

Probing microscopic origins of axions by the chiral magnetic effect

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Workshop on Physics of Dark Cosmos, Busan, Korea, Oct 22nd, 2022

Outline

- 3 possible classes of microscopic origins of axions
- (Measurable) distinctive patterns of axion couplings depending on the microscopic origins
- Measuring the axion-electron coupling by the chiral magnetic effect

Strong CP problem and QCD axion

$$y_u H Q_L u_R^c + y_d H^* Q_L d_R^c + \frac{g_s^2}{32\pi^2} \theta G \tilde{G}$$



$$\bar{\theta} = \theta + \arg \det (y_u y_d) < 10^{-10}$$

Non-observation
of neutron EDM

[Abel et al '20]

CPV in the QCD sector

while $\delta_{\text{CKM}} = \arg \det [y_u y_u^\dagger, y_d y_d^\dagger] \sim \mathcal{O}(1)$

The QCD vacuum energy is minimized at the CP-conserving point ($\bar{\theta} = 0$).

[Vafa, Witten '84]

$$V_{\text{QCD}} = -\Lambda_{\text{QCD}}^4 \cos \bar{\theta}$$

Promote $\bar{\theta}$ to a dynamical field (=QCD axion) : $\frac{g_s^2}{32\pi^2} \left(\theta + \frac{a}{f_a} \right) G \tilde{G}$

[Peccei, Quinn '77, Weinberg '78, Wilczek '78]

QCD axion lagrangian

$$\mathcal{L} = \frac{1}{2}(\partial_\mu a)^2 + \frac{g_s^2}{32\pi^2} c_G \frac{a}{f_a} G^{\mu\nu} \tilde{G}_{\mu\nu} \\ + \frac{a}{f_a} \sum_{A=W,B,\dots} \frac{g_A^2}{32\pi^2} c_A F^{A\mu\nu} \tilde{F}_{\mu\nu}^A + \frac{\partial_\mu a}{f_a} \left(\sum_{\psi=q,\ell,\dots} c_\psi \psi^\dagger \bar{\sigma}^\mu \psi + \sum_{\phi=H,\dots} c_\phi \phi^\dagger i \overleftrightarrow{D}^\mu \phi \right)$$

$$U(1)_{PQ} : a(x) \rightarrow a(x) + \alpha$$

broken by $c_G \neq 0$ non-perturbatively

$$\Rightarrow m_a^2 \simeq c_G^2 \frac{m_u m_d}{(m_u + m_d)^2} \frac{m_\pi^2 f_\pi^2}{f_a^2}$$

The axion couplings to the other SM particles

$c_W, c_B, c_q, c_\ell, c_H$ are UV model-dependent.

Axion-Like Particles (ALPs)

- Cousins of the QCD axion, while not being necessarily involved in the strong CP problem (so c_G can be 0)
- Ubiquitous in many BSM scenarios, in particular, string theory

[Arvanitaki, Dimopoulos, Dubovsky, Kaloper, Marsh-Russell, '09]

$$\frac{1}{2}(\partial_\mu a)^2 - \frac{1}{2}m_a^2 a^2 + \frac{a}{f_a} \sum_A \frac{g_A^2}{32\pi^2} c_A F^{A\mu\nu} \tilde{F}_{\mu\nu}^A + \frac{\partial_\mu a}{f_a} \left(\sum_\psi c_\psi \psi^\dagger \bar{\sigma}^\mu \psi + \sum_\phi c_\phi \phi^\dagger i \overleftrightarrow{D}^\mu \phi \right)$$

i) approximate shift symmetry $U(1)_{PQ}$ $a(x) \rightarrow a(x) + c$ ($c \in \mathbb{R}$)

: ALP can be naturally light.

ii) periodicity $\frac{a(x)}{f_a} \equiv \frac{a(x)}{f_a} + 2\pi$

: f_a characterizes typical size of ALP couplings

up to dimensionless parameters c_A, c_ψ, c_ϕ .

KSVZ model

Kim '79, Shifman, Vainshtein, Zakharov '80

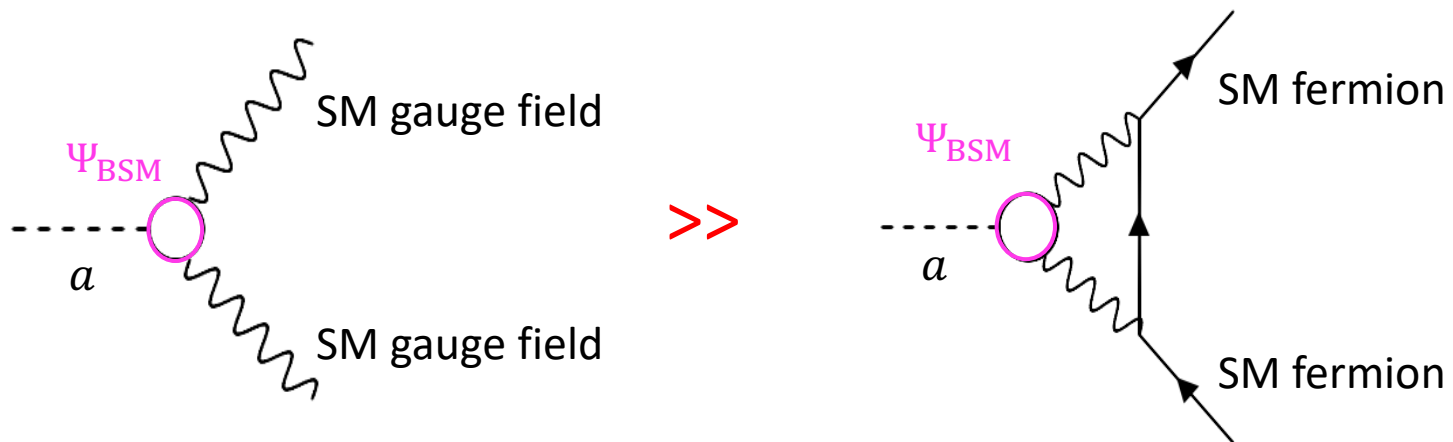
The axion couples to SM fields via a heavy BSM fermion charged under the SM gauge group.

$$y\Phi\Psi_{\text{BSM}}\Psi_{\text{BSM}}^c + \text{h.c.}$$

$$\Phi = \frac{1}{\sqrt{2}}(\rho + f_a)e^{ia/f_a}$$

$$\langle\Phi\rangle = \frac{1}{\sqrt{2}}f_a$$

$$m_\Psi = \frac{y}{\sqrt{2}}f_a$$



“KSVZ-like models”

: no tree-level couplings to the SM fermions

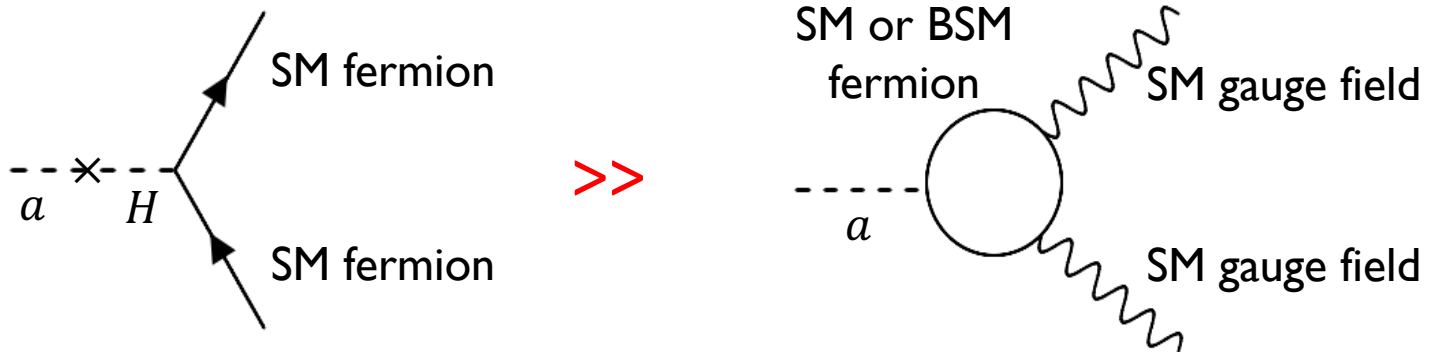
DFSZ model

Dine, Fischler, Srednicki '81, Zhitnitsky '80

The axion couples to the SM sector at tree-level through the Higgs portal.

$$y_u u_R^c Q_L H_u + y_d d_R^c Q_L H_d + y_e e_R^c L H_d + \lambda \Phi^2 H_u H_d + \text{h.c.}$$

$$\Phi = \frac{1}{\sqrt{2}}(\rho + f_a)e^{ia/f_a}$$



“DFSZ-like models”

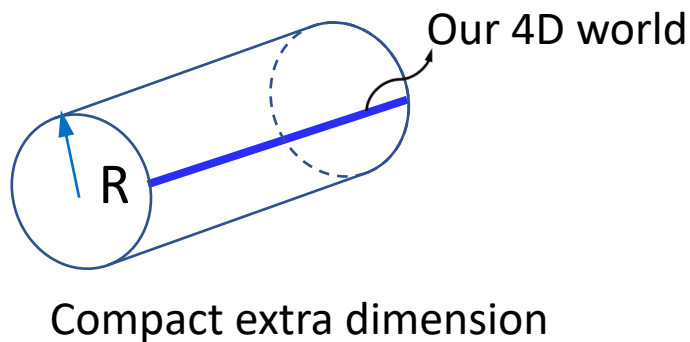
: $O(1)$ tree-level couplings to the SM fermions

String-theoretic models

$$A_{[m_1 m_2 \dots m_p]}(x^\mu, y^m) = a(x^\mu) \Omega_{[m_1 m_2 \dots m_p]}(y^m) \quad \Omega : \text{harmonic } p\text{-form on the compact internal space}$$

4D axions identified as zero modes of higher-dimensional p -form gauge field

- Simplified 5D toy model



$$A_M(x^\mu, y) \quad M = 0, 1, 2, 3, 5$$

$$\Rightarrow A_5(x^\mu, y) = a(x^\mu) \Omega_5(y)$$

$$F_{MN} = \partial_M A_N - \partial_N A_M$$

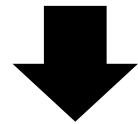
M_5 : 5D Planck mass (String scale)

$$S = \int d^5x \sqrt{-g} \left(-\frac{1}{4g_5^2} F^{MN} F_{MN} + \underbrace{c_1 \frac{\epsilon^{MNPQR}}{\sqrt{-g}} A_M G_{NP}^a G_{QR}^a}_{\text{5D Chern-Simon term (axion-gauge field coupling)}} + \underbrace{\frac{c_2}{M_5} F_{MN} \bar{\Psi} \gamma^M \gamma^N \Psi + \dots}_{\text{5D gauge-matter coupling (axion-matter coupling)}} \right)$$

