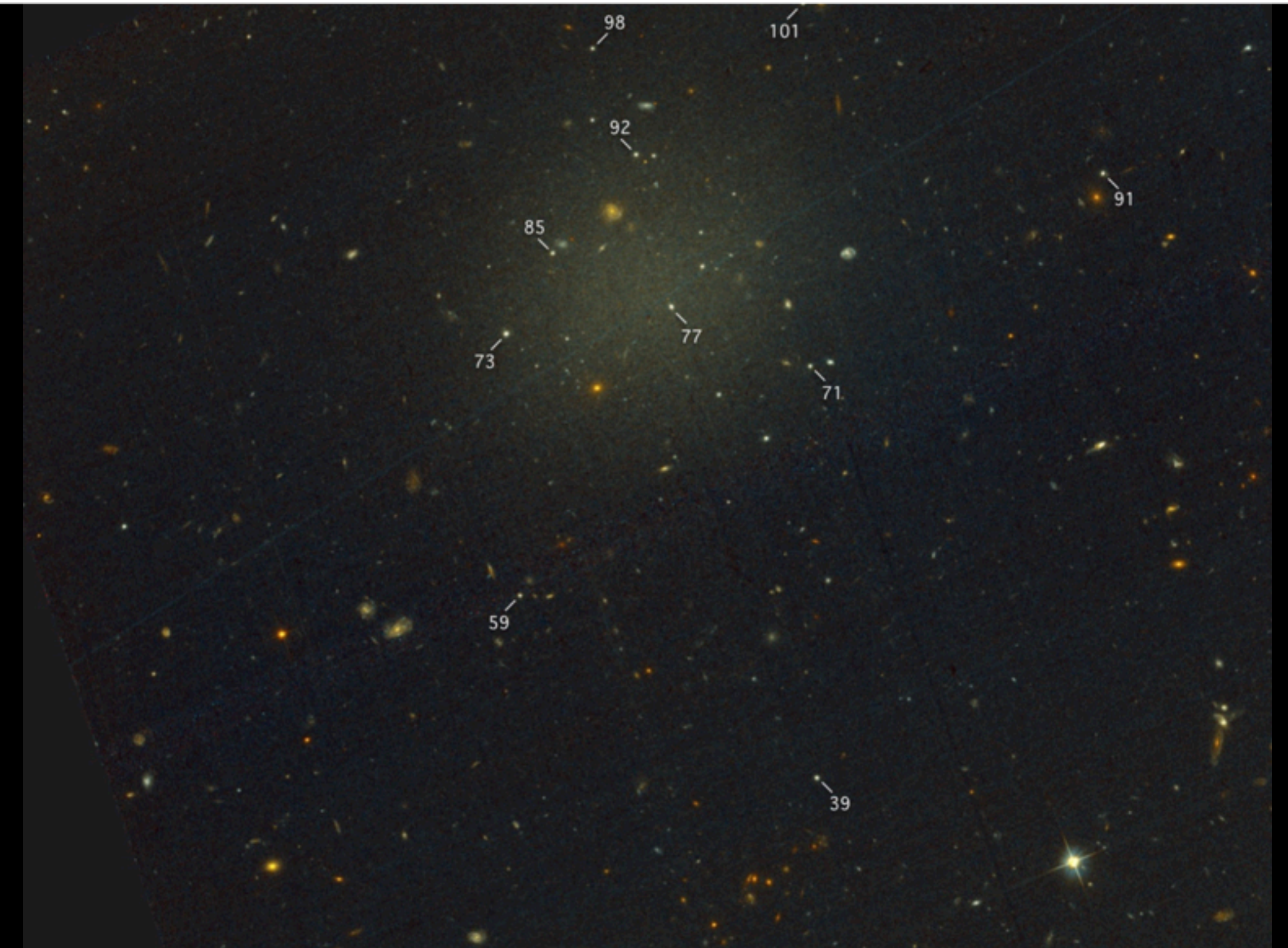
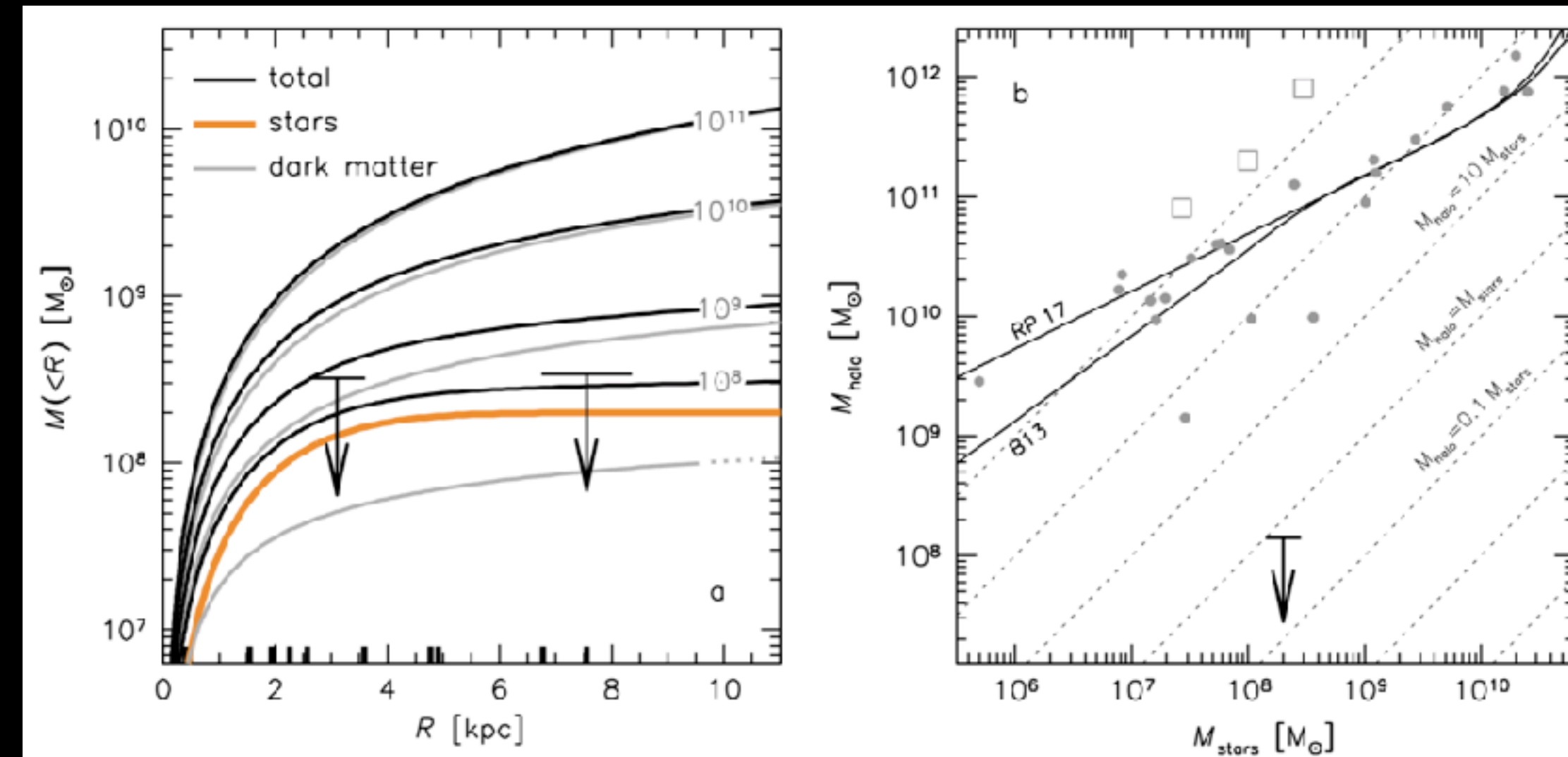
- Dragonfly 44 (van Dokkum+ 2016)
 - In Coma cluster
 - 33.5 hr with Deimos/Keck II
 - Based on 94 GCs
 - $M_{\text{dyn}} \sim MW$
 - $M_* \sim 1/20 MW$
 - DM dominant?



Shortcomings of models

3) ultra-diffuse galaxies

- NGC 1052-DF2 (van Dokkum+ 2018)
 - NGC 1052 group
 - HST + Deimos/Keck II
 - Based on 10 GCs
 - $M_{\text{dyn},3.5R_e} \lesssim 3.4 \times 10^8 M_{\odot}$
 - $M_* \sim 2 \times 10^8 M_{\odot} \sim 1/300 \text{ MW}$
 - $R_e \sim 2.2 \text{ kpc} \sim \text{MW}$
 - Compatible with no DM



NGC 1052-DF2 (van Dokkum et al. 2018)

Outstanding questions

- Is dark energy real?
- Is dark matter real? Is it cold?
- Are the issues of the strange properties of galaxies caused by DM or astrophysical nature?
- Is the missing satellite problem gone? (NewHorizon)
- Ultra diffuse galaxies (NewHorizon)
- Current efforts are heavily biased to astrophysics.