5D Chern-Simon term
(axion-gauge field coupling)

5D gauge-matter coupling
(axion-matter coupling)

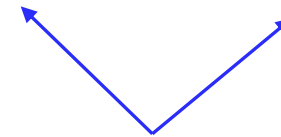
$$c_1 \sim c_2 \sim O(1)$$



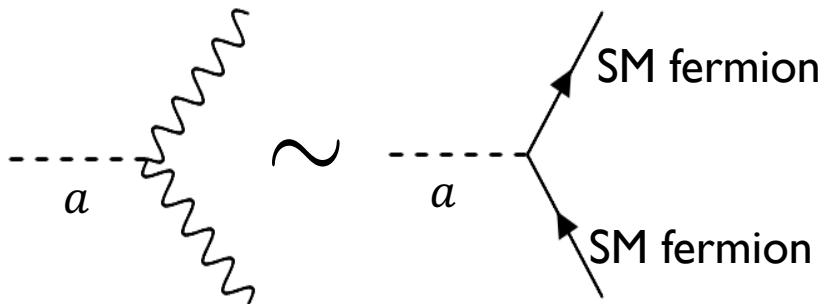
Integrating over the extra dimension

$$S_{\text{eff}} = \int d^4x \left(-\frac{1}{4g^2} F^{\mu\nu} F_{\mu\nu} + \frac{1}{2} \kappa M_5^2 \partial^\mu a \partial_\mu a + c_1 a G^{a\mu\nu} \tilde{G}_{\mu\nu}^a + c_2 \partial_\mu a \bar{\Psi} \gamma^\mu \gamma^5 \Psi + \dots \right)$$

$\kappa \sim O(1)$



String-theoretic axion couplings to matter fields and gauge fields are comparable.



$$S_{\text{eff}} = \int d^4x \left(-\frac{1}{4g^2} F^{\mu\nu} F_{\mu\nu} + \frac{1}{2} \kappa M_5^2 \partial^\mu a \partial_\mu a + c_1 a G^{a\mu\nu} \tilde{G}_{\mu\nu}^a + c_2 \partial_\mu a \bar{\Psi} \gamma^\mu \gamma^5 \Psi + \dots \right)$$

$$\kappa \sim c_1 \sim c_2 \sim O(1)$$



Canonical
normalization

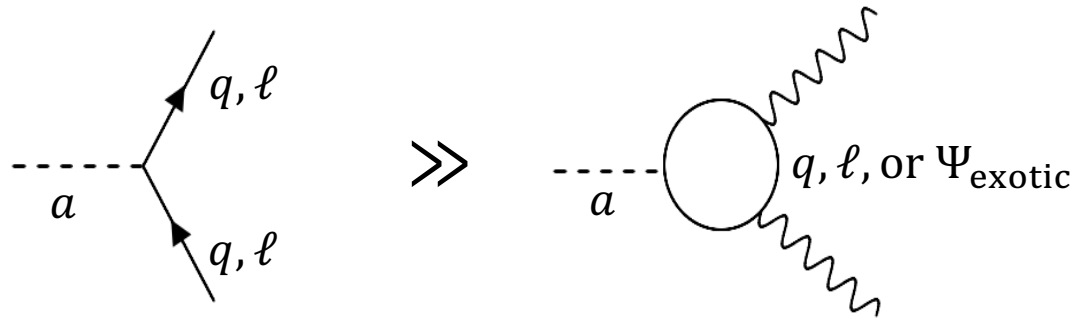
$$a \rightarrow \frac{a}{32\pi^2 f_a} \quad f_a = \frac{\sqrt{\kappa}}{32\pi^2} M_5$$

$$S_{\text{eff}} = \int d^4x \left(-\frac{1}{4g^2} F^{\mu\nu} F_{\mu\nu} + \frac{1}{2} \partial^\mu a \partial_\mu a + \frac{c_1}{32\pi^2} \frac{a}{f_a} G^{a\mu\nu} \tilde{G}_{\mu\nu}^a + \underbrace{\frac{c_2}{32\pi^2} \frac{\partial_\mu a}{f_a} \bar{\Psi} \gamma^\mu \gamma^5 \Psi + \dots}_{\text{Axon-matter couplings}} \right)$$

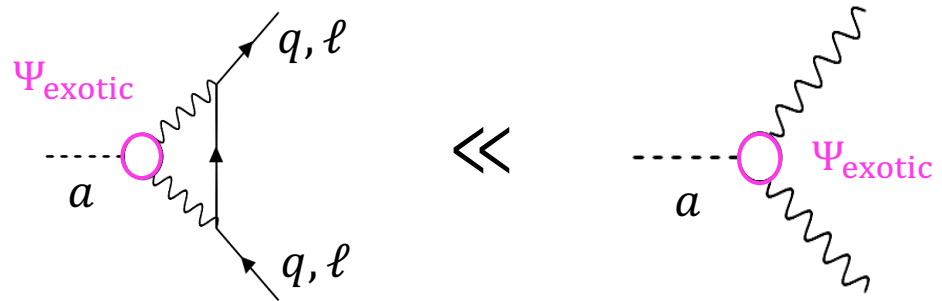
Axon-matter couplings are
suppressed by the 1-loop factor.

Summary: characteristic patterns of axion couplings to the SM depending on the microscopic origins

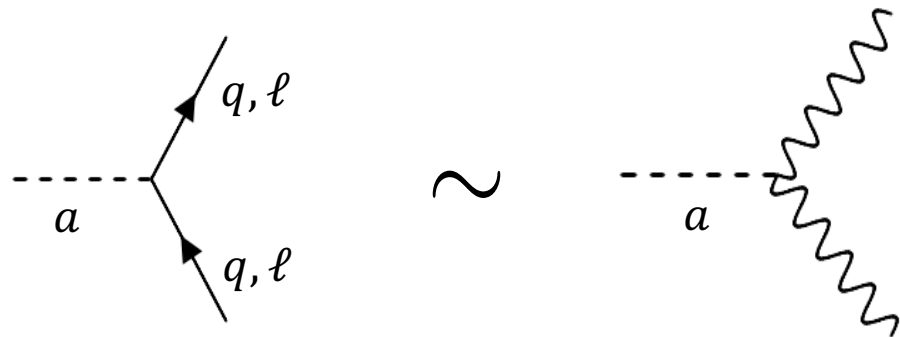
DFSZ-like models



KSVZ-like models

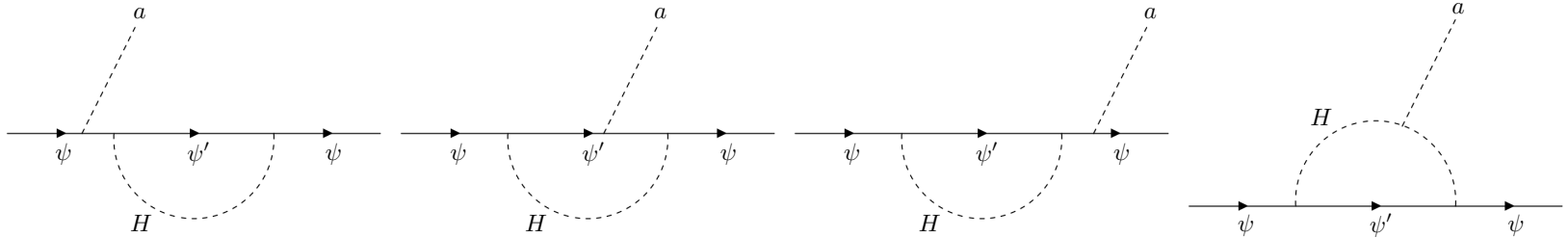


String-theoretic models



Running of axion couplings by Yukawa interactions

K Choi, SHI, CB Park, S Yun '17, Camalich, Pospelov, Vuong, Ziegler, Zupan '20
 Heiles, König, Neubert '20, Chala, Guedes, Ramos, Santiago '20



$$\frac{\partial_\mu a}{f_a} \left(\sum_\psi c_\psi \psi^\dagger \bar{\sigma}^\mu \psi + \sum_{\alpha=1,2} c_{H_\alpha} H_\alpha^\dagger i \overleftrightarrow{D}^\mu H_\alpha \right) + \sum_A \frac{g_A^2}{32\pi^2} c_A \frac{a}{f_a} F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$$y_t u_3^c Q_{L3} H_u$$



$$\frac{dc_{Q_3}}{d \ln \mu} \approx \frac{\xi_y}{16\pi^2} y_t^2 n_t$$

$$\frac{dc_{u_3^c}}{d \ln \mu} \approx \frac{\xi_y}{8\pi^2} y_t^2 n_t$$

$$\frac{dc_{H_u}}{d \ln \mu} \approx \frac{3}{8\pi^2} y_t^2 n_t$$

$$n_t \equiv c_{Q_3} + c_{u_3^c} + c_{H_u}$$

$$\xi_y = \begin{cases} 1 & \text{for non-SUSY models} \\ 2 & \text{for SUSY models} \end{cases}$$

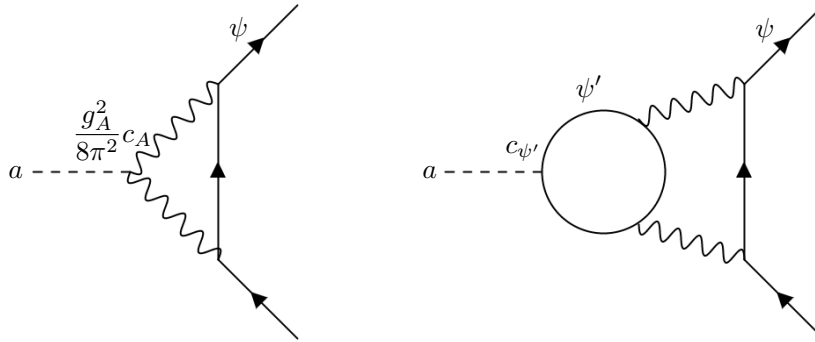
Running of axion couplings by gauge interactions

Srednicki '85, S Chang and K Choi '93

K Choi, SHI, CS Shin '20,

Chala, Guedes, Ramos, Santiago '20

Bauer, Neubert, Renner, Schnubel, Thamm '20



$$\frac{\partial_\mu a}{f_a} \left(\sum_\psi c_\psi \psi^\dagger \bar{\sigma}^\mu \psi + \sum_{\alpha=1,2} c_{H_\alpha} H_\alpha^\dagger i \overleftrightarrow{D}^\mu H_\alpha \right) + \sum_A \frac{g_A^2}{32\pi^2} c_A \frac{a}{f_a} F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$$\left. \frac{dc_\psi}{d \ln \mu} \right|_{\text{gauge}} = -\xi_g \sum_A \frac{3}{2} \left(\frac{g_A^2}{8\pi^2} \right)^2 \mathbb{C}_A(\psi) \tilde{c}_A \quad \tilde{c}_A \equiv c_A - \sum_{\psi'} c_{\psi'}$$

$$\left. \frac{dc_{H_\alpha}}{d \ln \mu} \right|_{\text{gauge}} = -\xi_H \sum_A \frac{3}{2} \left(\frac{g_A^2}{8\pi^2} \right)^2 \mathbb{C}_A(H_\alpha) \tilde{c}_A \quad \mathbb{C}_A(\Phi) : \text{quadratic Casimir}$$

$$\xi_g = \begin{cases} 1 & \text{for non-SUSY models} \\ 2/3 & \text{for SUSY models} \end{cases}, \quad \xi_H = \begin{cases} 0 & \text{for non-SUSY models} \\ 2/3 & \text{for SUSY models} \end{cases}$$

Consequences in low energy observables

Axion couplings to the photon, electron, neutron, and proton below GeV

$$\frac{1}{4}g_{a\gamma}a\vec{E}\cdot\vec{B} + \partial_\mu a \left[\frac{g_{ae}}{2m_e}\bar{e}\gamma^\mu\gamma_5e + \frac{g_{an}}{2m_n}\bar{n}\gamma^\mu\gamma_5n + \frac{g_{ap}}{2m_p}\bar{p}\gamma^\mu\gamma_5p \right]$$

$$g_{a\gamma} \simeq \frac{\alpha_{\text{em}}}{2\pi} \frac{1}{f_a} \left(c_W + c_B - \frac{2}{3} \frac{m_u + 4m_d}{m_u + m_d} c_G \right) \simeq \frac{\alpha_{\text{em}}}{2\pi} \frac{1}{f_a} \left(c_W + c_B - 1.92c_G \right),$$

$$g_{ap} \simeq \frac{m_p}{f_a} \left(C_u\Delta u + C_d\Delta d - \left(\frac{m_d}{m_u + m_d}\Delta u + \frac{m_u}{m_u + m_d}\Delta d \right) c_G \right),$$

$$\simeq \frac{m_p}{f_a} \left(0.90(3) C_u(2 \text{ GeV}) - 0.38(2) C_d(2 \text{ GeV}) - 0.48(3) c_G \right), \quad \underbrace{\langle p | \bar{u}\gamma^\mu\gamma_5u | p \rangle}_{s^\mu \Delta u}$$

$$g_{an} \simeq \frac{m_n}{f_a} \left(C_d\Delta u + C_u\Delta d - \left(\frac{m_u}{m_u + m_d}\Delta u + \frac{m_d}{m_u + m_d}\Delta d \right) c_G \right),$$

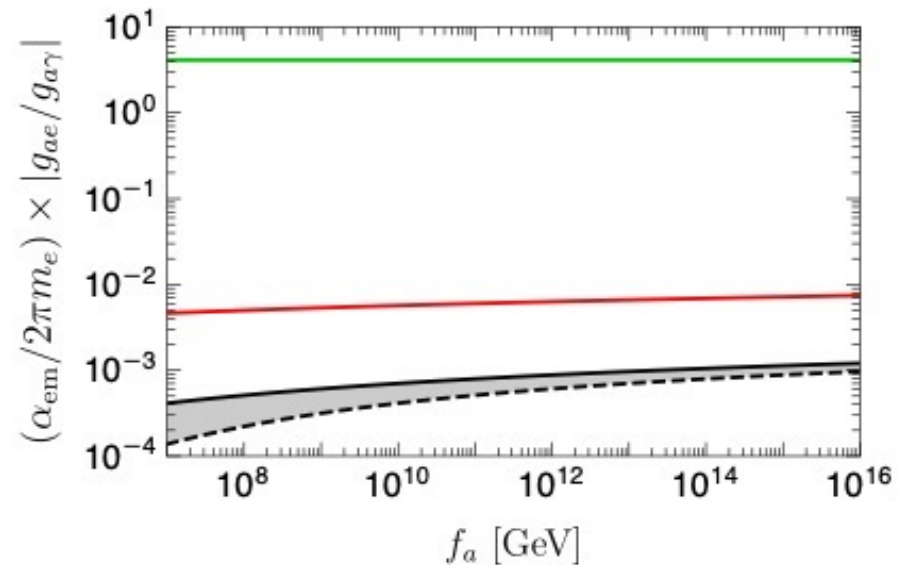
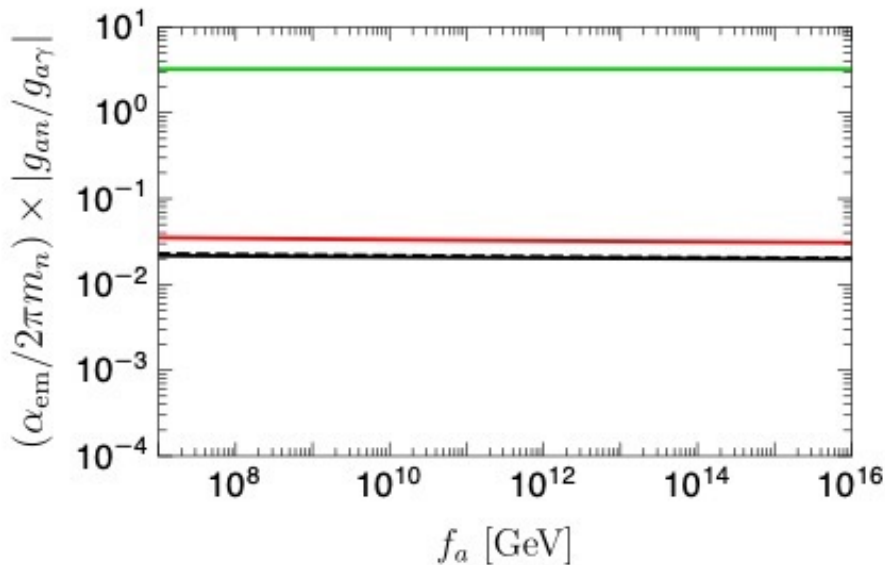
$$\simeq \frac{m_n}{f_a} \left(0.90(3) C_d(2 \text{ GeV}) - 0.38(2) C_u(2 \text{ GeV}) - 0.04(3) c_G \right), \quad \underbrace{\langle p | \bar{d}\gamma^\mu\gamma_5d | p \rangle}_{s^\mu \Delta d}$$

$$g_{ae} \simeq \frac{m_e}{f_a} C_e(m_e),$$

Cortona, Hardy, Vega, Villadoro '15

Distinguishing the models of an axion by coupling ratios

For QCD axion ($c_G \neq 0$), $g_{ap} \sim \frac{m_p}{f_a}$ regardless of the classes of models

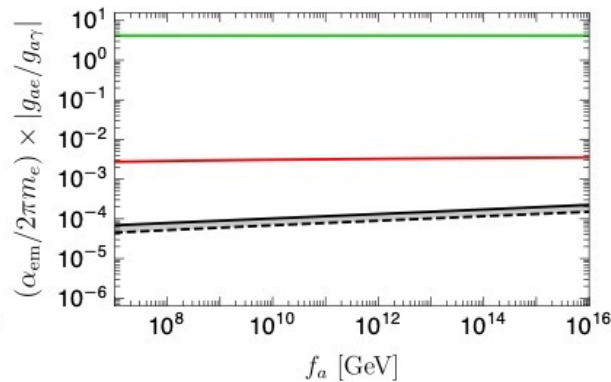
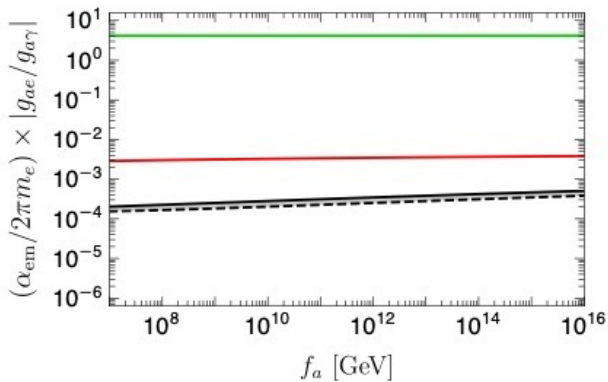
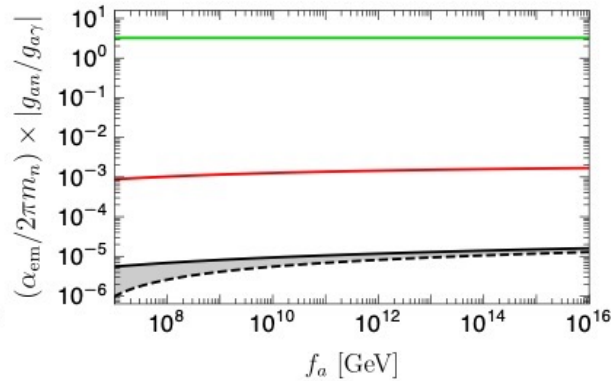
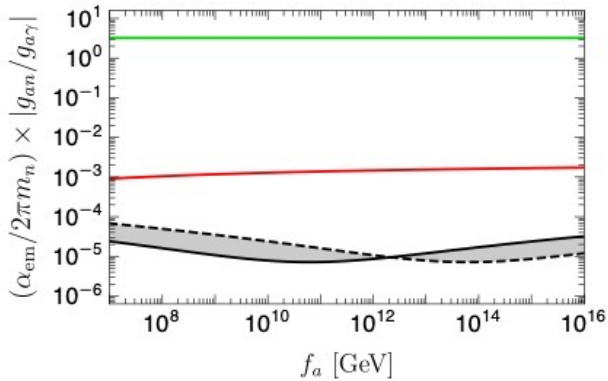
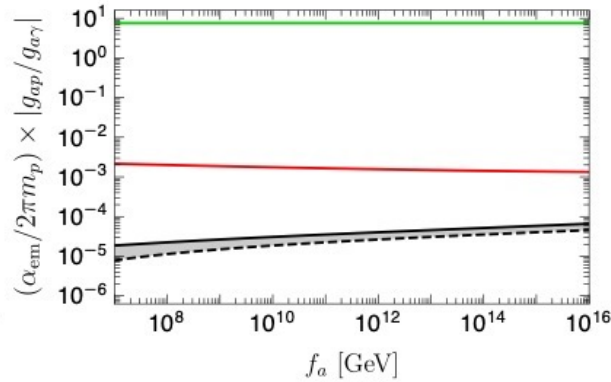
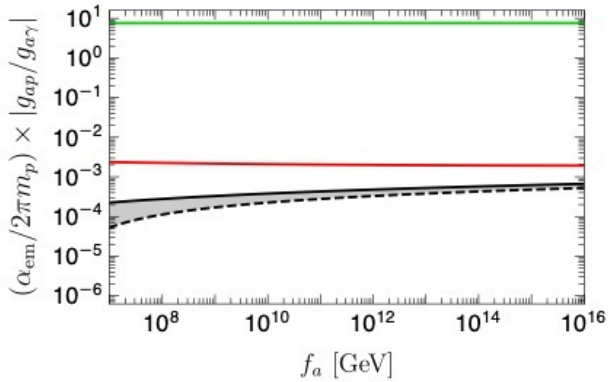


Green : DFSZ-like model

Red : String-theoretic model

Black : KSVZ-like model (dashed : $m_\Psi = 10^{-3}f_a$, solid : $m_\Psi = f_a$)

For ALPs with ($c_G = 0$),



Green : DFSZ-like model
 Red : String-theoretic model
 Black : KSVZ-like model
 (dashed : $m_\Psi = 10^{-3} f_a$,
 solid : $m_\Psi = f_a$)

$$c_W = 1 \quad (c_G = c_B = 0)$$

$$c_B = 1 \quad (c_G = c_W = 0)$$

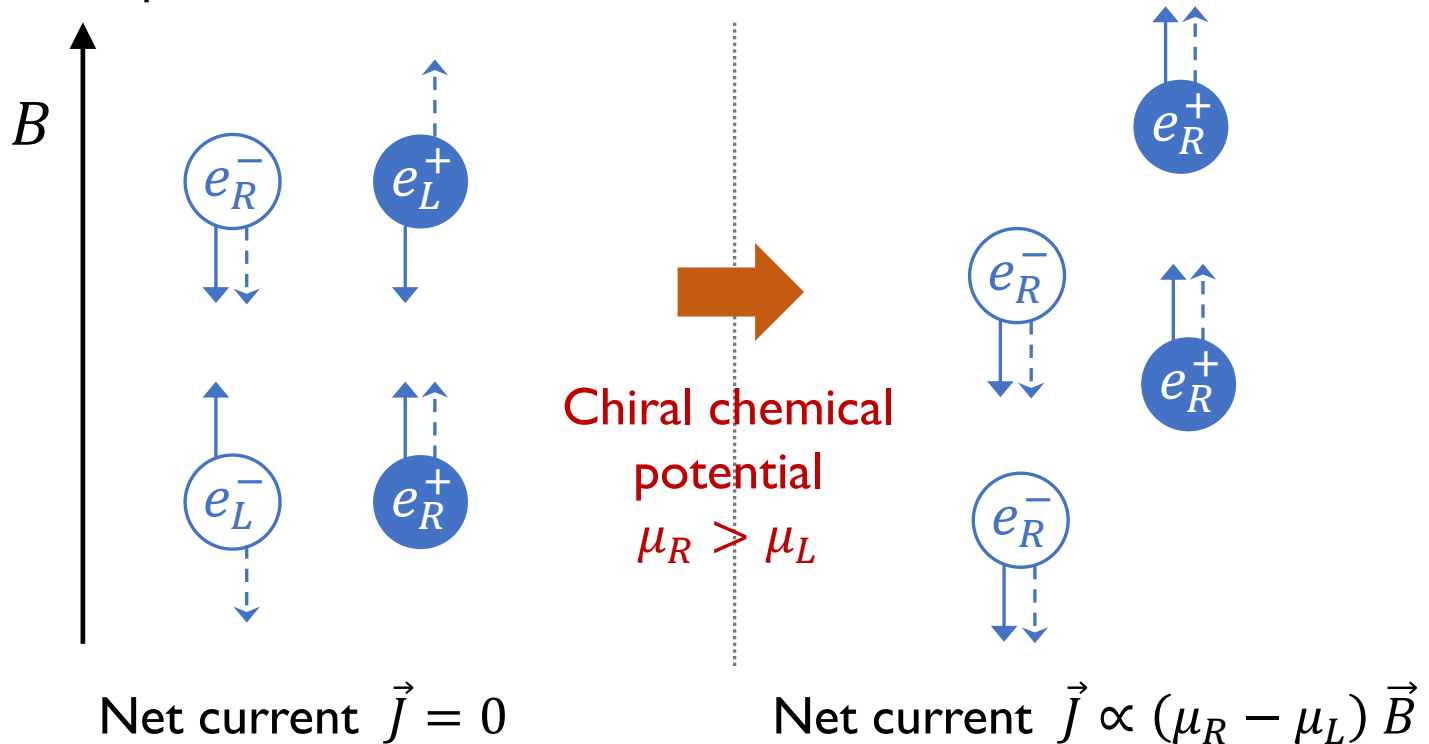
Take-home message I

- In principle, we have 3 possible classes of UV physics for an axion : KSVZ-like, DFSZ-like, and string-theoretic.
- Those 3 classes of UV physics may be experimentally distinguishable by measuring the ratio of an axion-fermion coupling to the axion-photon coupling.
- For the QCD axion, *the measurement of the axion-electron coupling is crucial* for the distinction.

Chiral Magnetic Effect (CME) in a nutshell

— Momentum direction

- - - Spin direction



- The magnetic field aligns the spin directions depending on particles and antiparticles.
- The helicity imbalance causes a non-zero electric current along the B-field direction.

- Chiral chemical potential

$$\begin{aligned}\mathcal{L} \supset \mu_5(n_R - n_L) &= \mu_5 \left(\psi_R^\dagger \psi_R - \psi_L^\dagger \psi_L \right) \\ &= \mu_5 \bar{\Psi} \gamma^0 \gamma^5 \Psi\end{aligned}\quad \Psi = \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix}$$

It makes *helicity imbalance*.

- (Vector) chemical potential

$$\begin{aligned}\mathcal{L} \supset \mu(n_R + n_L) &= \mu \left(\psi_R^\dagger \psi_R + \psi_L^\dagger \psi_L \right) \\ &= \mu \bar{\Psi} \gamma^0 \Psi\end{aligned}$$

It makes *charge imbalance* (i.e. particles vs antiparticles).

The *charge imbalance alone cannot induce a current*. However, μ might be still relevant for the magnitude of the current for a given helicity imbalance.

Time-varying axion field as a source for μ_5

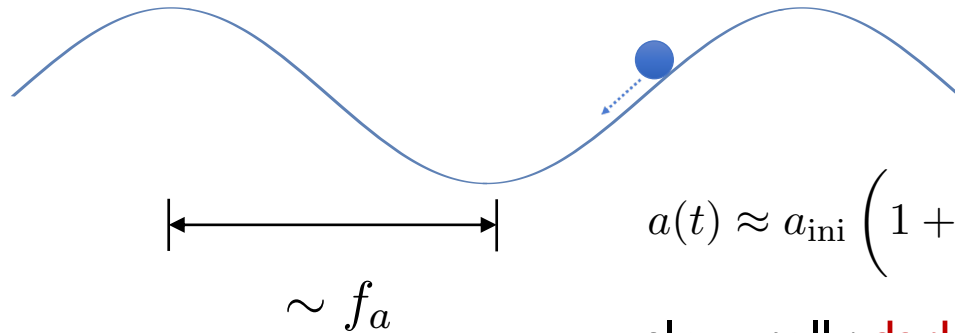
Suppose that we have an axion-fermion coupling:

$$c_\Psi \frac{\partial_\mu a}{f_a} \bar{\Psi} \gamma^\mu \gamma^5 \Psi \quad \longrightarrow \quad c_\Psi \frac{\dot{a}}{f_a} \bar{\Psi} \gamma^0 \gamma^5 \Psi$$

$$\mu_5(t) = c_\Psi \frac{\dot{a}}{f_a}$$

Cosmological evolution of an axion field

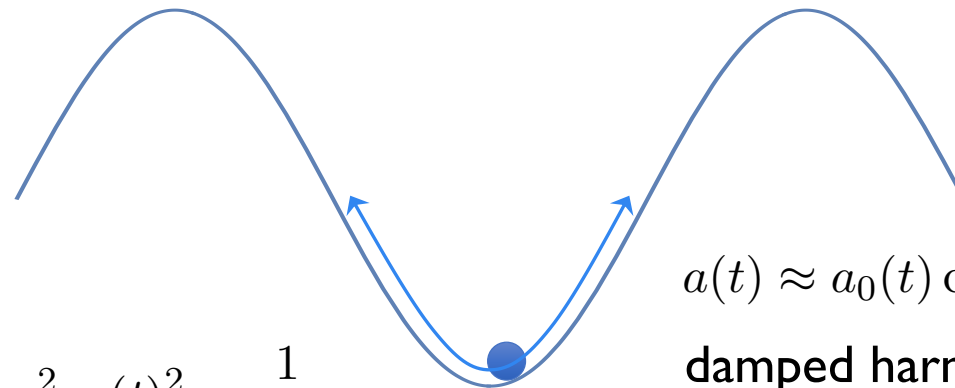
$$m_a \ll H(t)$$



$$a(t) \approx a_{\text{ini}} \left(1 + \frac{1}{20} \left[\frac{m_a^2}{H_{\text{ini}}^2} - \frac{m_a^2}{H(t)^2} \right] \right)$$

slow-roll : **dark energy**

$$m_a \gg H(t)$$

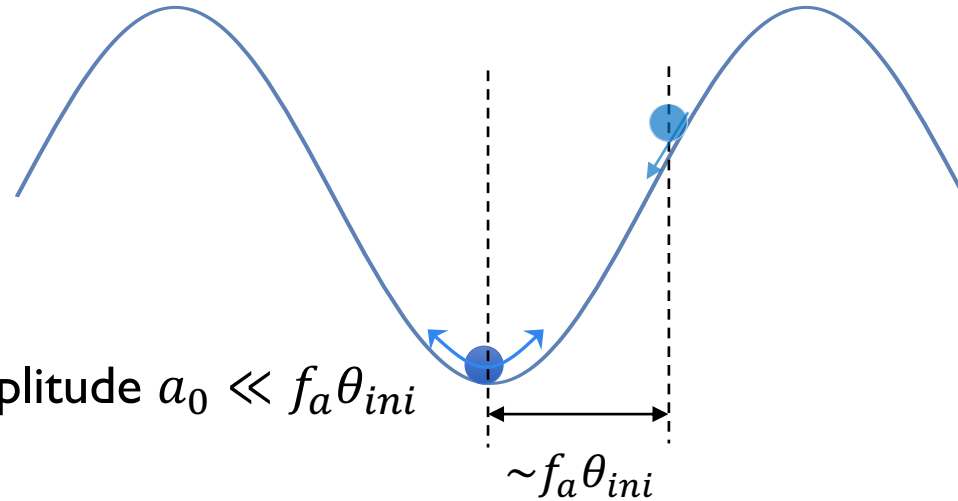


$$\rho_a \approx \frac{1}{2} m_a^2 a_0(t)^2 \propto \frac{1}{R^3}$$

$$a(t) \approx a_0(t) \cos(m_a t)$$

damped harmonic
oscillator : **cold dark matter**

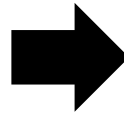
Misalignment production of axion dark matter and chiral chemical potential



Present oscillation amplitude $a_0 \ll f_a \theta_{ini}$

$$a(t, \vec{x}) \approx a_0 \cos m_a t$$

$$\rho_a = \frac{1}{2} m_a^2 a_0^2$$



$$\mu_5(t) \approx -c_\Psi \frac{\sqrt{2\rho_a}}{f_a} \sin m_a t$$

$$\mu_5(t) = c_\Psi \frac{\dot{a}}{f_a}$$

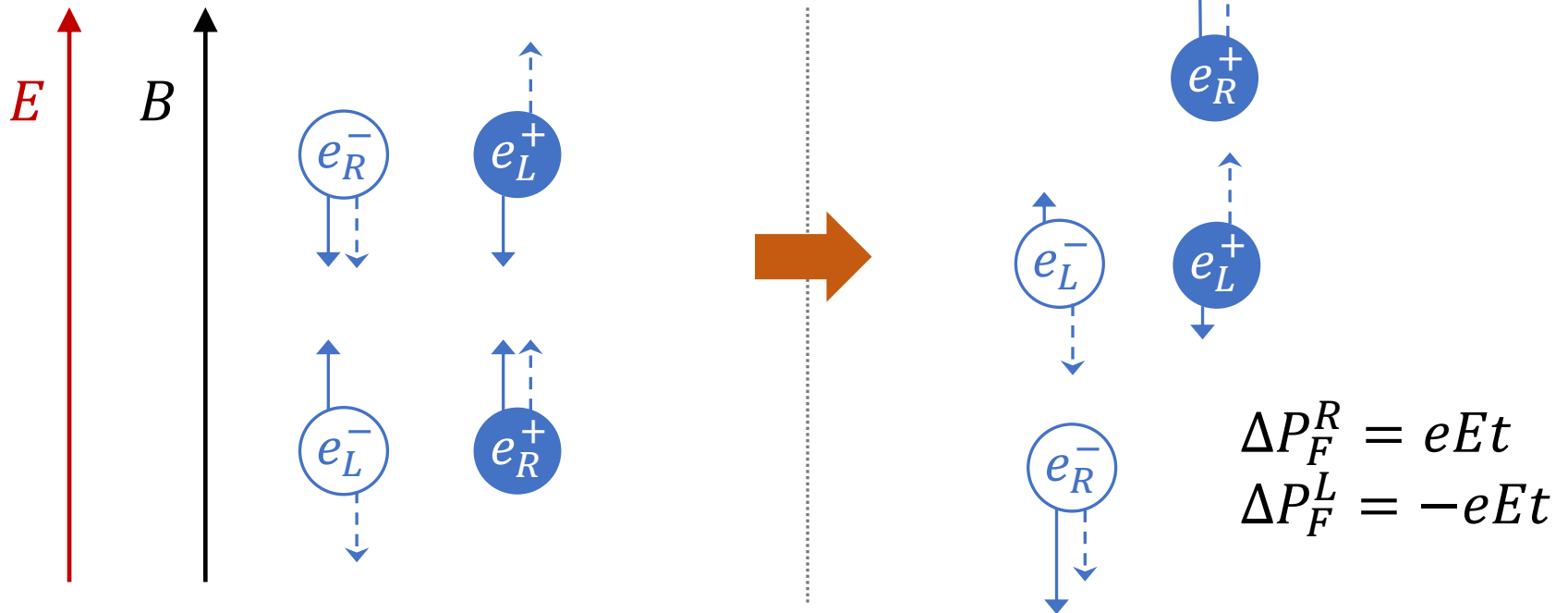
The axion dark matter background gives rise to an oscillating chiral chemical potential for fermions coupled to the axion.

CME-induced current

I. Energy balance argument

Nielsen and Ninomiya '83
Fukushima, Kharzeev, Warringa '08

— Momentum direction
- - - Spin direction



Density change of
right-handed fermion states

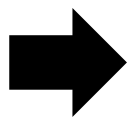
$$\underbrace{\frac{\Delta P_F^R}{2\pi}} \cdot \underbrace{\frac{eB}{2\pi}} = \frac{e^2}{4\pi^2} \vec{E} \cdot \vec{B} t$$

Longitudinal phase
space density

The density of Landau levels in
the transverse direction

Density change of
left-handed fermion states

$$\frac{\Delta P_F^L}{2\pi} \cdot \frac{eB}{2\pi} = -\frac{e^2}{4\pi^2} \vec{E} \cdot \vec{B} t$$



$$\frac{d}{dt}(n_R - n_L) = \frac{e^2}{2\pi^2} \vec{E} \cdot \vec{B}$$

which partially reproduces the axial anomaly equation

$$\partial_\mu(\bar{\Psi} \gamma^\mu \gamma^5 \Psi) = \frac{e^2}{2\pi^2} \vec{E} \cdot \vec{B} + 2m\bar{\Psi}\Psi$$

$$\frac{d}{dt}(n_R - n_L) = \frac{e^2}{2\pi^2} \vec{E} \cdot \vec{B}$$

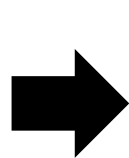
$$\blackrightarrow \underbrace{\mathcal{E}_{LR} \frac{d}{dt}(N_R - N_L)} = \frac{e^2}{2\pi^2} \mathcal{E}_{LR} \int d^3x \vec{E} \cdot \vec{B}$$

Energy cost per unit time needed for adding a right-handed fermion and removing a left-handed fermion

||

$$\int d^3x \vec{j} \cdot \vec{E}$$

This energy cost has to be supplied by the electric power.



$$\boxed{\vec{j} = \frac{e^2}{2\pi^2} \mathcal{E}_{LR} \vec{B}}$$

\mathcal{E}_{LR} : Energy cost for converting a left mode to a right mode

Essentially

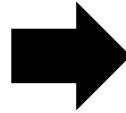
$$\begin{aligned} j &= \rho v = e(n_R - n_L) v_F \\ &= e \left(\frac{(p_F^R - p_F^L)}{2\pi} \frac{eB}{2\pi} \right) v_F \\ &= \frac{e^2}{2\pi^2} \cdot \underbrace{\frac{1}{2}(p_F^R - p_F^L) v_F}_{\mathcal{E}_{LR}} \cdot B \end{aligned}$$

Landau level
quantization

Fermi surface ($\omega = \mu$)

$$\omega_R = \sqrt{(|p_z| - \mu_5)^2 + m^2}$$

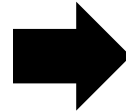
$$\omega_L = \sqrt{(|p_z| + \mu_5)^2 + m^2}$$



$$p_F^R = \sqrt{\mu^2 - m^2} + \mu_5$$

$$p_F^L = \sqrt{\mu^2 - m^2} - \mu_5$$

$$\mathcal{E}_{LR} = \frac{1}{2}(p_F^R - p_F^L)v_F$$



$$\mathcal{E}_{LR} = \mu_5 v_F$$

Therefore,

$$\vec{j} = \frac{e^2}{2\pi^2} \mu_5 v_F \vec{B}$$

(Caveat) v_F dependence is missing in the original literature of Fukushima, Kharzeev, Warringa (0808.3382).

CME-induced current

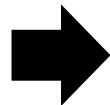
II. Effective lagrangian approach for non-relativistic fermions

DK Hong, SHI, KS Jeong, D Yeom '22

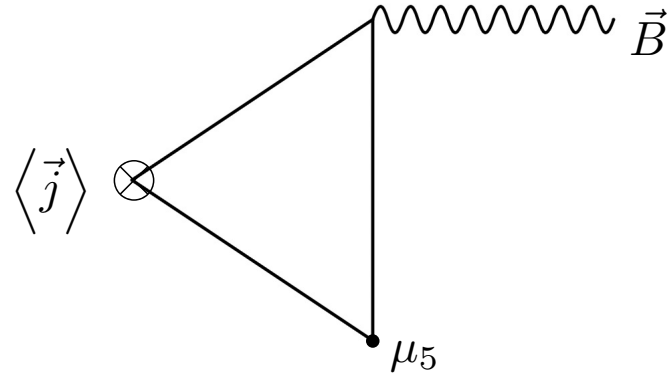
$$\mathcal{L} = i\bar{\Psi}\gamma^\mu D_\mu\Psi - m\bar{\Psi}\Psi + \mu_5\bar{\Psi}\gamma^0\gamma^5\Psi$$

$$\Psi = e^{-imt}\tilde{\Psi} \quad \tilde{\Psi} = \begin{pmatrix} \psi \\ \chi \end{pmatrix}$$

With the standard rep of the γ matrices, χ can be integrated out for the momentum scale below m .


$$\mathcal{L}_{\text{eff}} = \psi^\dagger \left[i\partial_t + \frac{1}{2m}(\vec{\sigma} \cdot \vec{D})^2 + \mu_5 \frac{(-i\vec{\nabla})}{m} \cdot \vec{\sigma} \right] \psi$$

(Caveat) This part was not noticed in our draft v1.
(to be corrected : thanks to K. Choi)



$$\vec{j}_L(\mu_L, B) \equiv \left\langle \frac{\delta S}{\delta \vec{A}} \right\rangle_{\vec{A}=0} = e \langle \Psi_L^\dagger \vec{\sigma} v_F \Psi_L \rangle = -ie \int_p \text{Tr}_L \left(\frac{\sigma^i v_F}{(1+i\epsilon)p^0 - \frac{p^2}{2m} + \mu_B \vec{\sigma} \cdot \vec{B}_{\text{ext}} + \mu_L} \right)$$

$$\begin{aligned} \blacktriangleright \quad j_L^i(\mu_L, B) &= \frac{e^2}{2\pi^2} \mu_L B \delta^{i3} \\ j_R^i(\mu_R, B) &= -\frac{e^2}{2\pi^2} \mu_R B \delta^{i3} \end{aligned}$$

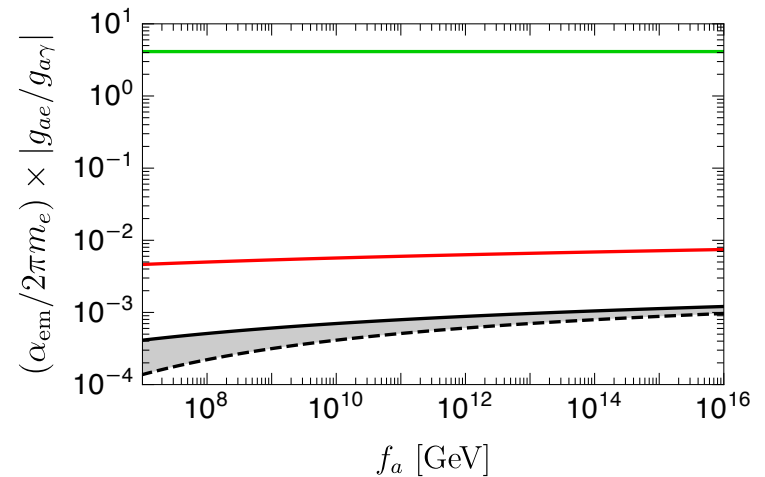
$$\blacktriangleright \quad \boxed{j^3 = j_L^3 + j_R^3 = \frac{e^2}{2\pi^2} (\mu_L - \mu_R) B}$$

$$\mu_L - \mu_R = \mu_5 v_F \neq \mu_5 \quad (\text{to be corrected})$$

The axion-electron coupling

$$\mathcal{L}_{\text{int}} = C_e \frac{\partial_\mu a}{f} \bar{\psi} \gamma^\mu \gamma_5 \psi$$

$$C_e \simeq \begin{cases} \mathcal{O}(1) & \text{DFSZ-like models} \\ \mathcal{O}(10^{-4} \sim 10^{-3}) & \text{KSVZ-like models} \\ \mathcal{O}(10^{-3} \sim 10^{-2}) & \text{string-theoretic axions} \end{cases}$$



The measurement of the axion-electron coupling will give us an important clue for underlying high energy physics.

K Choi, SHI, HJ Kim, H Seong '21

Detecting axion dark matter via the CME

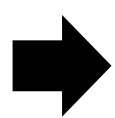
DK Hong, SHI, KS Jeong, D Yeom '22

$$C_e \frac{\partial_\mu a}{f} \bar{\psi} \gamma^\mu \gamma_5 \psi \quad \longrightarrow \quad \mu_5 = C_e \frac{\dot{a}}{f}$$

$$a(t) = \frac{\sqrt{2\rho_{\text{DM}}}}{m_a} \sin(m_a t)$$

The axion dark matter field induces an oscillating chiral chemical potential if the axion couples to the electrons.

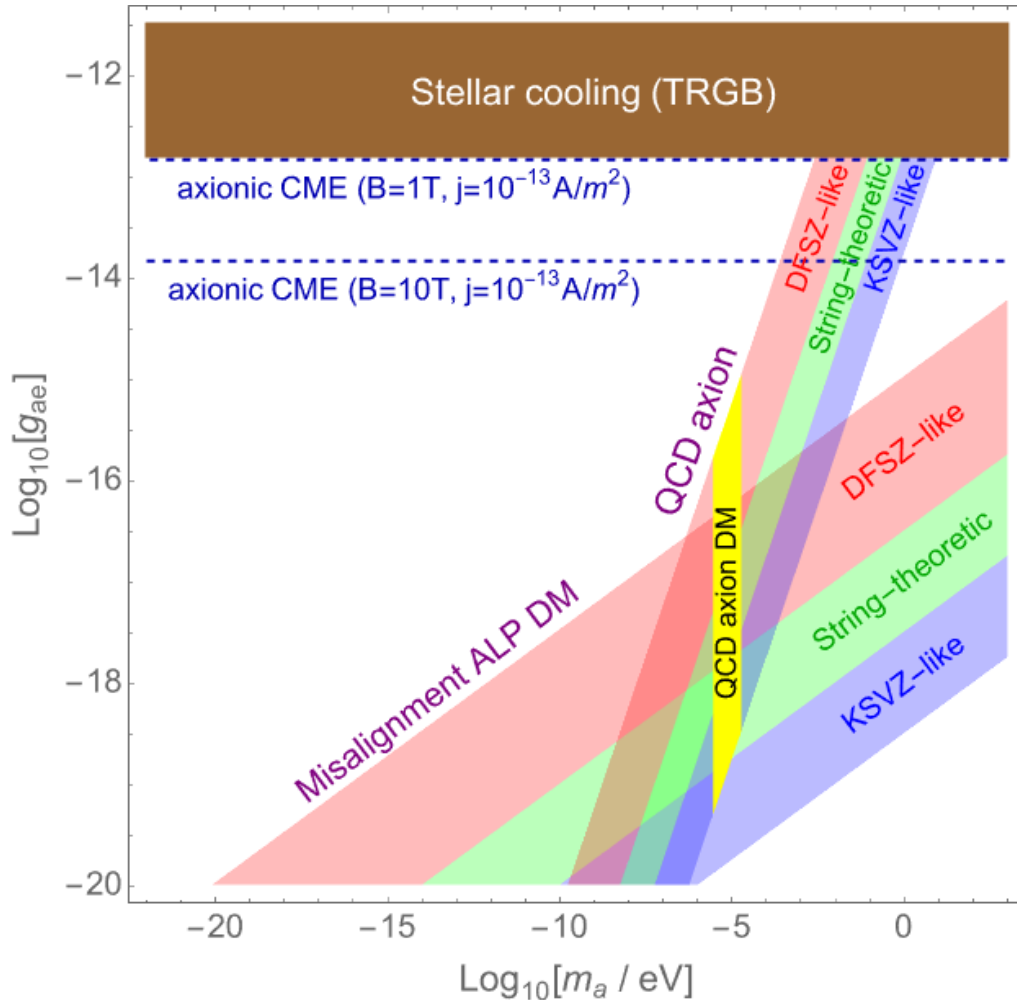
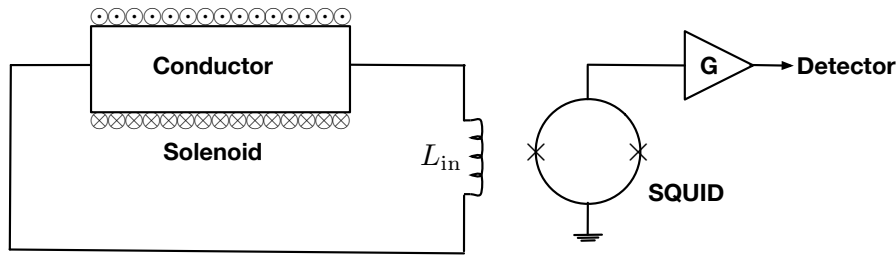
$$\mu_5 = C_e \frac{\sqrt{\rho_{\text{DM}}}}{f} \cos(m_a t) \sim 0.25 \times 10^{-32} \text{ GeV} \cdot \left(\frac{\rho_{\text{DM}}}{0.4 \text{ GeV cm}^{-3}} \right)^{1/2} \cdot \left(\frac{10^{12} \text{ GeV}}{f/C_e} \right)$$



$$j^3 = \frac{e^2}{2\pi^2} \mu_5 v_F B = 6.8 \times 10^{-15} \text{ A m}^{-2} \cos(m_a t)$$

CME

$$\times \left(\frac{v_F}{10^{-2}} \right) \cdot \left(\frac{\rho_{\text{DM}}}{0.4 \text{ GeV cm}^{-3}} \right)^{1/2} \cdot \left(\frac{10^{12} \text{ GeV}}{f/C_e} \right) \cdot \left(\frac{B}{10 \text{ Tesla}} \right)$$



Such a tiny CME-induced current may be measurable by exploiting a highly sensitive superconducting SQUID coil or a semiconductor diode.

Conclusions

- Axions are theoretically well-motivated new particles which may be an important clue for underlying UV physics when they are discovered.
- The underlying UV physics may be distinguishable by precision measurements of low energy axion couplings.
- The measurement of the axion-electron coupling is particularly important for pinning down the microscopic origin of the QCD axion.
- The chiral magnetic effect (CME) offers an intriguing possibility for the measurement of the axion-electron coupling, when the axion comprises a major fraction of dark matter.
- At the moment, the size of the induced current seems rather controversial for the case of massive fermions.

Back-up slides

String-theoretic models

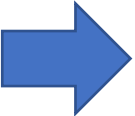
$$C_{[m_1 m_2 \dots m_p]}(x^\mu, y^m) = a(x^\mu) \Omega_{[m_1 m_2 \dots m_p]}(y^m) \quad \Omega : \text{harmonic } p\text{-form on the compact internal space}$$

4D axions identified as zero modes of higher-dimensional p -form gauge field

 SUSY-preserving compactification

$$\left\{ \begin{array}{l} T = \tau + ia \quad \text{Axion chiral superfield } (\tau : \text{volume modulus of } p\text{-cycle dual to } \Omega) \\ U(1)_{PQ} : a \rightarrow a + \text{const} \\ \quad \quad \quad : \text{remnant of a higher-dimensional gauge symmetry} \end{array} \right.$$

$$\delta C_{[m_1 m_2 \dots m_p]} = \partial_{[m_1} \Lambda_{m_2, \dots, m_p]}$$

 4D Low energy effective action

$$K = K_0(T + T^*) + Z_I(T + T^*)\Phi_I^*\Phi_I$$

$$\mathcal{F}_A = c_A T \quad Z_I \propto (T + T^*)^{\omega_I} \quad \omega_I \sim \mathcal{O}(1) \quad \text{scaling weight of } \Phi_I$$

$$c_A \sim \mathcal{O}(1)$$

$$K = K_0(T + T^*) + Z_I(T + T^*)\Phi_I^*\Phi_I$$

$$\mathcal{F}_A = c_A T \quad Z_I \propto (T + T^*)^{\omega_I} \quad \omega_I \sim \mathcal{O}(1) \quad c_A \sim \mathcal{O}(1)$$



$$T = \tau + ia$$

$$\mathcal{L}_{\text{eff}} = \frac{M_P^2}{4} \partial_\tau^2 K_0 (\partial_\mu a)^2 + \frac{\omega_I}{2\tau} \partial_\mu a \left(\psi_I^\dagger \bar{\sigma}^\mu \psi_I + \phi_I^\dagger i \overleftrightarrow{D}^\mu \phi_I \right) - \frac{1}{4} \frac{c_A \tau}{c_A g_A^2} F^{A\mu\nu} F_{\mu\nu}^A + \frac{c_A}{4} a F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$\sim \mathcal{O}(1)$ $\sim \mathcal{O}(1)$

$\tau = \frac{1}{c_A g_A^2} \sim \mathcal{O}(1)$

String-theoretic axion couplings to matter fields and gauge fields are comparable to each other.



Canonical normalization

$$a \rightarrow \frac{a}{8\pi^2 f_a} \quad f_a = \frac{M_P}{8\pi^2} \sqrt{\frac{\partial_\tau^2 K_0}{2}}$$

$$\mathcal{L}_{\text{eff}} = -\frac{1}{4g_A^2} F^{A\mu\nu} F_{\mu\nu}^A + \frac{1}{2} (\partial_\mu a)^2 + \omega_I \frac{c_A g_A^2}{16\pi^2} \frac{\partial_\mu a}{f_a} \left(\psi_I^\dagger \bar{\sigma}^\mu \psi_I + \phi_I^\dagger i \overleftrightarrow{D}^\mu \phi_I \right) + \frac{c_A}{32\pi^2} \frac{a}{f_a} F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$\sim \mathcal{O}(g^2/16\pi^2)$

Laboratory searches for axion DM -photonic probes

$$\frac{g_{a\gamma}}{4} a F \tilde{F} \quad \longrightarrow \quad \nabla \times \vec{B} = \frac{\partial \vec{E}}{\partial t} \underbrace{-g_{a\gamma} \vec{B} \partial_t a}_{\vec{J}_{\text{eff}}} \quad \text{effective current}$$

Background axion DM field

$$a \approx a_0 \cos [m_a (t - \vec{v} \cdot \vec{x})]$$

$$\rho_a = \frac{1}{2} m_a^2 a_0^2 \quad |\vec{v}| \sim 10^{-3} c$$

$$\vec{J}_{\text{eff}} \approx g_{a\gamma} \sqrt{2\rho_a} \vec{B} \sin m_a t$$

The best experimental sensitivity on $g_{a\gamma}$ is obtained when $\rho_a = \rho_{\text{DM}}$.

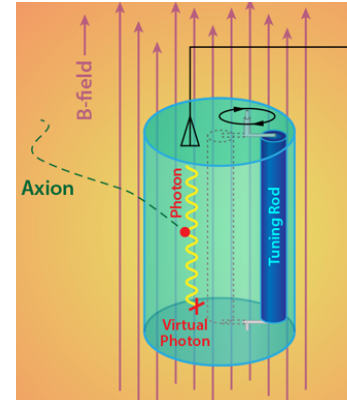
Misalignment axion DM

$$f_a \simeq 10^{17} \text{ GeV} \left(\frac{10^{-22} \text{ eV}}{m_a} \right)^{1/4} \sqrt{\frac{\rho_a}{\rho_{\text{DM}}}} \quad \longrightarrow \quad g_{a\gamma} = \frac{e^2}{8\pi^2} \frac{1}{f_a} c_{a\gamma}$$

Given axion DM mass,
 $g_{a\gamma}$ is determined for $c_{a\gamma} \sim O(1)$.

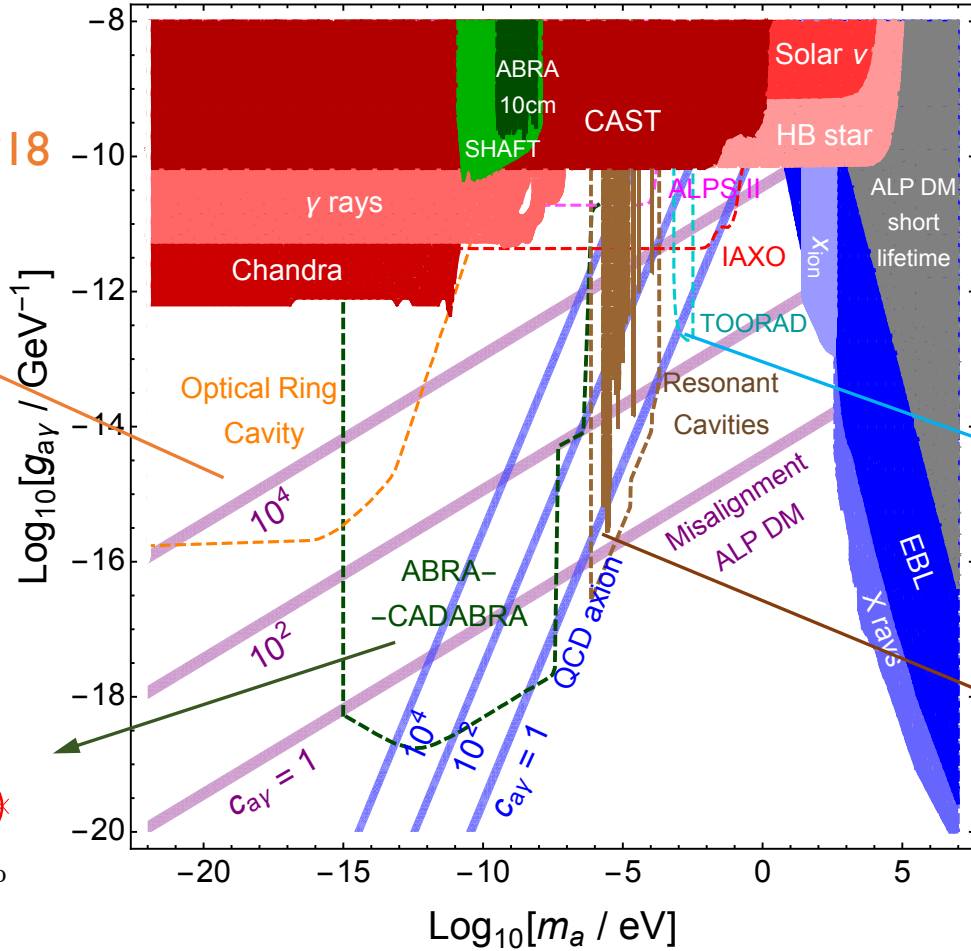
Current and future limits on $g_{a\gamma}$

Choi, SHI, Shin '20

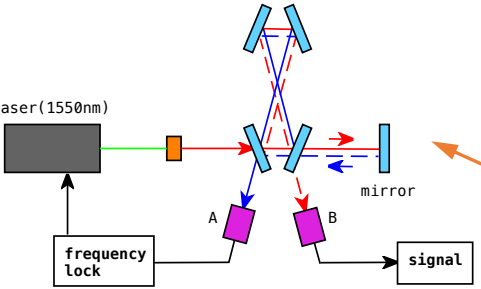


Marsh, Fong, Lenz, Smejkal, Ali '18

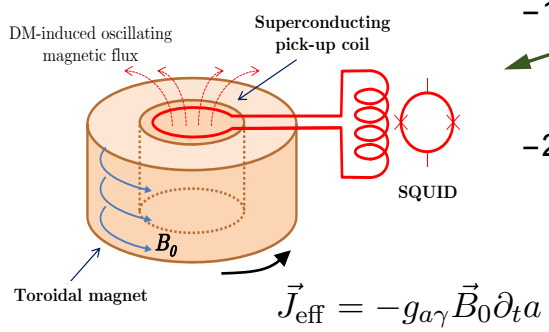
ADMX, IBS-CAPP, MADMAX...



Obata, Fujita, Michimura '18



Kahn, Safdi, Thaler '16



Laboratory searches for axion DM -nucleonic probes

$$g_{aN} \frac{\partial_\mu a}{2m_N} \bar{N} \gamma^\mu \gamma^5 N \quad \longrightarrow \quad \underbrace{g_{aN} \frac{\nabla a}{\gamma_N m_N}}_{\vec{B}_{\text{eff}}} \cdot \gamma_N \vec{S}_N \quad \gamma_N : \text{nucleon gyromagnetic ratio}$$

Background axion DM field

$$a \approx a_0 \cos [m_a (t - \vec{v} \cdot \vec{x})]$$

$$\rho_a = \frac{1}{2} m_a^2 a_0^2 \quad |\vec{v}| \sim 10^{-3} c$$

$$\vec{B}_{\text{eff}} \approx g_{aN} \frac{\sqrt{2\rho_a}}{\gamma_N m_N} \vec{v}_a \sin m_a t$$

The best experimental sensitivity on g_{aN} is obtained when $\rho_a = \rho_{DM}$.

Misalignment axion DM

$$f_a \simeq 10^{17} \text{ GeV} \left(\frac{10^{-22} \text{ eV}}{m_a} \right)^{1/4} \sqrt{\frac{\rho_a}{\rho_{DM}}} \quad \longrightarrow \quad g_{aN} = \frac{m_N}{f_a} c_{aq} \times \mathcal{O}(1)$$

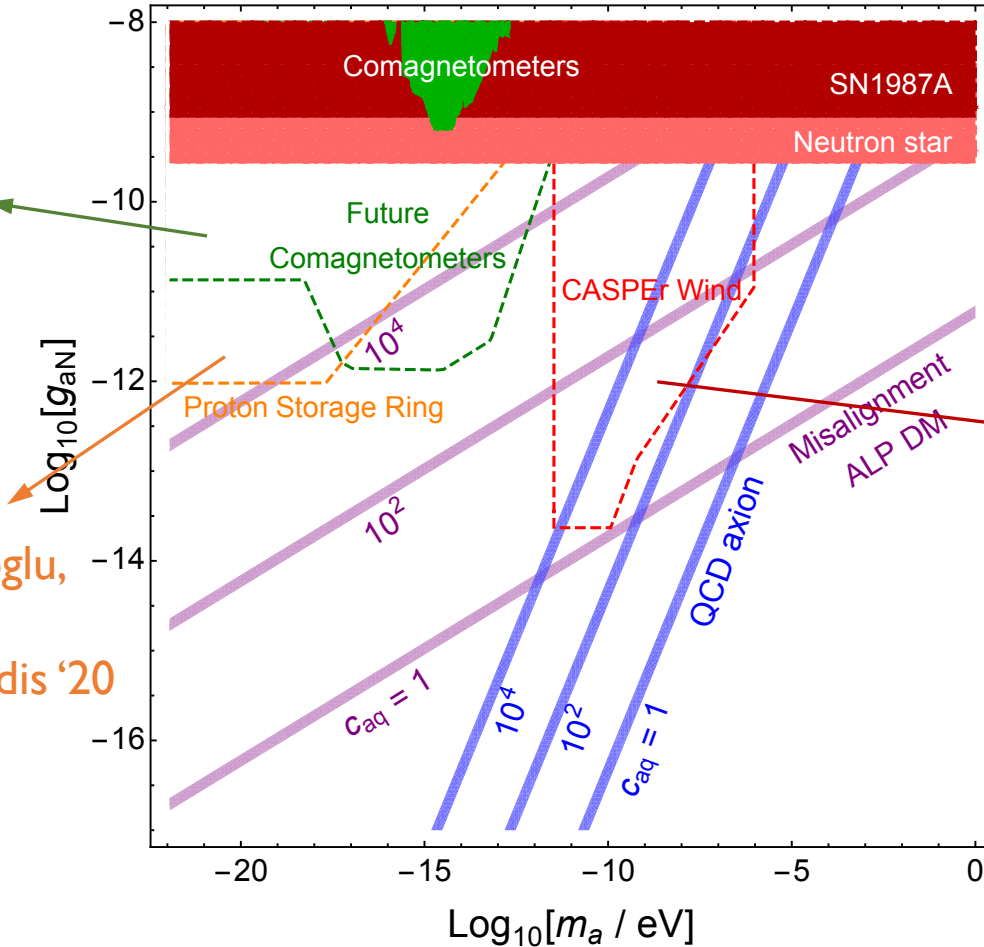
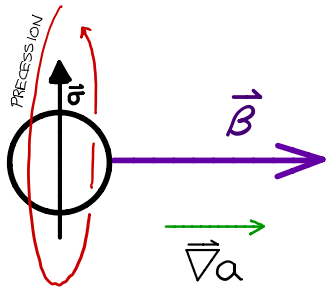
Given axion DM mass,
 g_{aN} is determined for $c_{aq} \sim \mathcal{O}(1)$.

Current and future limits on g_{aN}

Bloch, Hochberg,
Kuflik, Volansky '19

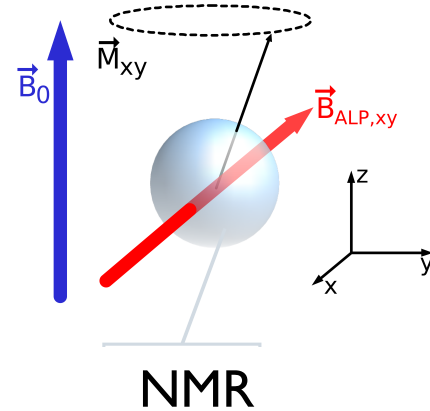
$$\frac{B_{\text{eff}}^e}{B_{\text{eff}}^N} \sim \frac{c_{ae}m_e}{c_{aN}m_N} \neq 1$$

Graham, Hacıomeroglu,
Kaplan, Omarov,
Rajendran, Semertzidis '20



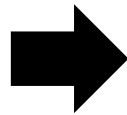
Choi, Shi, Shin '20

Kimball et al '17



Dirac equation for a massless fermion

$$\begin{pmatrix} 0 & \omega - p_z \sigma^3 + \mu_5 \\ \omega + p_z \sigma^3 - \mu_5 & 0 \end{pmatrix} \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix} = 0$$

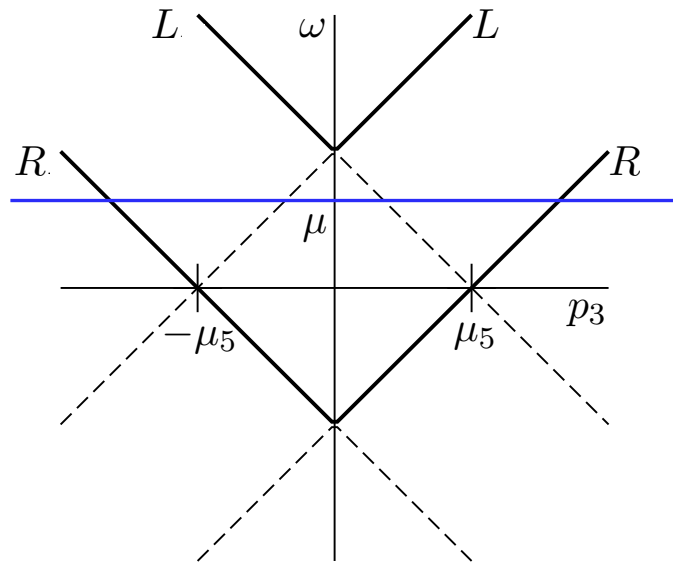


$$\omega_R = |p_z| - \mu_5$$

$$\omega_L = |p_z| + \mu_5$$

Fermi surface

$$n_i = \frac{1}{e^{\beta(\omega_i - \mu)} + 1}$$

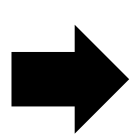


At low temperatures and low μ , only the right chiral modes are occupied for a positive μ_5 .

→ helicity imbalance

Dirac equation for a massive fermion

$$\begin{pmatrix} -m & \omega - p_z \sigma^3 + \mu_5 \\ \omega + p_z \sigma^3 - \mu_5 & -m \end{pmatrix} \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix} = 0$$

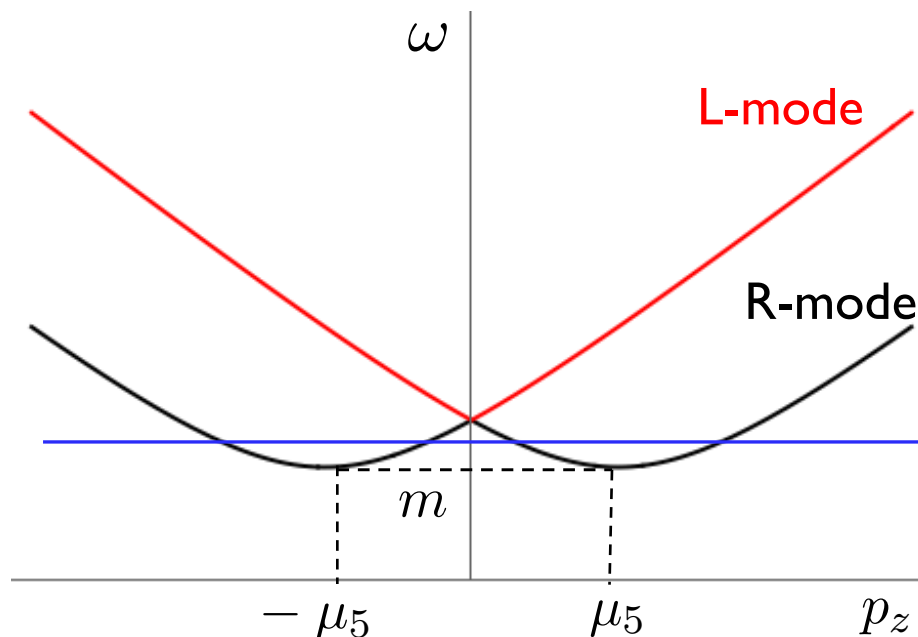


$$\omega_R = \sqrt{(|p_z| - \mu_5)^2 + m^2}$$

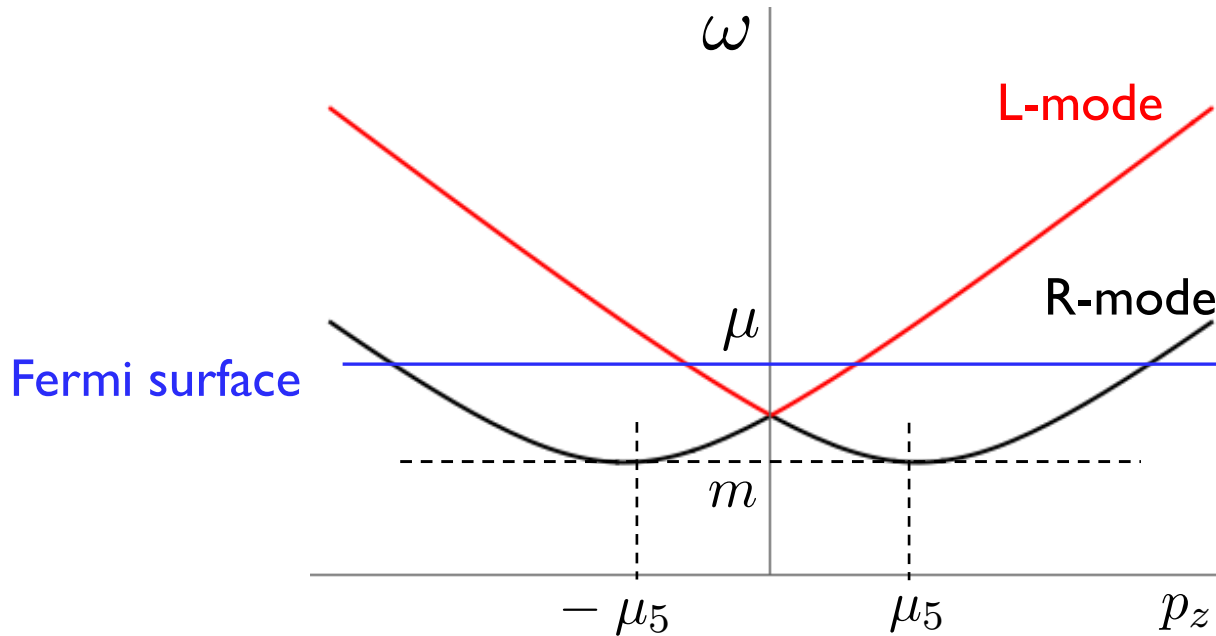
$$\omega_L = \sqrt{(|p_z| + \mu_5)^2 + m^2}$$

Helicity eigenstates

Fermi surface



Qualitatively the same as the massless case



- The difference between the L-mode and the R-mode disappears for $\mu_5 = 0$.
- Without a magnetic field, the spin directions of particles will not be fixed.
- The occupation number difference depends on the chemical potential μ and the fermion mass m .

$$J(\mu_5, B, \mu, m, T) = \mu_5 B \times f(\mu, m, T) \\ + \text{possible higher order terms in } (\mu_5, B)$$